

Neutrino-Nucleus Interaction

Physics around 1-10 GeV

outline

1. Neutrino interaction physics - introduction
2. Charged-Current Quasi-Elastic (CCQE) interaction
3. Neutrino baryonic resonance interaction
4. Neutrino shallow- and deep-inelastic scatterings
5. Conclusion

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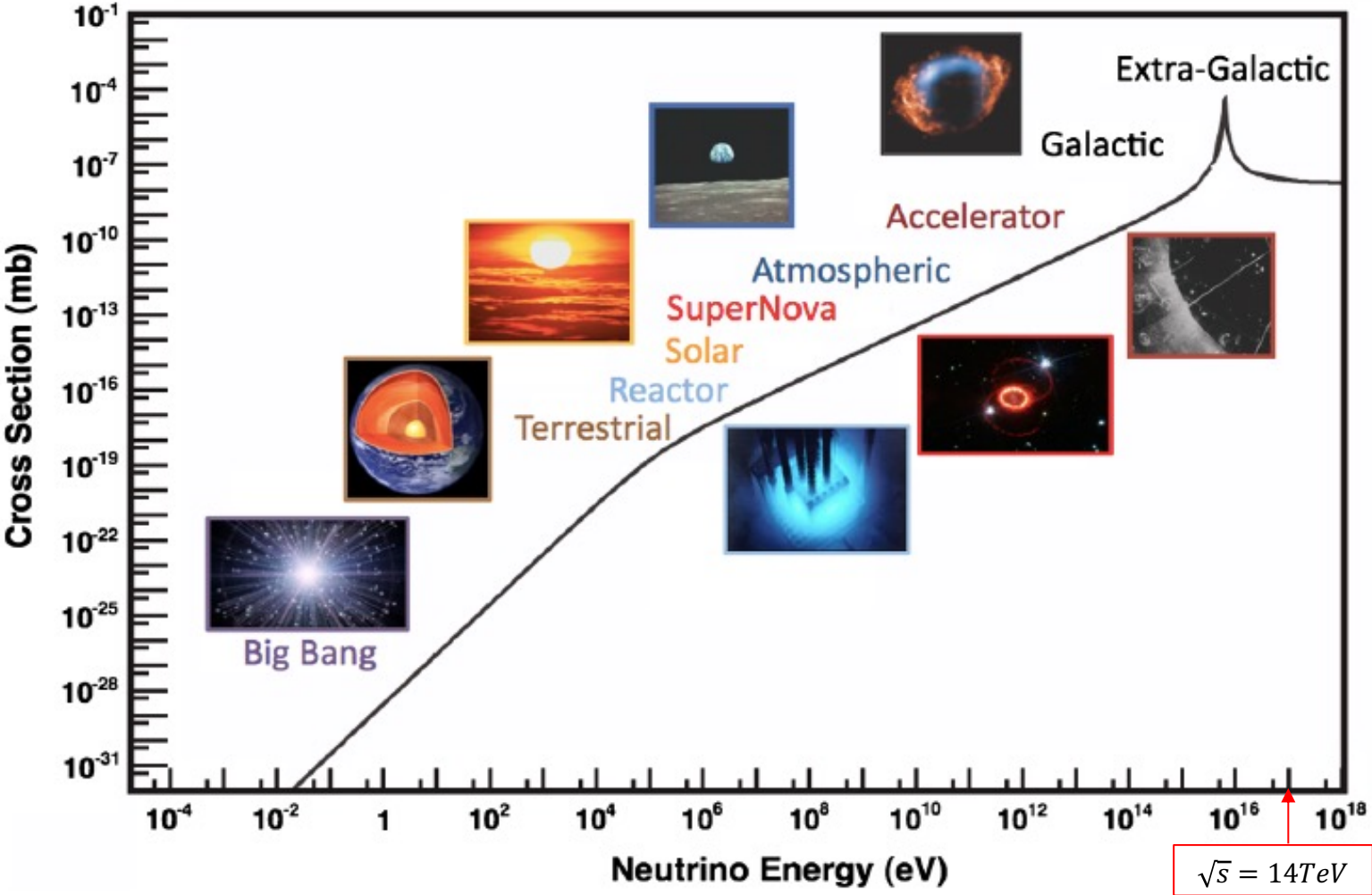
Neutrino Group seminar, Ruđer Bošković Institute, Zagreb, April 10, 2025

teppei.katori@cern.ch

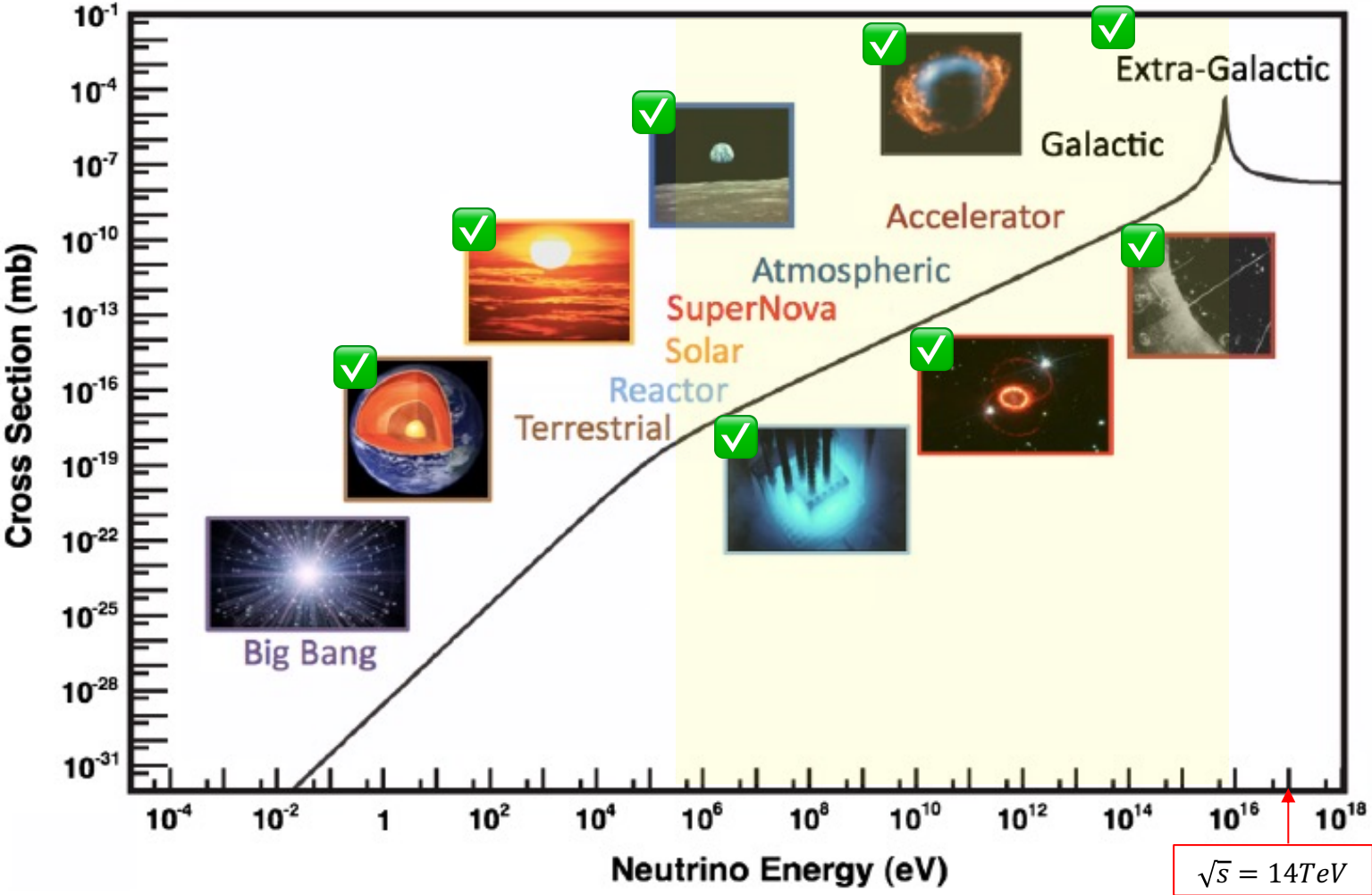
2025/04/10

- 1. Neutrino interaction physics - introduction**
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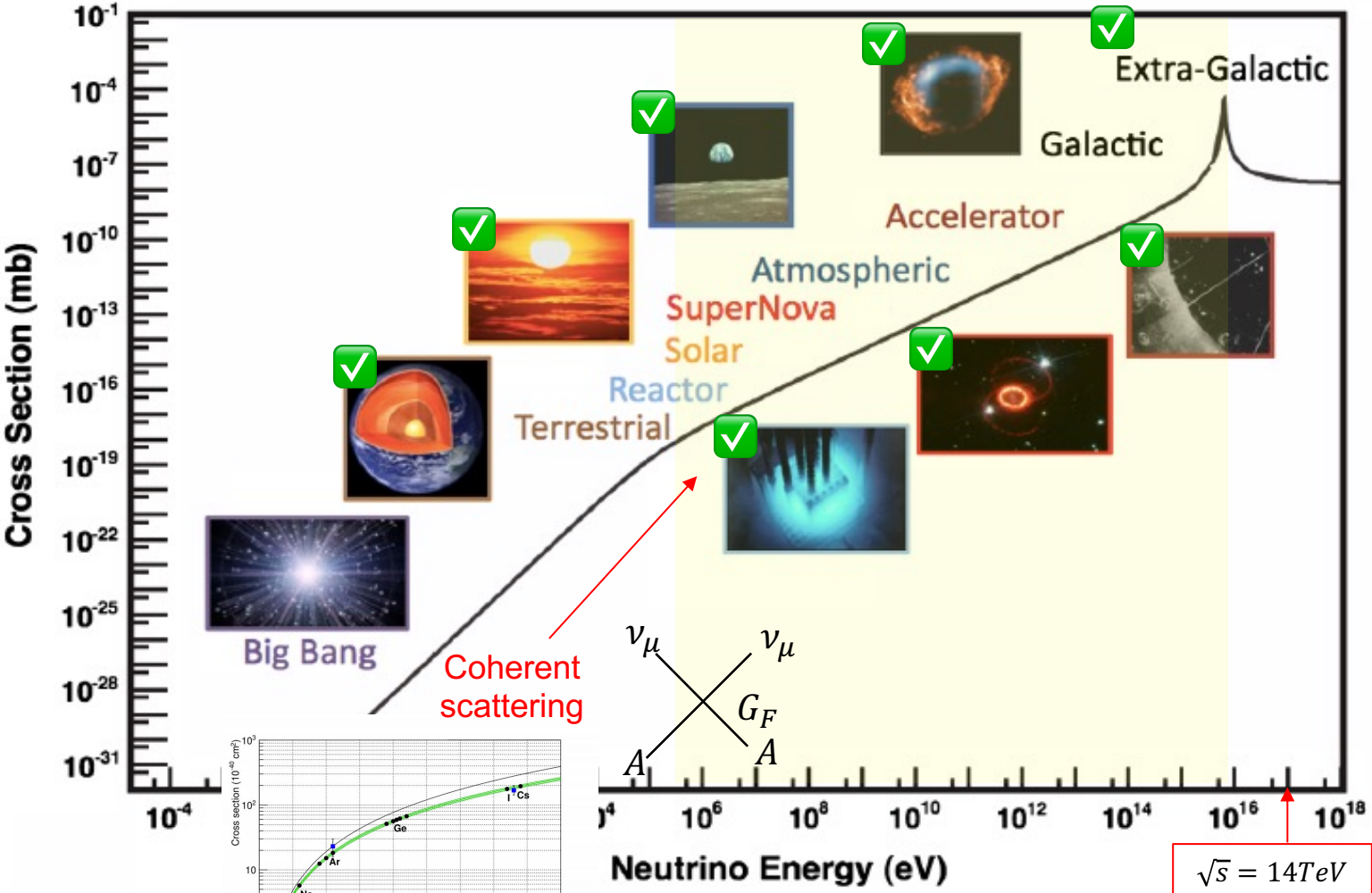
1. From eV to EeV: Neutrino cross sections across energy scales



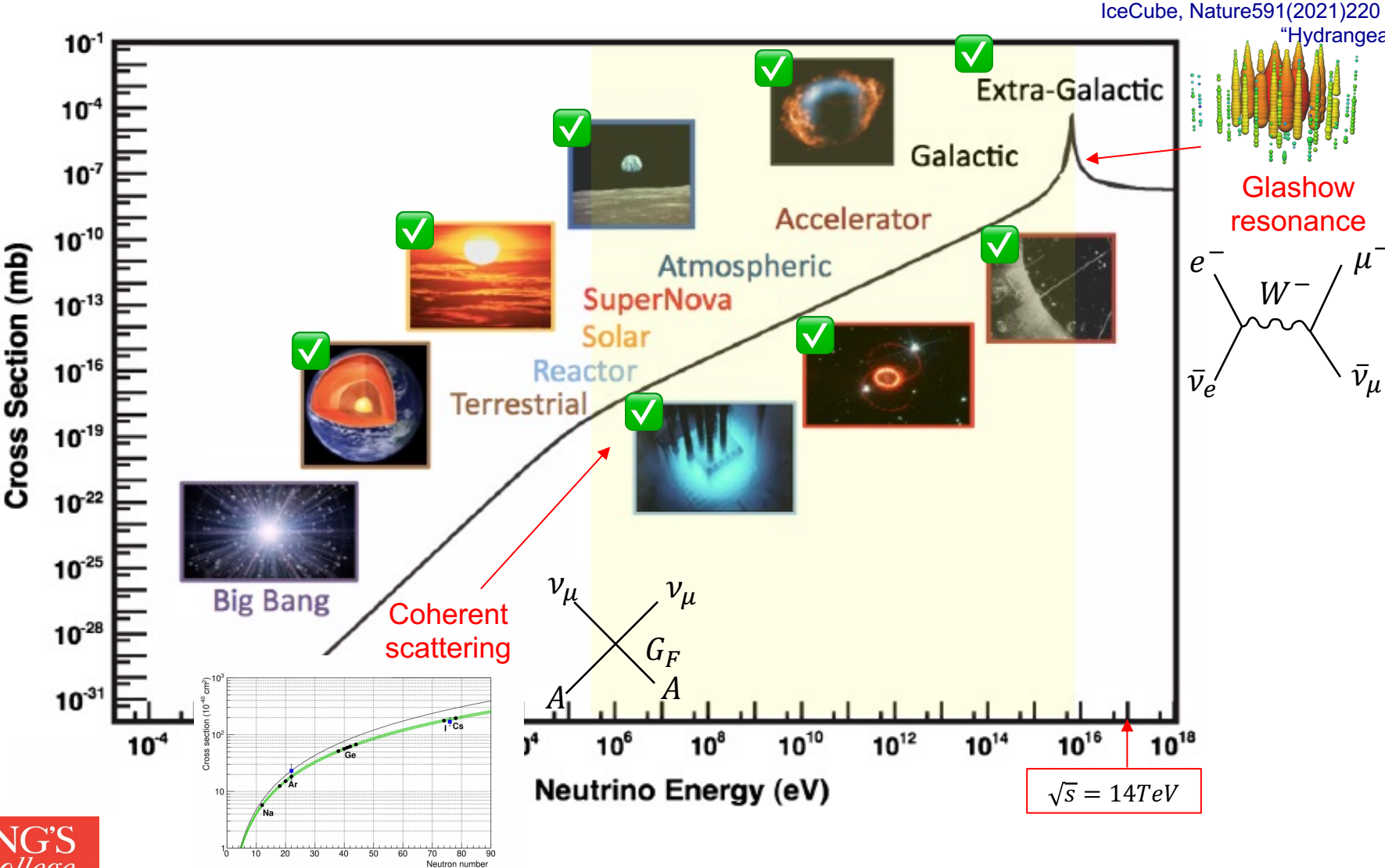
1. From eV to EeV: Neutrino cross sections across energy scales



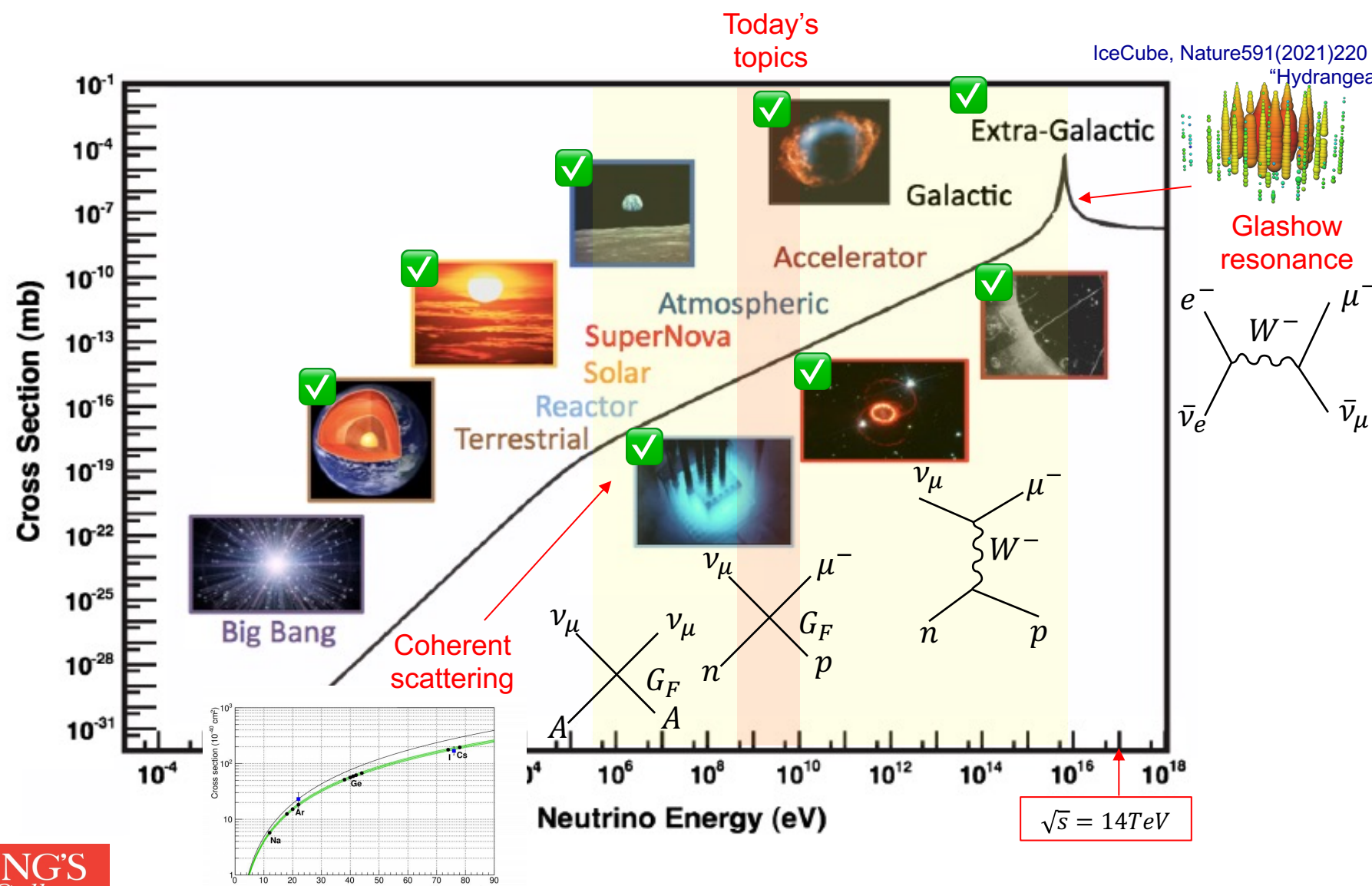
1. From eV to EeV: Neutrino cross sections across energy scales



1. From eV to EeV: Neutrino cross sections across energy scales



1. From eV to EeV: Neutrino cross sections across energy scales



1. PDG: Neutrino Cross Section Measurements

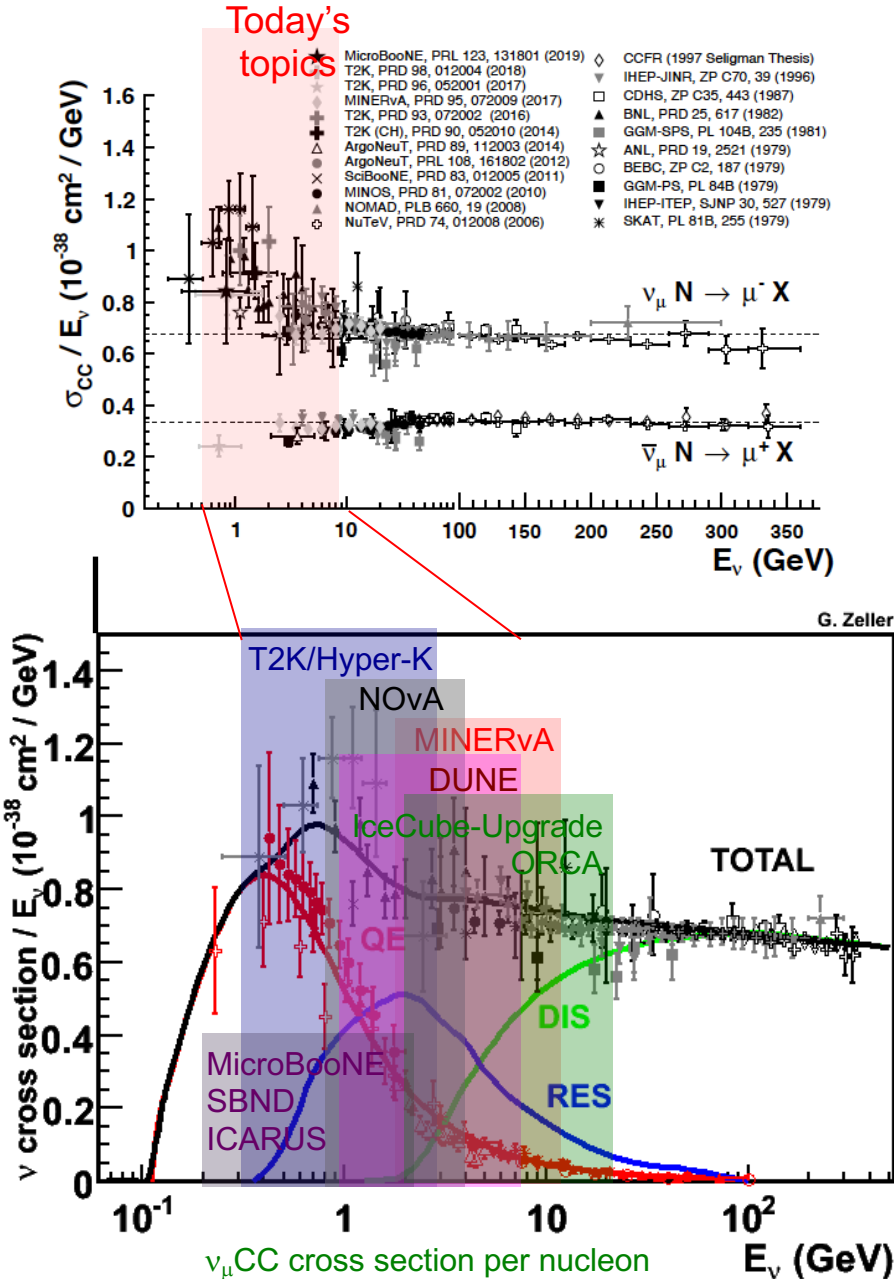
PDG has a summary of neutrino cross-section data since 2012!

Focus of this talk is around a few GeV

Table 52.2: Published measurements of neutrino and antineutrino CC inclusive cross sections from modern accelerator-based neutrino experiments.

experiment	measurement	target
ArgoNeuT	ν_μ [6, 7], $\bar{\nu}_\mu$ [7]	Ar
MicroBooNE	ν_μ [8, 26], ν_e [22]	Ar
MINERvA	ν_μ [9–11, 16, 17, 27], $\bar{\nu}_\mu$ [27], $\bar{\nu}_\mu/\nu_\mu$ [28]	CH, C/CH, Fe/CH, Pb/CH
MINOS	ν_μ [29], $\bar{\nu}_\mu$ [29]	Fe
NINJA	ν_μ [12], $\bar{\nu}_\mu$ [12]	H ₂ O
NOMAD	ν_μ [30]	C
SciBooNE	ν_μ [31]	CH
T2K	ν_μ [13, 14, 32–34], ν_e [23–25], $\bar{\nu}_\mu/\nu_\mu$ [15]	CH, H ₂ O, Fe

Today's
topics

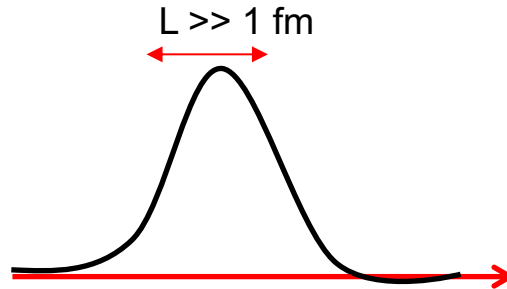


1. Neutrino interaction physics around 1-10 GeV

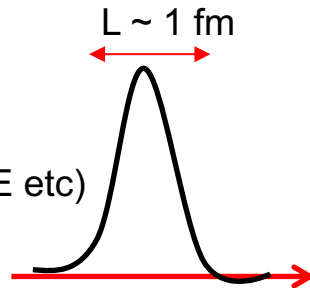
Size of wave packet $\sim$ momentum transfer ($\sim$ energy)

$$\hbar c = 197 \text{ MeV} \cdot \text{fm} \rightarrow 200 \text{ MeV} \sim 1 \text{ fm (size of nucleon)}$$

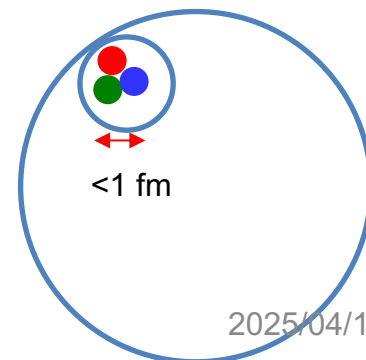
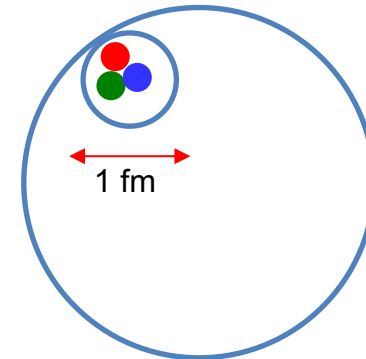
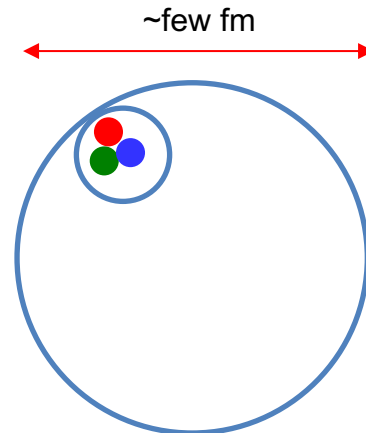
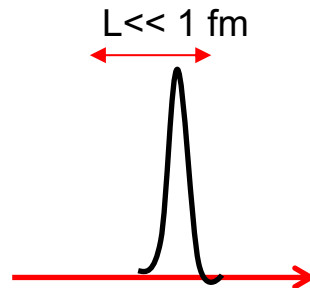
$\ll 1$ GeV neutrino beam
(solar neutrinos, etc)



~ 1 GeV neutrino beam
(T2K, NOvA, HyperK, DUNE etc)



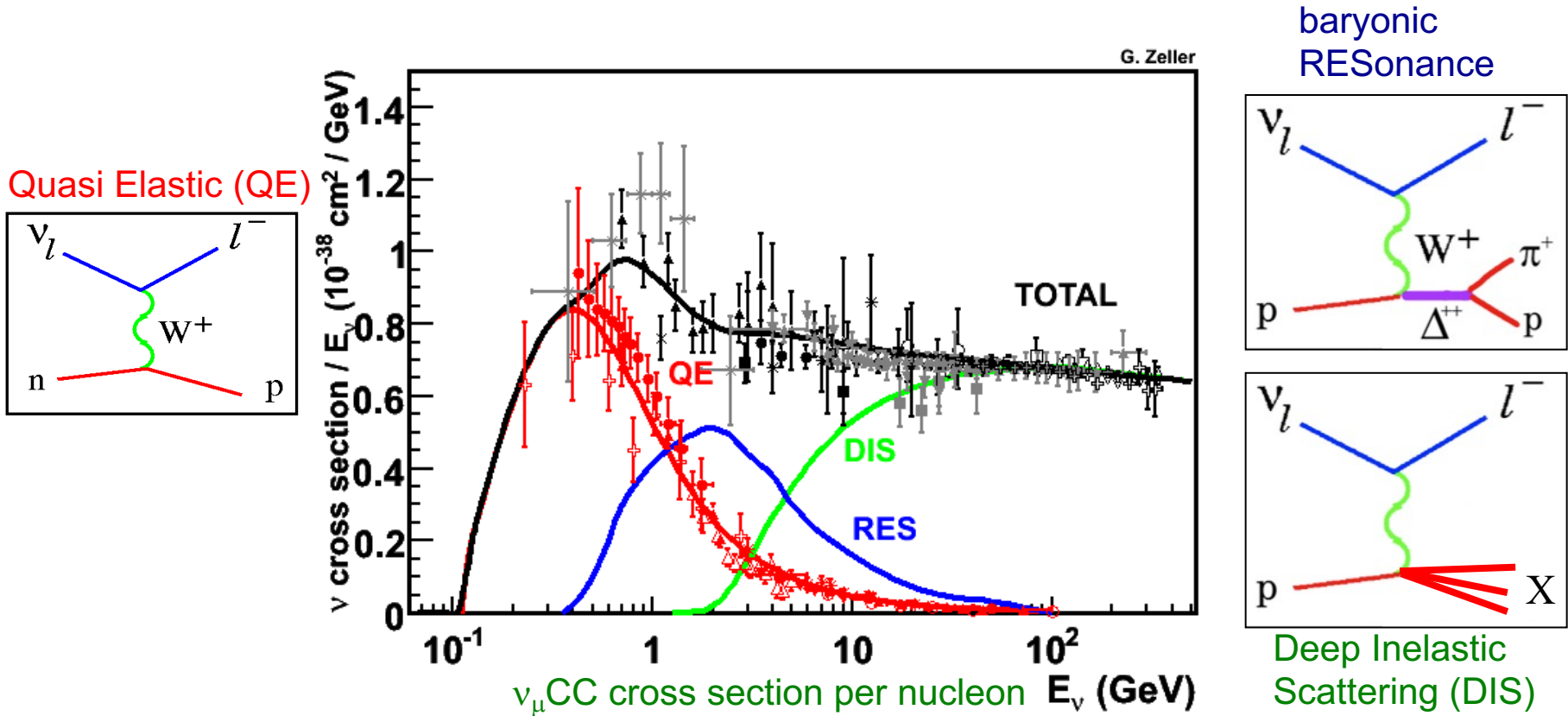
$\gg 1$ GeV neutrino beam
(LHC, astrophysical)



1. Neutrino interaction physics around 1-10 GeV

Neutrino interaction physics around 1-10 GeV

- degree of freedom change from nucleus $\rightarrow$ nucleon $\rightarrow$ parton
- There is no cut off (they all interfere)



1. Next goal of high energy physics

Establish Neutrino Standard Model (nSM)

- SM + 3 active massive neutrinos

Unknown parameters of nSM

1. Dirac CP phase
 2. θ_{23} ($\theta_{23}=40^\circ$ and 50° are same for $\sin^2\theta_{23}$, but not for $\sin\theta_{23}$)
 3. normal mass ordering $m_1 < m_2 < m_3$ or inverted mass ordering $m_3 < m_1 < m_2$
 4. Dirac or Majorana
 5. Majorana phases (x2)
 6. Absolute neutrino mass
- } not relevant to neutrino oscillation experiment

We need higher precision neutrino experiments around 1-10 GeV.

Low energy beam (~1 GeV)

- shorter baseline (lower flux reduction)
- lower neutrino production
- lower interaction rate
- kinematic energy reconstruction

High energy beam (~few GeV)

- longer baseline (higher flux reduction)
- higher neutrino production
- higher interaction rate
- calorimetric energy reconstruction

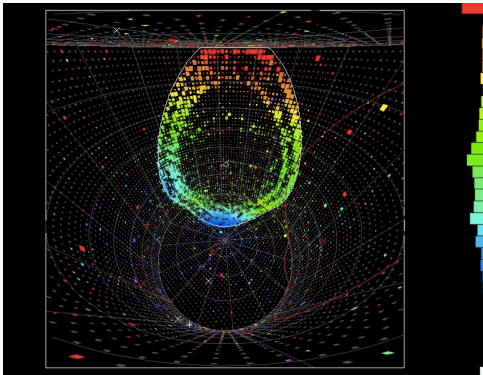
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$$P_{\mu \rightarrow e}(L / E) = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 (eV^2) \frac{L(km)}{E(GeV)} \right)$$

1. Next goal of high energy physics

Kinematics energy reconstruction

- problem: it assume 2-body neutrino interaction with single nucleon



$$E_{\nu}^{QE} = \frac{ME_{\nu} - 0.5m_{\mu}^2}{M - E_{\mu} + p_{\mu}\cos\theta}$$

Low energy beam (~1 GeV)

- shorter baseline (lower flux reduction)
- lower neutrino production
- lower interaction rate
- kinematic energy reconstruction

Calorimetric energy reconstruction

- problem: you need to measure energy deposit from all outgoing particles



$$E_{\nu}^{Cal} = E_{\mu} + \sum_{i=1}^{all} E_{had}^i$$

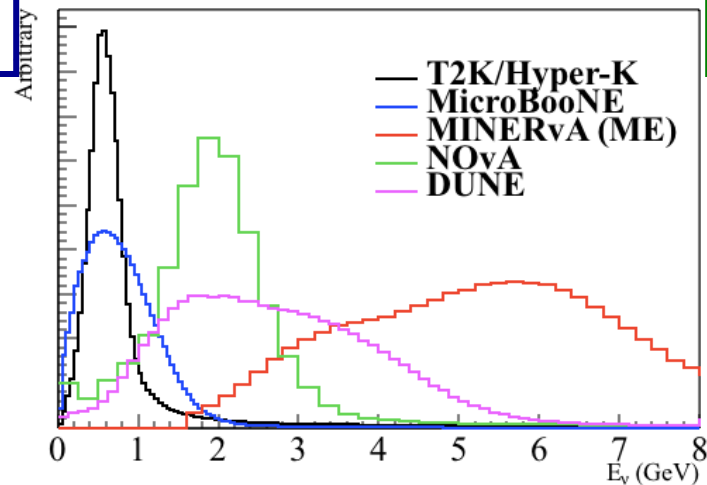
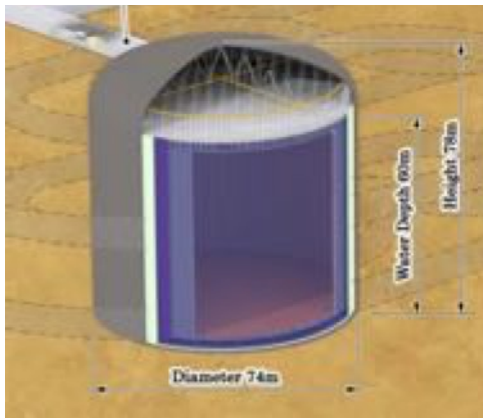
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1. Next goal of high energy physics

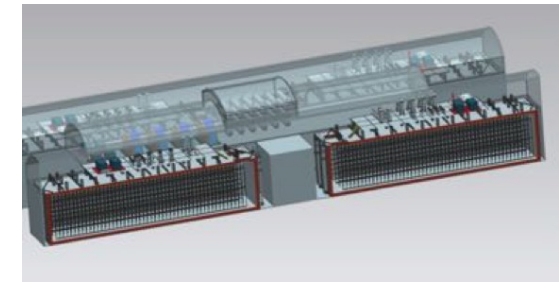
Hyper-Kamiokande (Japan)

- Water target
- 300 km baseline
- Narrow band 0.6 GeV
- Low spatial resolution
- High time resolution



DUNE (USA)

- Argon target
- 1300 km baseline
- wide band 1-4 GeV
- High spatial resolution
- Low time resolution



Low energy beam (~1 GeV)

- shorter baseline (lower flux reduction)
- lower neutrino production
- lower interaction rate
- kinematic energy reconstruction

High energy beam (~few GeV)

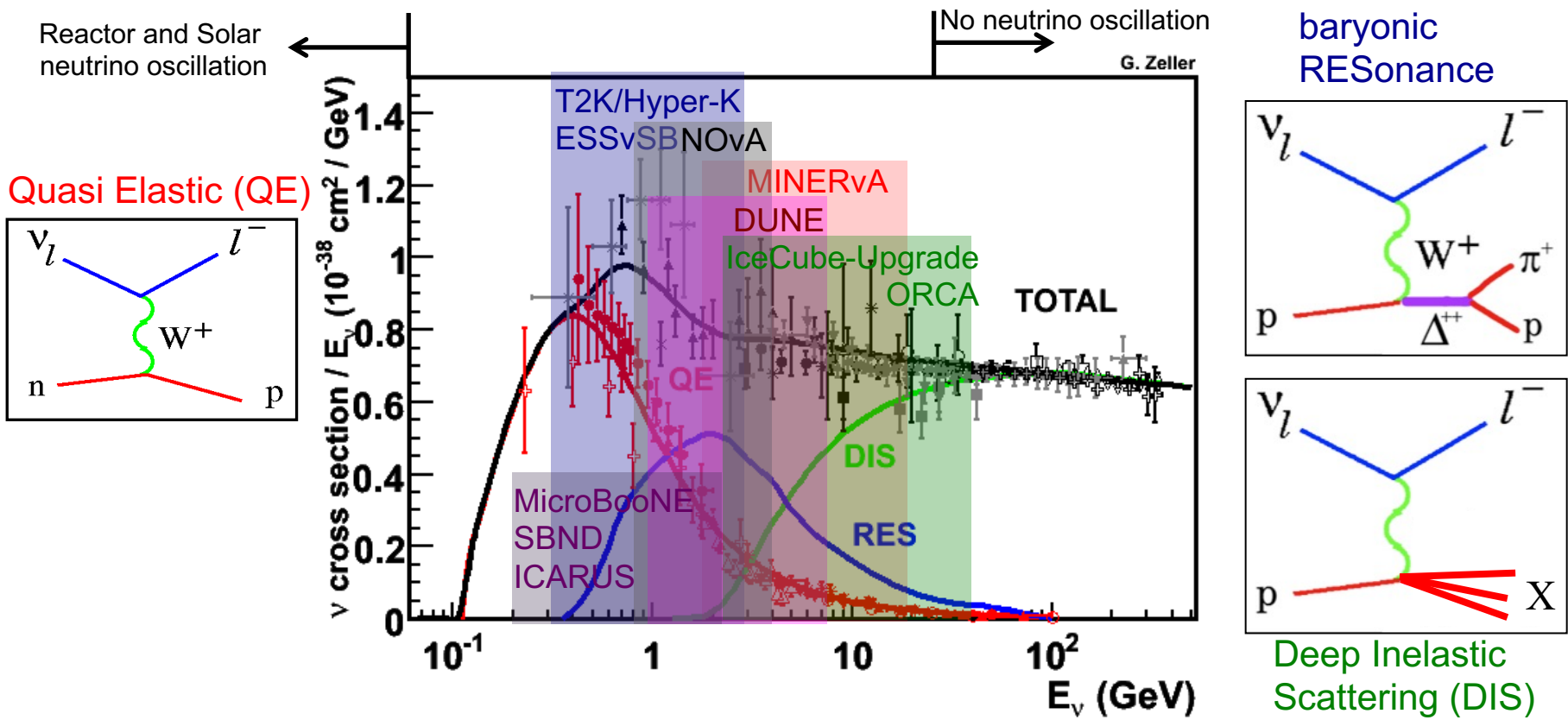
- longer baseline (higher flux reduction)
- higher neutrino production
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$$P_{\mu \rightarrow e}(L/E) = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 (eV^2) \frac{L(km)}{E(GeV)} \right)$$

1. Next generation neutrino oscillation experiments

Current and future neutrino oscillation experiments

- J-PARC: T2K, Hyper-Kamiokande, ESSvSB
- Fermilab: MicroBooNE/SBND/ICARUS, MINERvA, NOvA, DUNE
- Atmospheric: Hyper-Kamiokande, ORCA, IceCube-Upgrade



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$$P_{\mu \rightarrow e}(L / E) = \sin^2 2 \theta \sin^2 \left(1.27 \Delta m^2 (eV^2) \frac{L(km)}{E(GeV)} \right)$$

1. Neutrino cross-section formula

Cross-section

- product of Leptonic and Hadronic tensor

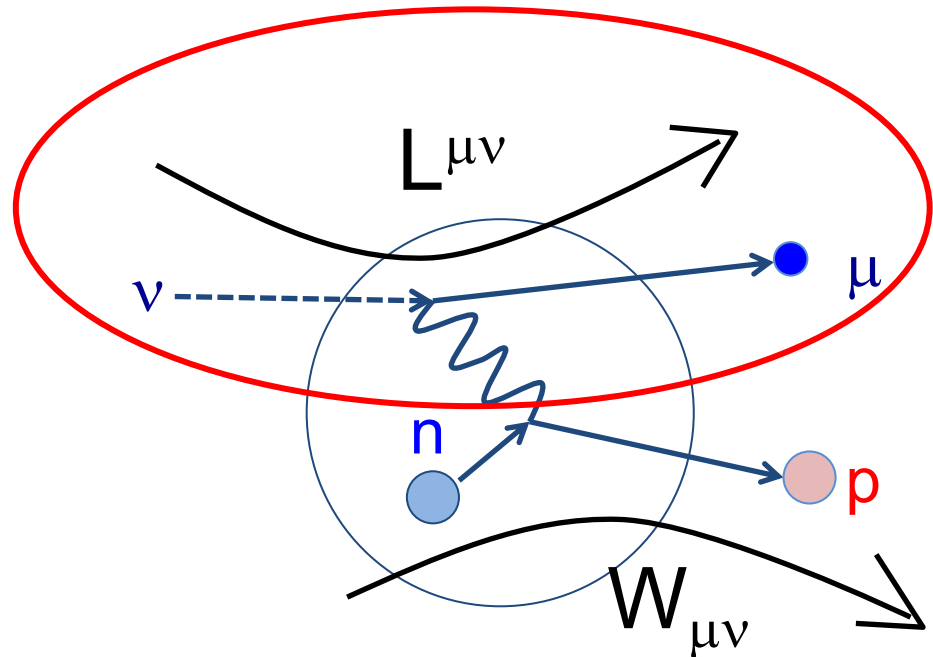
$$d\sigma \sim L^{\mu\nu} W_{\mu\nu}$$

Leptonic tensor

→ the Standard Model (easy)

Hadronic tensor

→ nuclear physics (hard)



1. Neutrino cross-section formula

Cross-section

- product of Leptonic and Hadronic tensor

$$d\sigma \sim L^{\mu\nu} W_{\mu\nu}$$

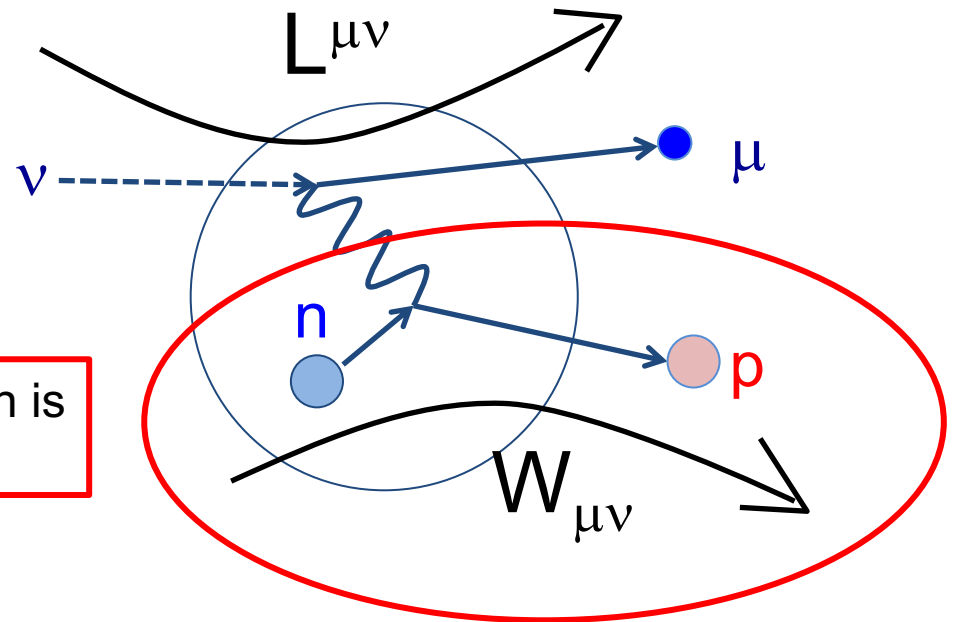
Leptonic tensor

→ the Standard Model (easy)

Hadronic tensor

→ nuclear physics (hard)

All complication of neutrino cross-section is how to model the hadronic tensor part



1. Neutrino event generator

GENIE

<https://github.com/GENIE-MC>

- Used by Fermilab experiments

NEUT

(no public website)

- Used by Japanese neutrino experiments

NuWro

<https://nuwro.github.io/user-guide/>

- Independent generator

GIBUU

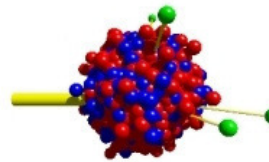
<https://gibuu.hepforge.org/trac/wiki>

- BUU transport to simulate hadron final states

NUISANCE

<https://nuisance.hepforge.org/>

- Data-Neutrino generator comparison framework



GIBUU

The Giessen Boltzmann-Uehling-Uhlenbeck Project



Achilles (New!)

<https://arxiv.org/abs/2205.06378>

- Theory-driven better factorization

1. Neutrino event generator

Fast simulation

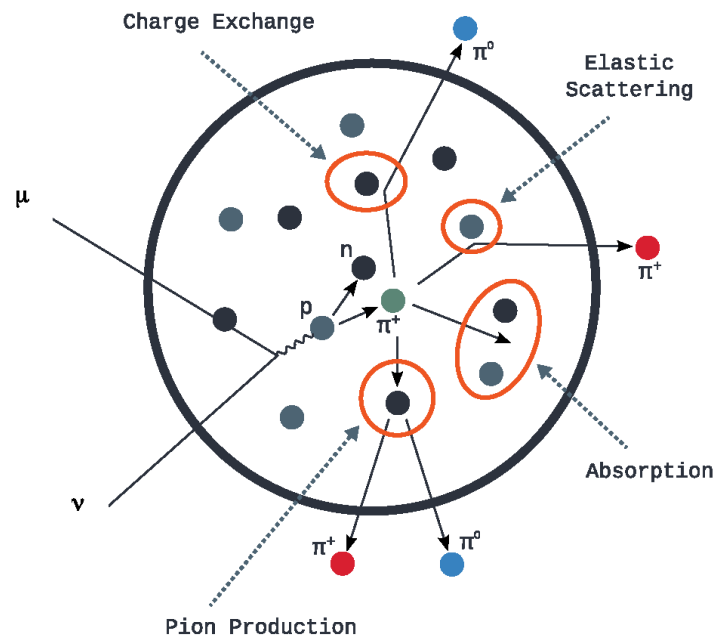
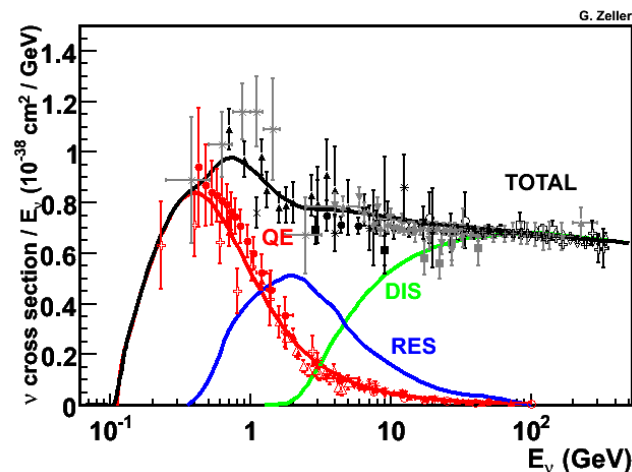
- Monte Carlo method

Merge models to cover all kinematic phase space

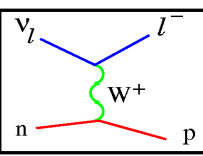
- Inverse beta decay (IBD)
- Charged-current quasi-elastic (CCQE)
- Resonance baryon production (RES)
- Deep-inelastic scattering (DIS)
- etc

Nuclear effects

- Pauli blocking
- Fermi motion
- Final state interactions (scattering, absorption, etc)
- Nucleon correlations
- etc



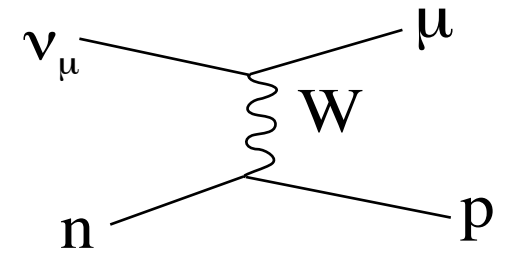
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2. Charged Current Quasi-Elastic scattering (CCQE)

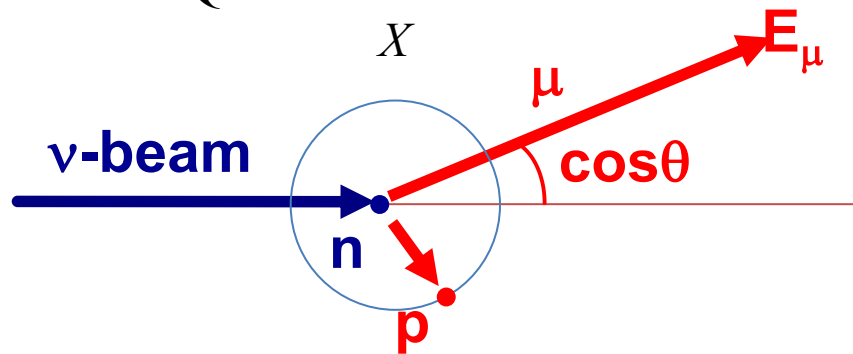
The simplest and the most abundant interaction around ~ 1 GeV.

$$\nu_\mu + n \rightarrow p + \mu^- \quad (\nu_\mu + X \rightarrow X' + \mu^-)$$



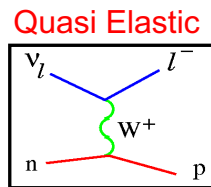
Neutrino energy is reconstructed from the observed lepton kinematics
“QE assumption”

1. assuming neutron at rest
2. assuming interaction is CCQE



$$E_\nu^{QE} = \frac{ME_\nu - 0.5m_\mu^2}{M - E_\mu + p_\mu \cos\theta}$$

CCQE is the single most important channel of neutrino oscillation physics
T2K, NOvA, MicroBoonE, Hyper-Kamiokande,...etc



2. CCQE puzzle

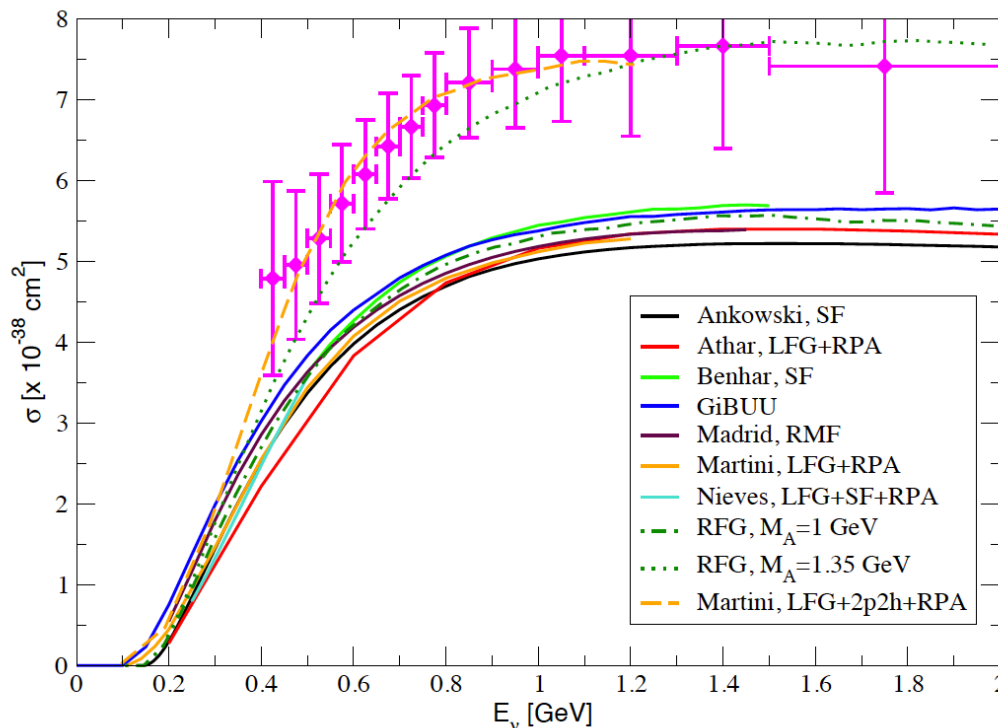
Simplest channel, but both shape and normalization disagree

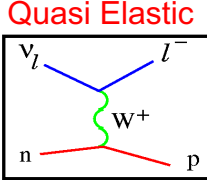
1. low Q^2 suppression $\rightarrow$ Low forward efficiency? (detector?)
2. high Q^2 enhancement $\rightarrow$ Axial mass > 1.0 GeV? (physics?)
3. large normalization $\rightarrow$ Beam simulation is wrong? (flux?)

CCQE interaction on nuclear targets are precisely measured by electron scattering

- Lepton universality = precise prediction for neutrino CCQE cross-section...?

MiniBooNE ν_μ CCQE cross section on ^{12}C target (per nucleon)





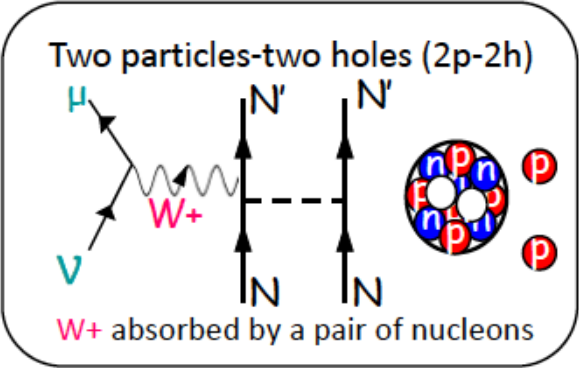
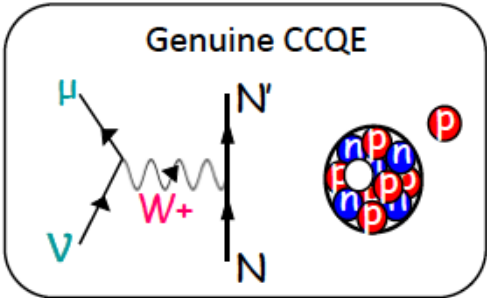
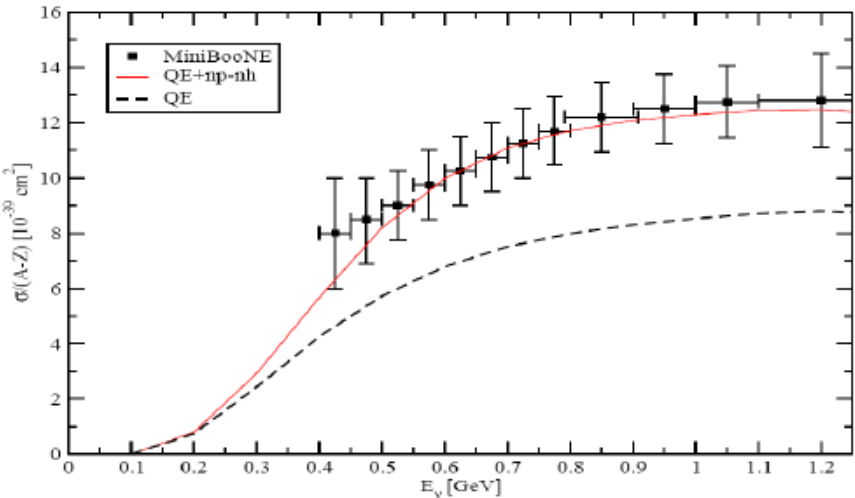
2. Solution of CCQE puzzle

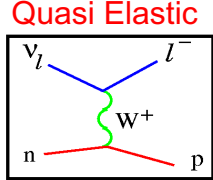
Presence of 2-body current

- CCQE is identified from single outgoing charged lepton events
- Significant fraction of events are not from 2-body neutrino-nucleon interactions
- Martini et al showed 2p-2h effect can add up ~30% more cross section

An explanation of this puzzle

Inclusion of the multinucleon emission channel (np-nh)





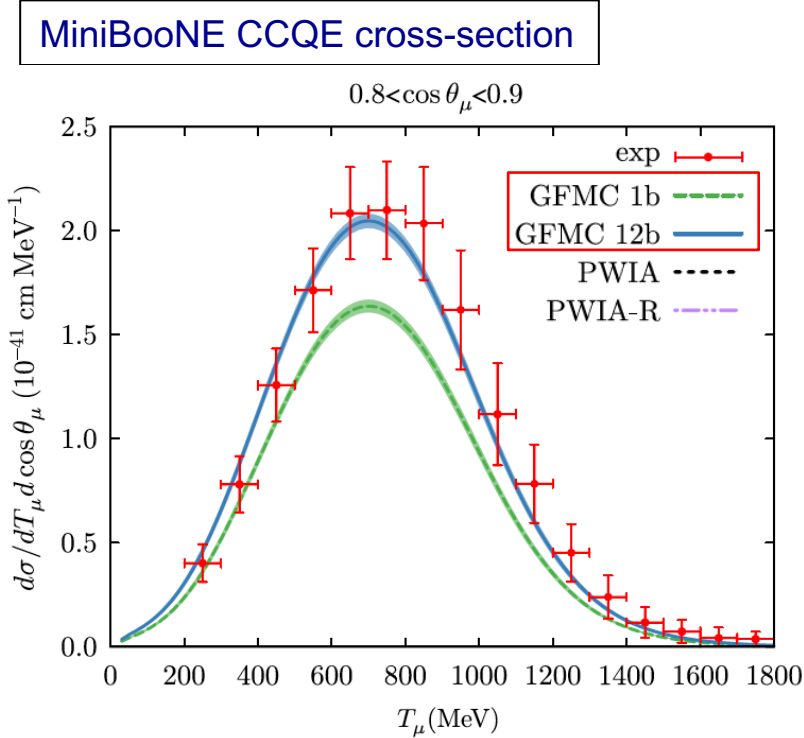
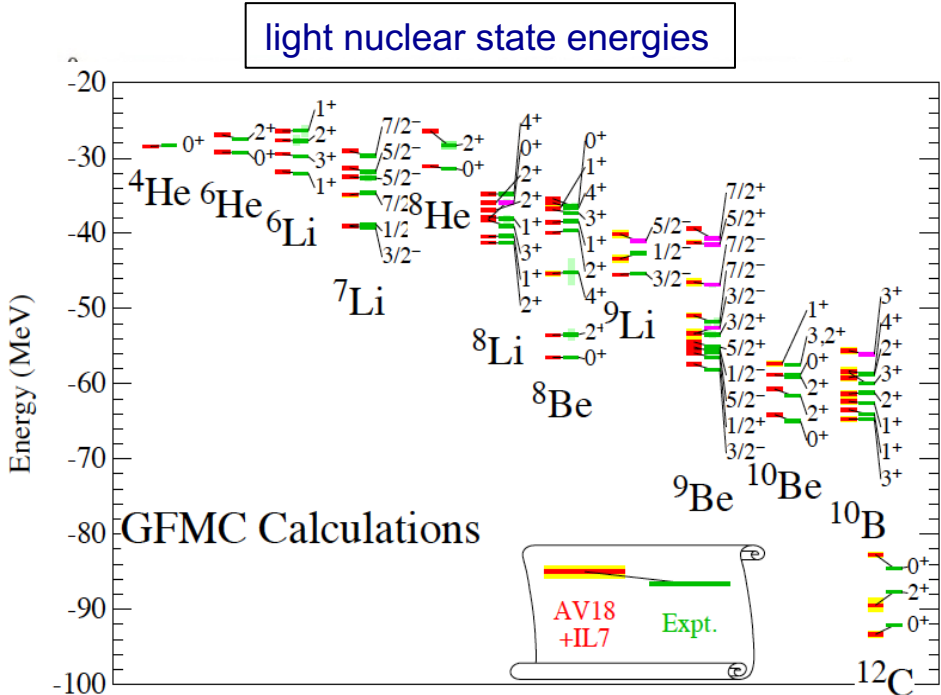
2. Nucleon correlations in neutrino physics

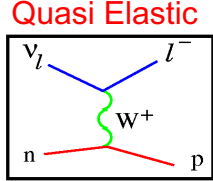
Ab-initio calculation

- Quantum Monte Carlo (QMC)
- Predicts energy levels of all light nuclei
- Consistent result with phenomenological models
- Ground state includes correct nucleon correlations

$$|\Psi_V\rangle = \mathcal{S} \prod_{i < j}^A \left[1 + \boxed{U_{ij}} + \sum_{k \neq i, j}^A \boxed{\tilde{U}_{ijk}^{TNI}} \right] |\Psi_J\rangle$$

2N potential (Av18) 3N potential (IL7)





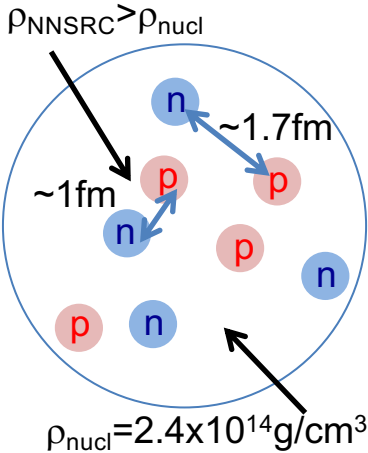
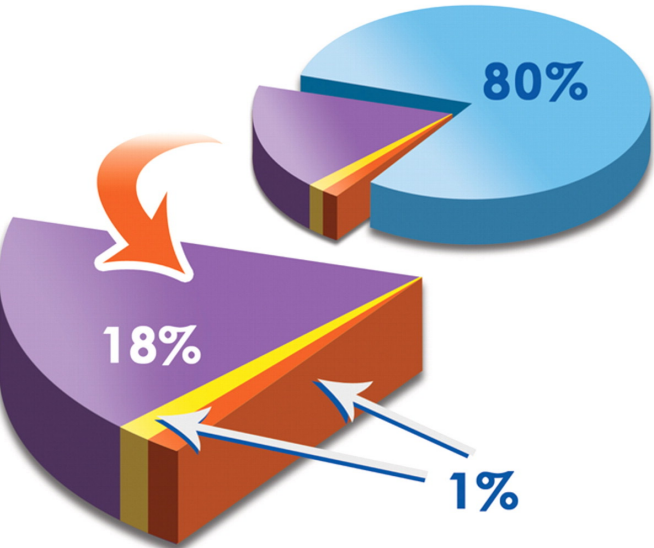
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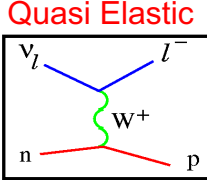
2N potential (Av18) 3N potential (IL7)



Physics of nucleon correlation

- neutrino interaction
- $0\nu\beta\beta$
- Direct WIMP detection
- EMC effect
- etc

Nucleon correlation is a very hot topics!

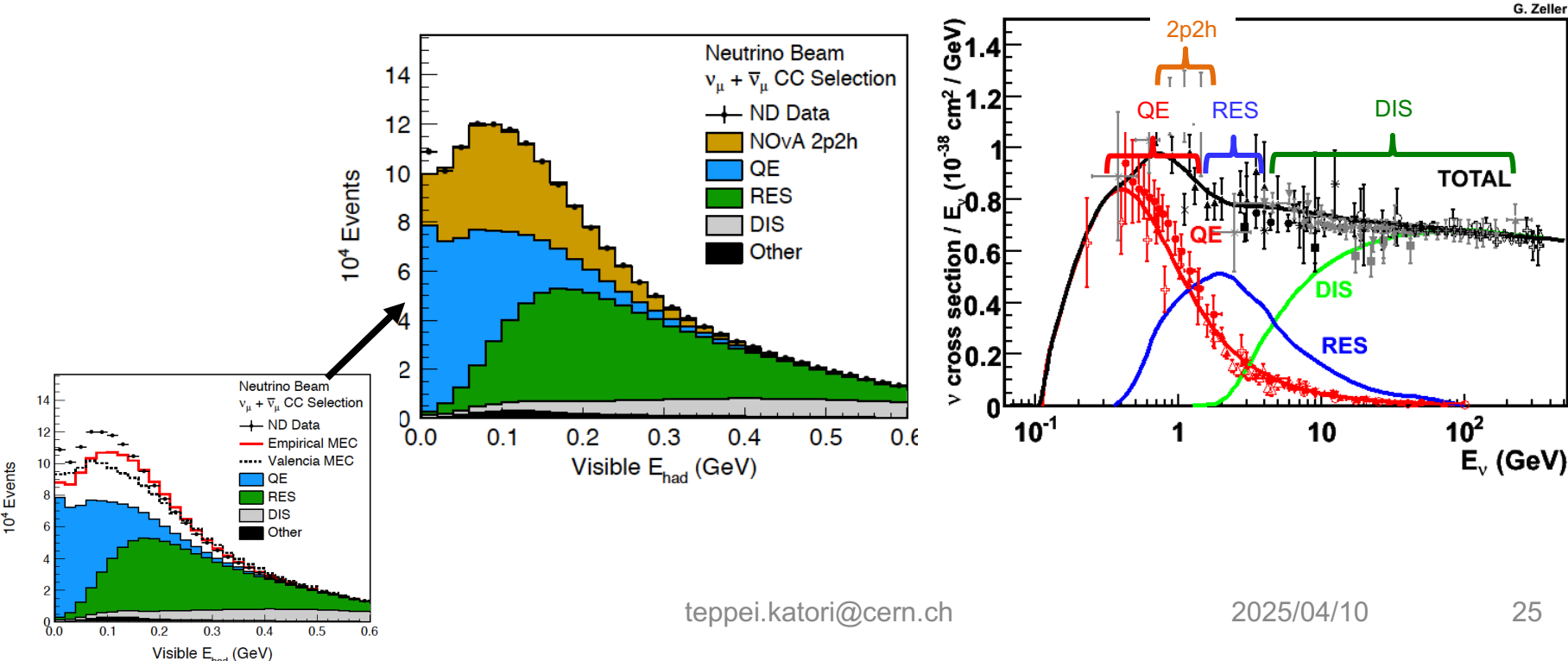


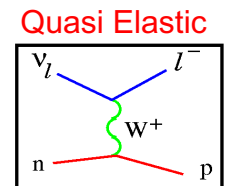
2. Nucleon correlations in neutrino physics

2-particle 2-hole (2p2h) effect

- Essential to describe data
- The biggest topic in nuxsec community (T2K, NOvA, MINERvA, MicroBooNE, etc)
- 2p2h models in generators don't describe data without heavy tuning
- High resolution detector (LArTPC, emulsion, etc) can find what is going on?

NOvA near detector data-MC comparison after fit



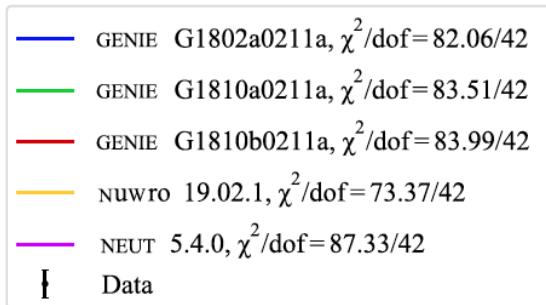
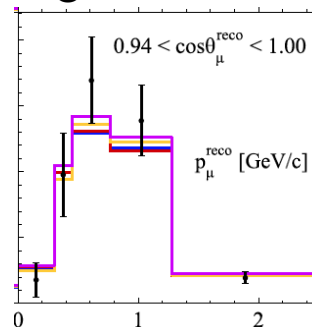


2. Data tension

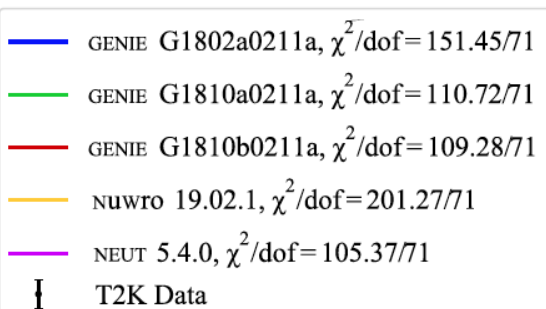
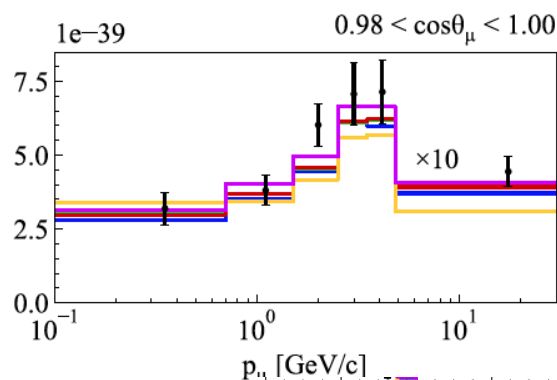
Data tension – external: T2K vs. MINERvA vs. MicroBooNE

- Different kinematic coverage, different target

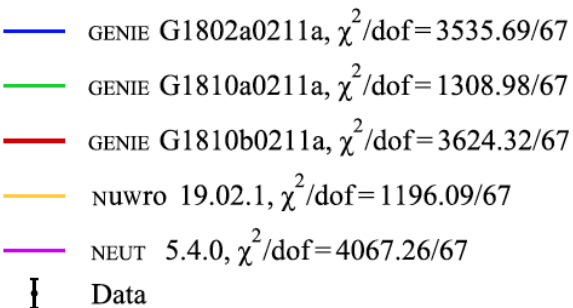
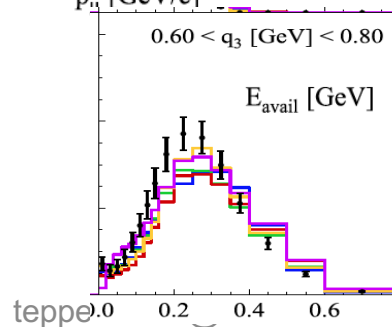
MicroBooNE CC inclusive double differential cross-section

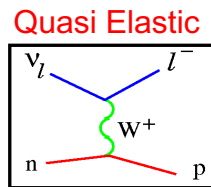


T2K CC inclusive double differential cross-section



MINERvA CC inclusive double differential cross-section

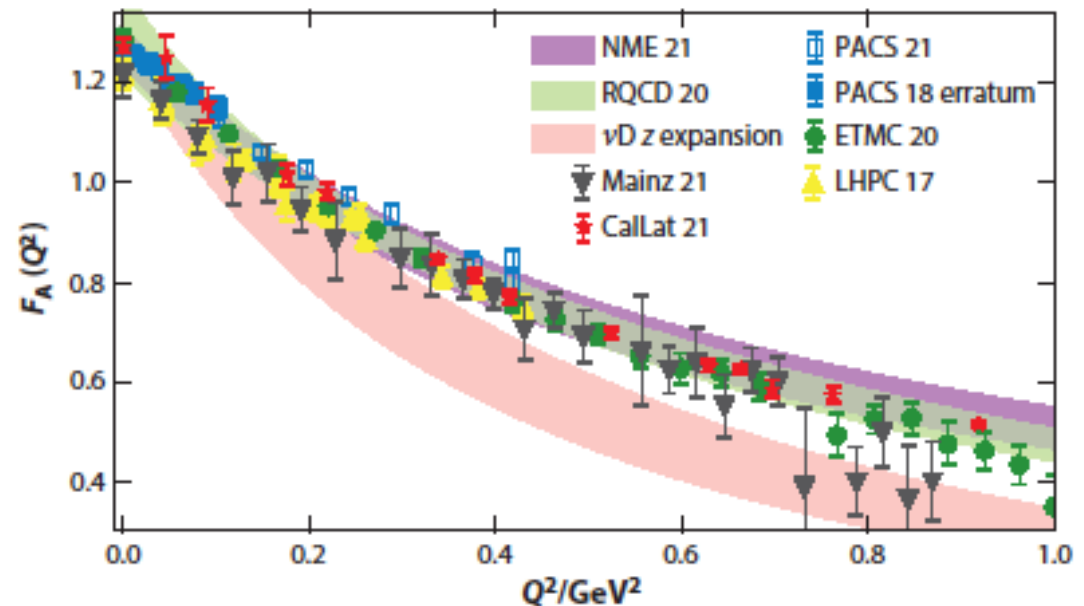




2. Large axial mass?

Nucleon axial mass (nucleon parameter)

- 2p2h effect was interpreted with large $M_A^{\text{QE}} \sim 1.3 \text{ GeV}$
- With 2p2h effect, we expect $M_A^{\text{QE}} \sim 1.0 \text{ GeV}$
- Latest lattice QCD calculations suggests different value, $M_A^{\text{QE}} \sim 1.3 \text{ GeV}$?
- Role of 2p2h is smaller than expected?
- $M_A^{\text{QE}} \sim 1.0 \text{ GeV}$ is supported by photo-pion production experiments
- $M_A^{\text{QE}} \sim 1.0 \text{ GeV}$ is supported by neutrino bubble chamber data



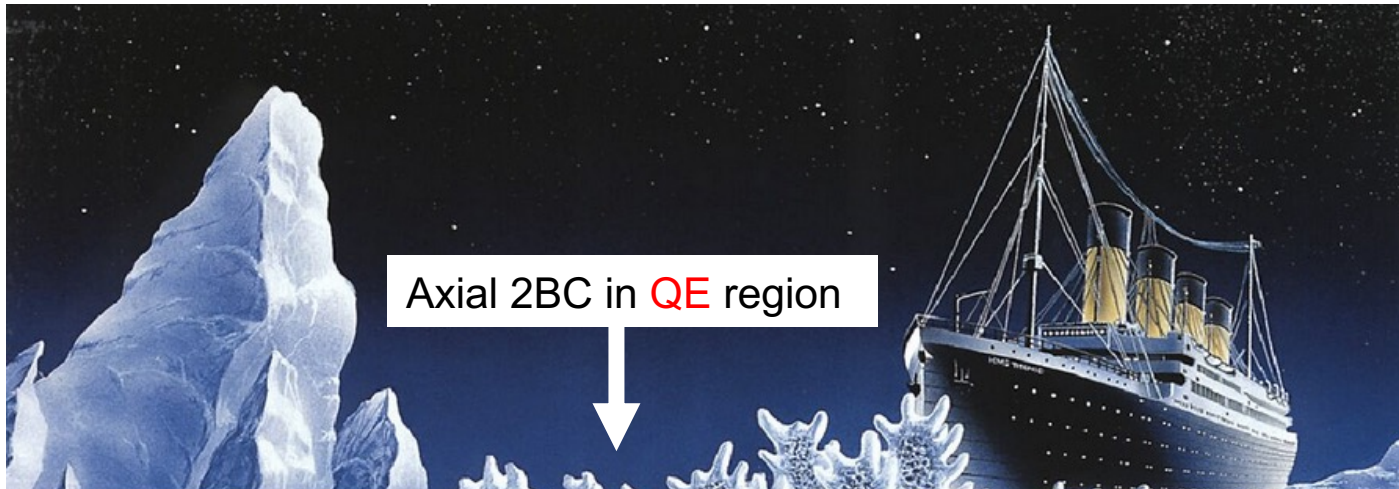
Neutrino data anomalies are (mostly) by Strong interaction



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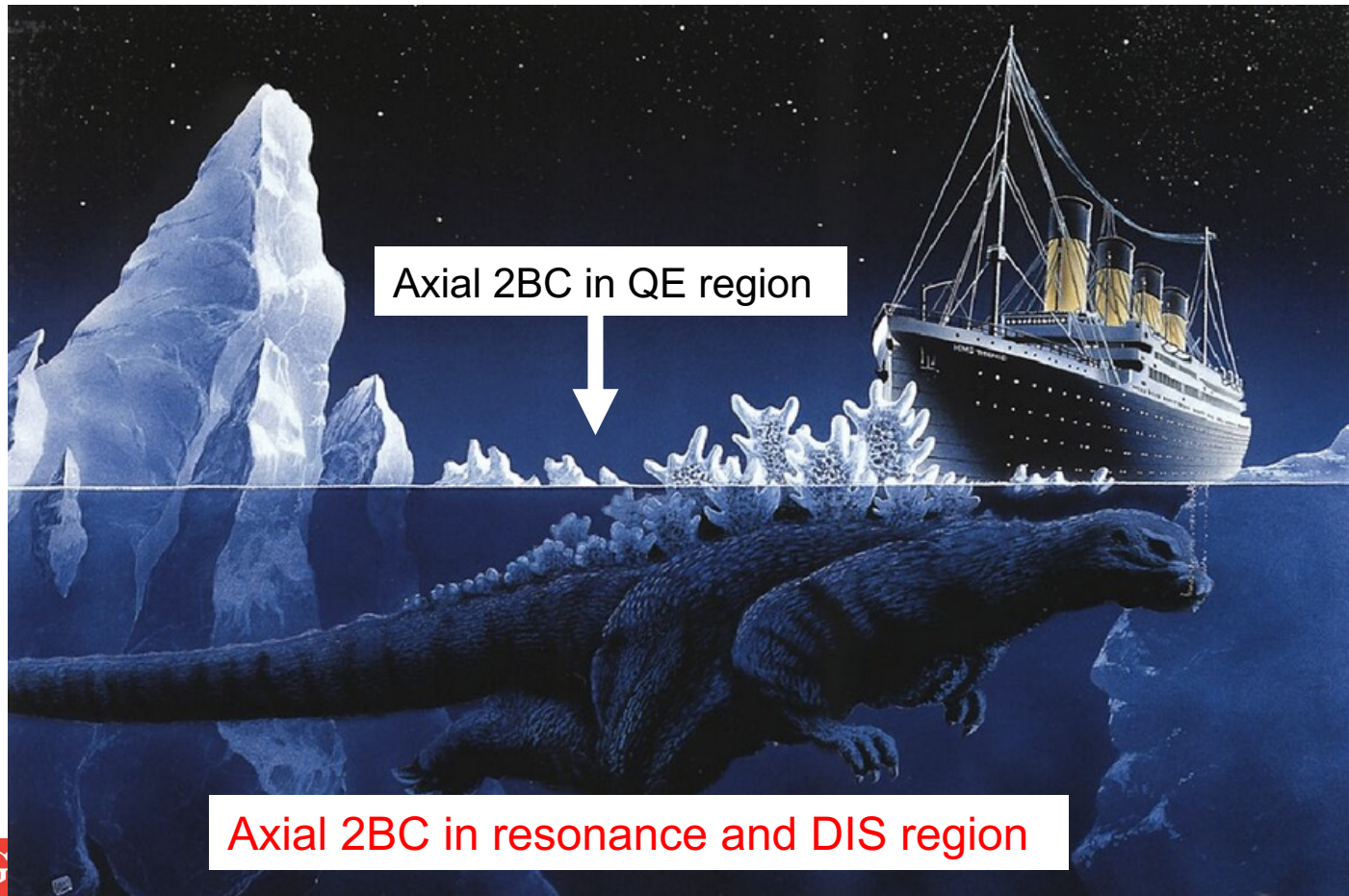
3. Beyond QE peak

Axial 2-body current in QE region may be a tip of the iceberg...



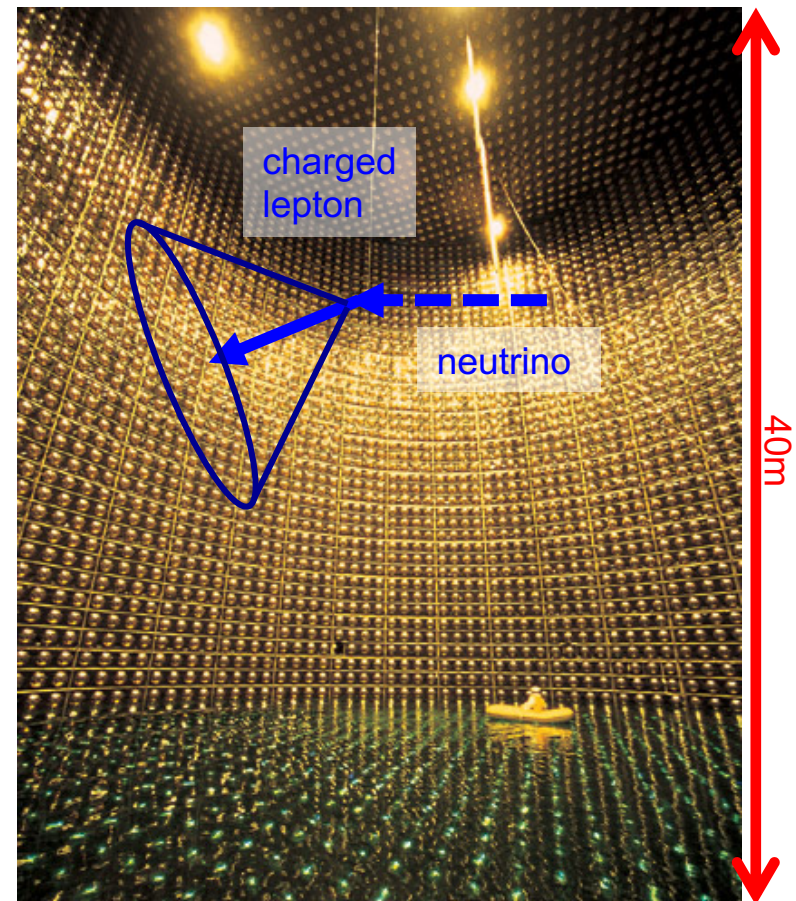
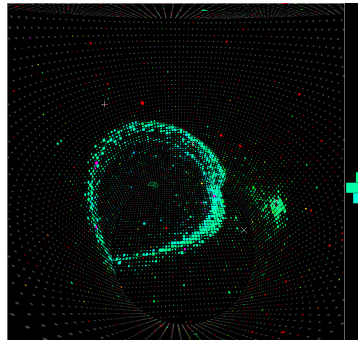
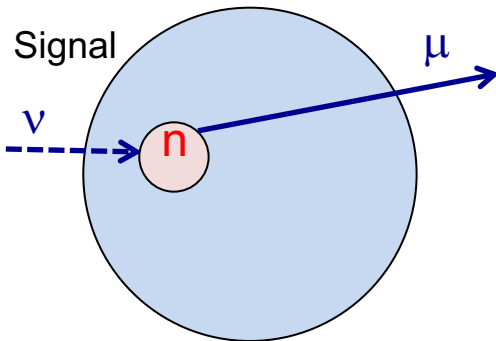
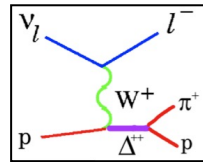
3. Beyond QE peak

Axial 2-body current in QE region may be a tip of the iceberg..., or maybe a tip of gozilla!



3. non-QE background (resonance pion production)

non-QE background $\rightarrow$ shift spectrum

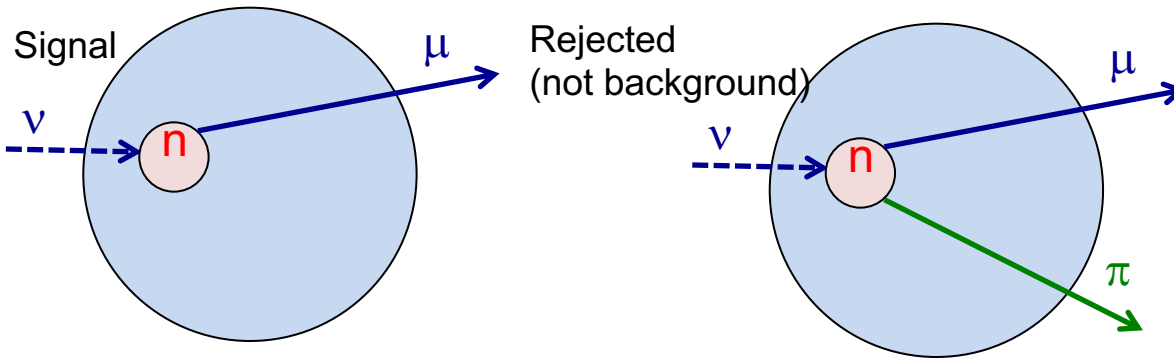
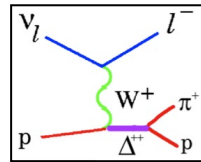


Water Cherenkov detector

- Big and dense, to maximize interaction rate
- Coarsely instrumented, to minimize cost (not great detector to measure hadrons)

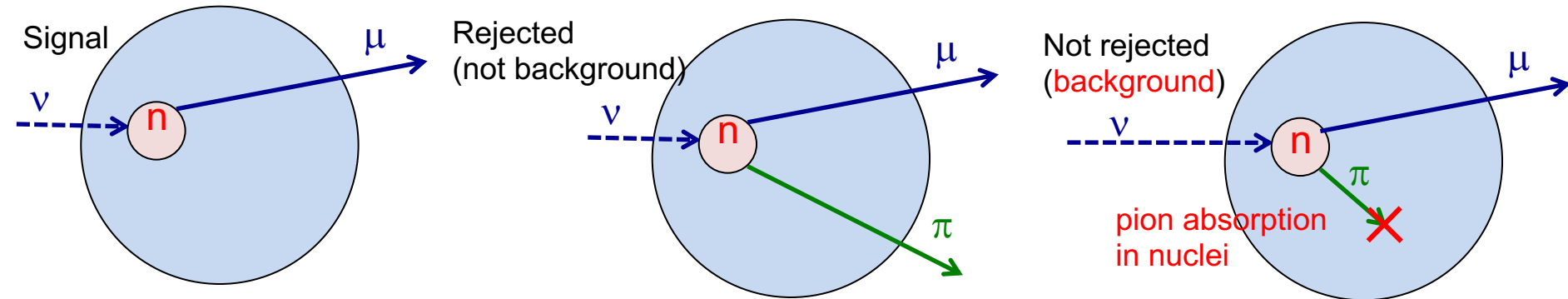
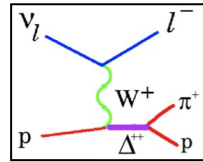
3. non-QE background (resonance pion production)

non-QE background $\rightarrow$ shift spectrum



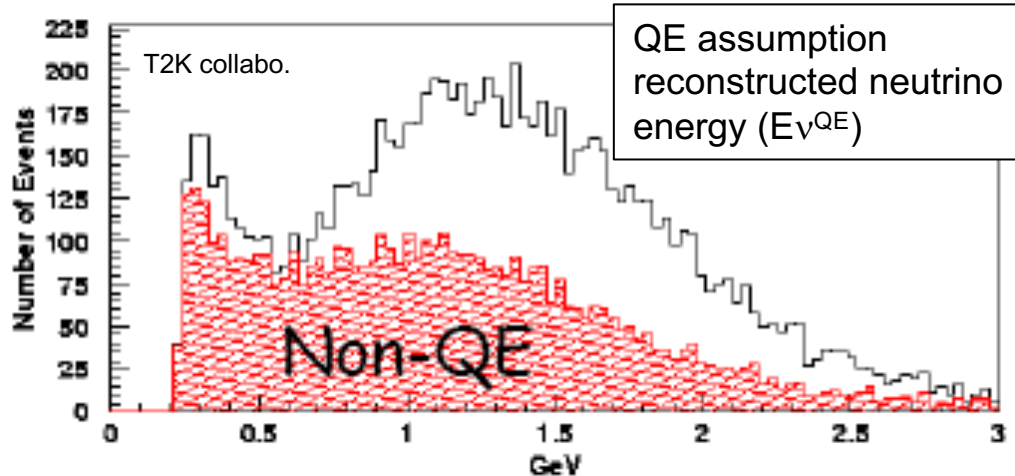
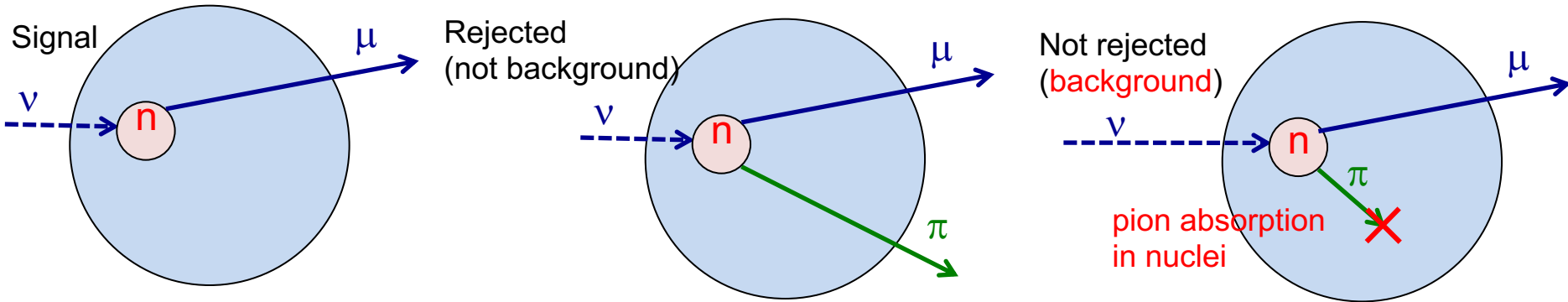
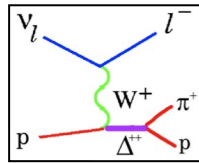
3. non-QE background (resonance pion production)

non-QE background $\rightarrow$ shift spectrum

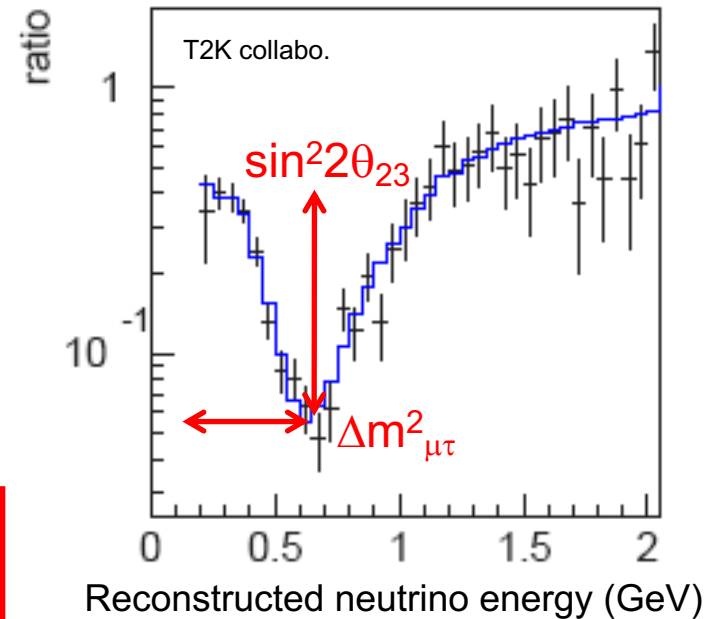


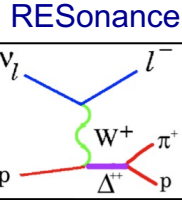
3. non-QE background (resonance pion production)

non-QE background $\rightarrow$ shift spectrum



Solution: you need a good prediction of
neutrino-induced pion production (hard)

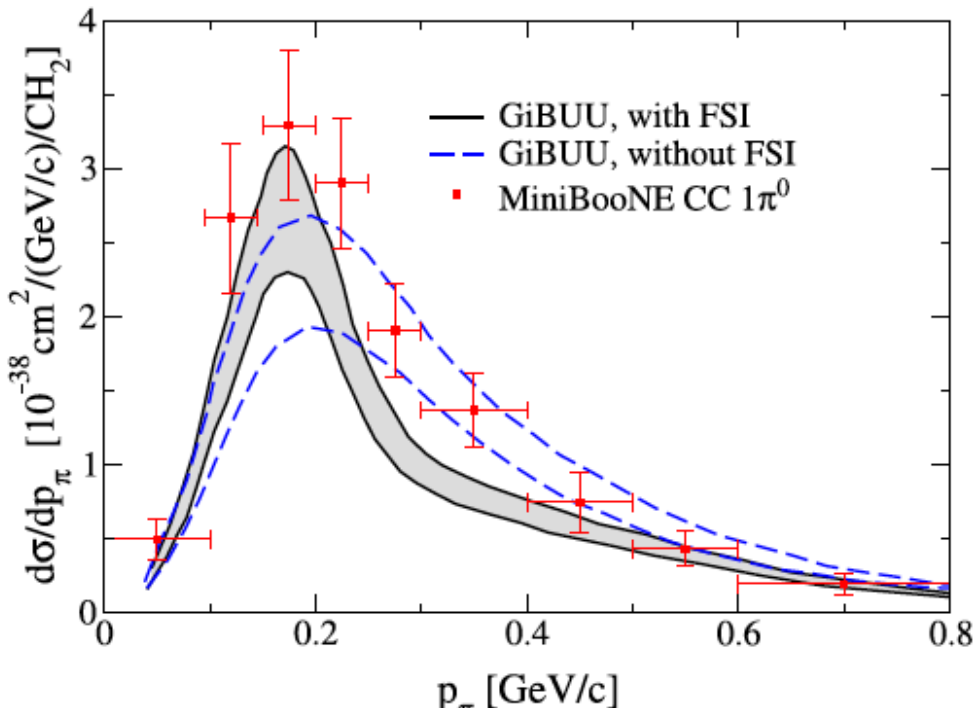
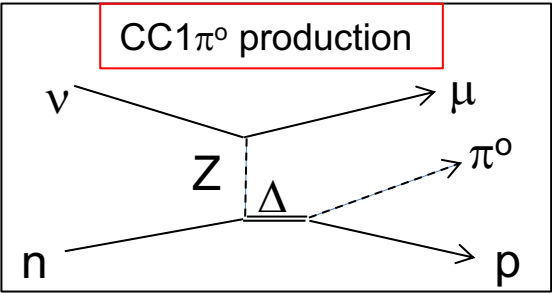




3. Pion puzzle

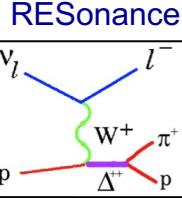
Final state interaction

- Cascade model as a standard of the community
- Advanced models are not available for event-by-event simulation



ex) Giessen BUU transport model

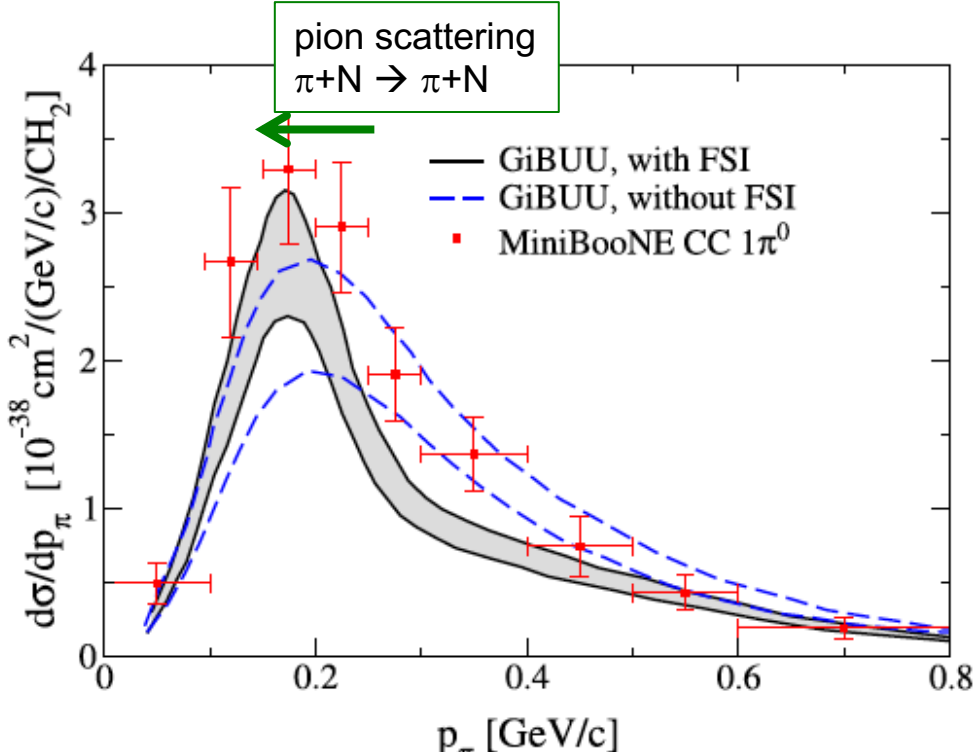
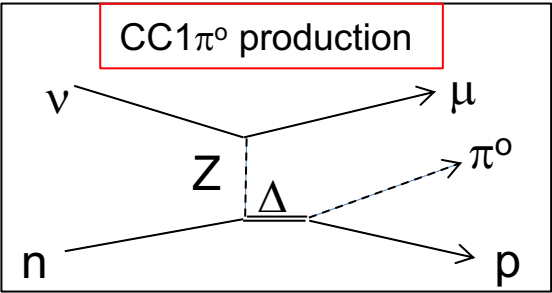
- Developed for heavy ion collision, and now used to calculate final state interactions of pions in nuclear media



3. Pion puzzle

Final state interaction

- Cascade model as a standard of the community
- Advanced models are not available for event-by-event simulation



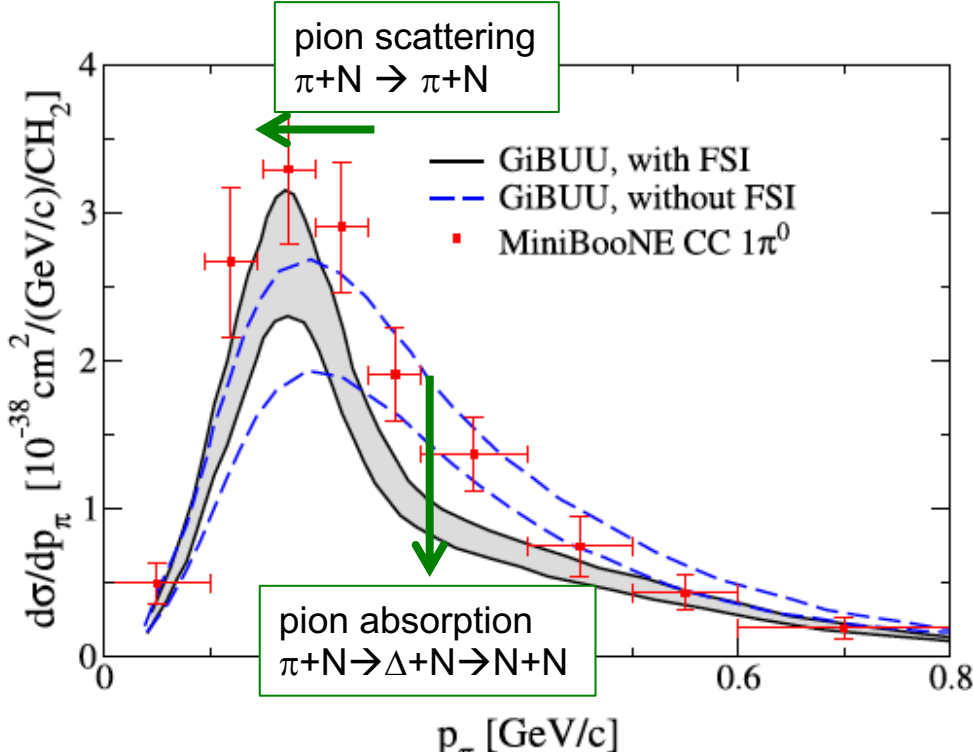
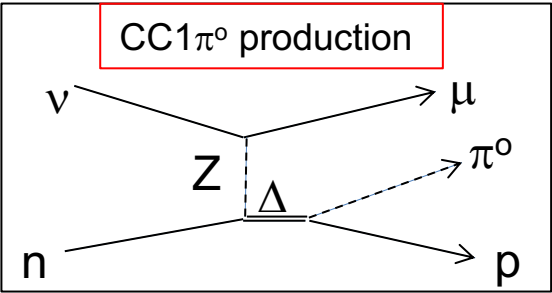
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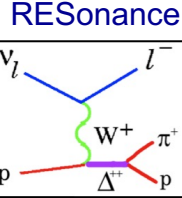
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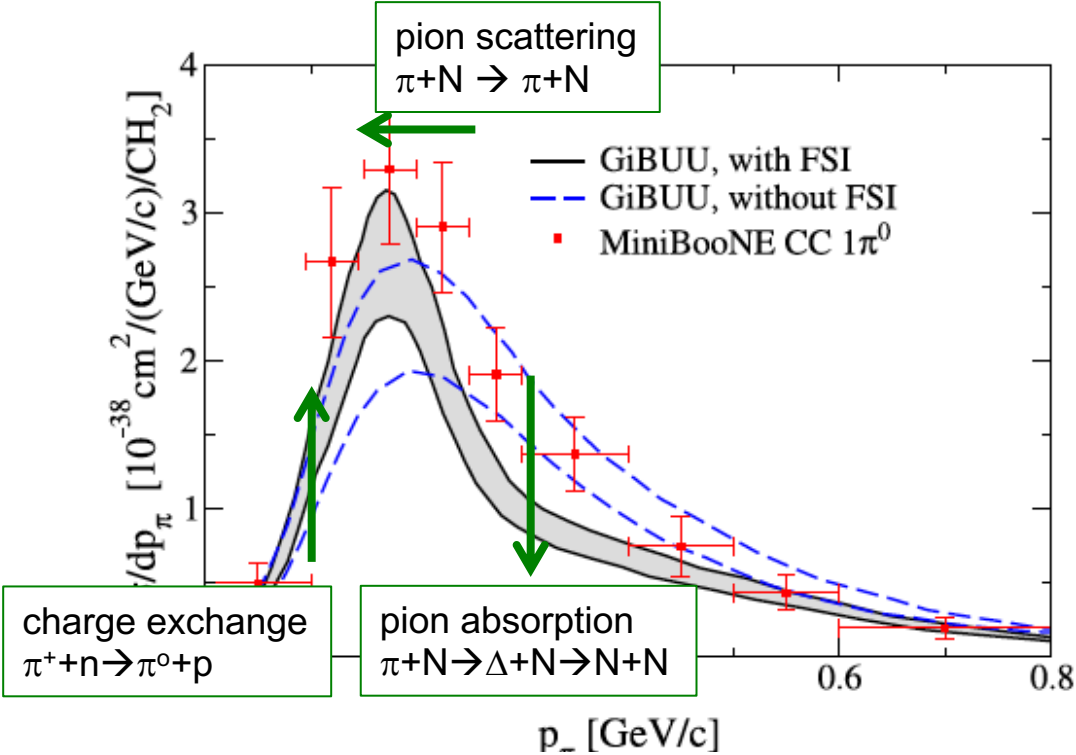
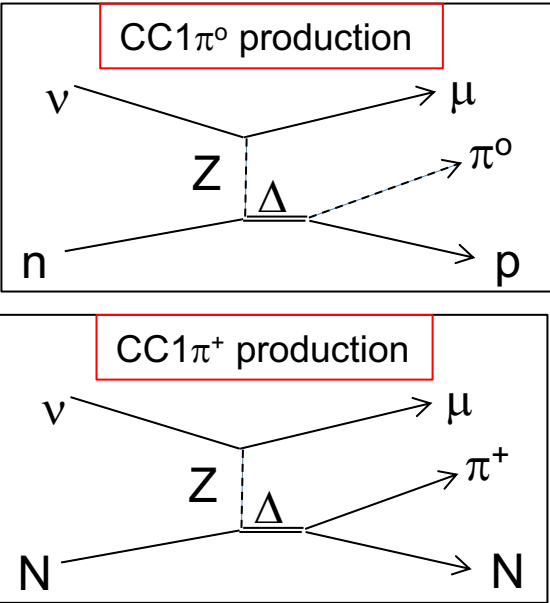
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3. Pion puzzle

Final state interaction

- Cascade model as a standard of the community
- Advanced models are not available for event-by-event simulation



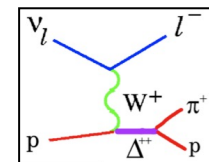
ex) Giessen BUU transport model

- Developed for heavy ion collision, and now used to calculate final state interactions of pions in nuclear media

You need to predict both

1. all pion production channels
2. all final state interaction

3. Pion puzzle

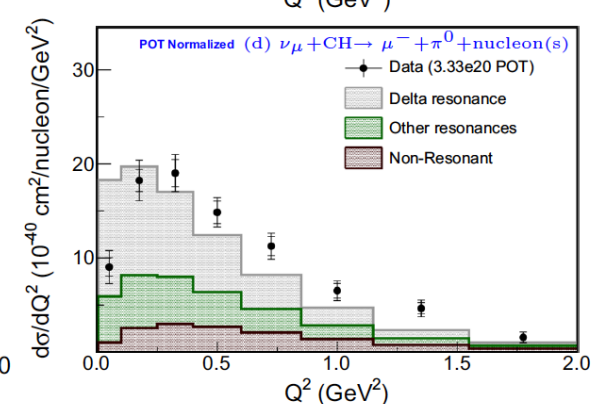
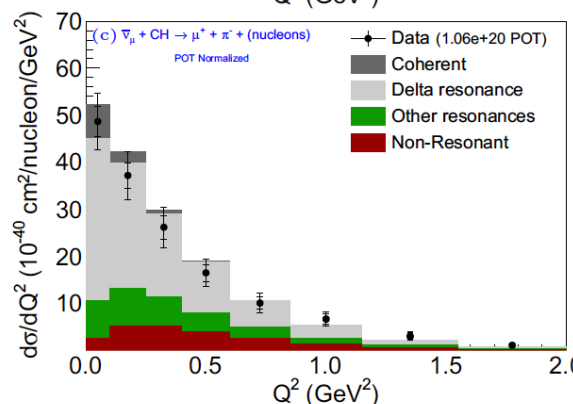
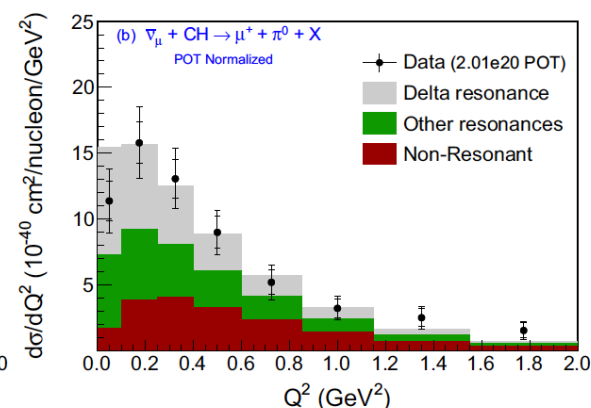
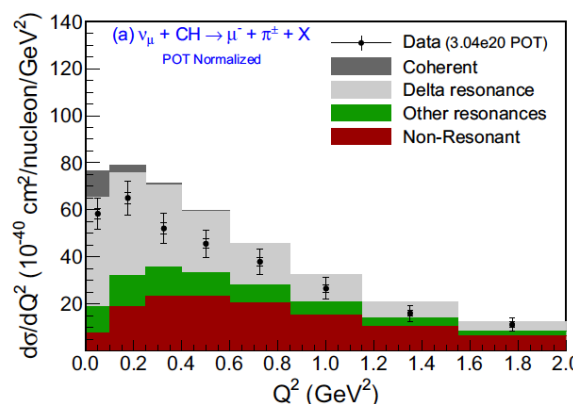


Data tension – internal: MINERvA pion data

- It is extremely difficult to tune pion and/or FSI parameters to fit all pion data
- $\nu_\mu CC \pi^\pm$, low Q^2 suppression, over-predicted
- $\nu_\mu CC \pi^0$, strong low Q^2 suppression
- $\bar{\nu}_\mu CC \pi^-$, no low Q^2 suppression
- $\bar{\nu}_\mu CC \pi^0$, low Q^2 suppression, under-predicted

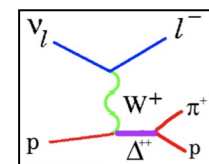
The study relies of available knobs in the generator

It looks the simulation doesn't have good “knobs” to tune



You need to predict both

1. all pion production channels
2. all final state interaction

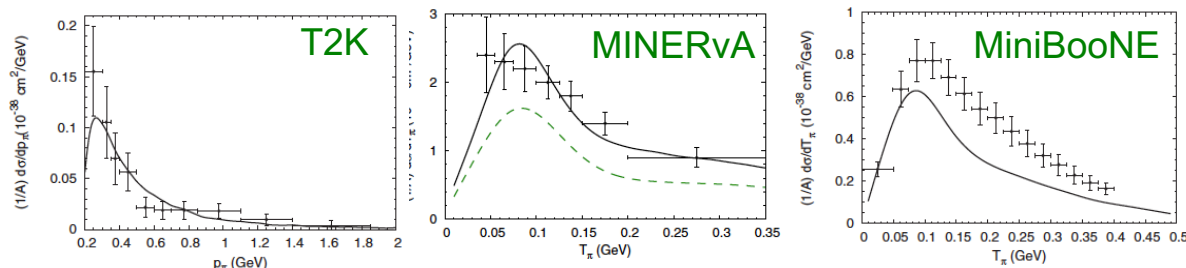


3. Data tension

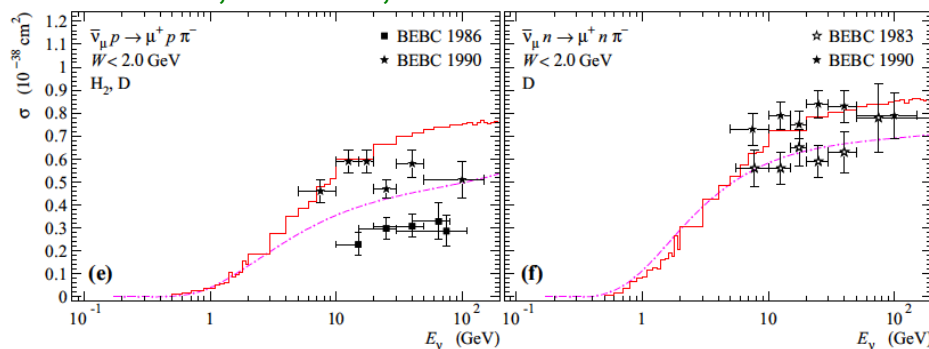
Data tension – external

- Tension between different experiments
- Tension between different targets
- Tension between different analyses

MiniBooNE vs. T2K vs. MINERvA (GiBUU)

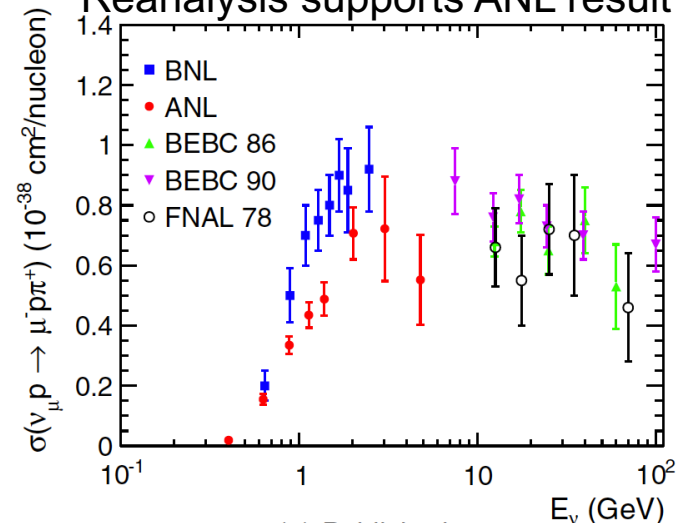


BEBC, H vs. D, 1983 vs. 1986 vs. 1990

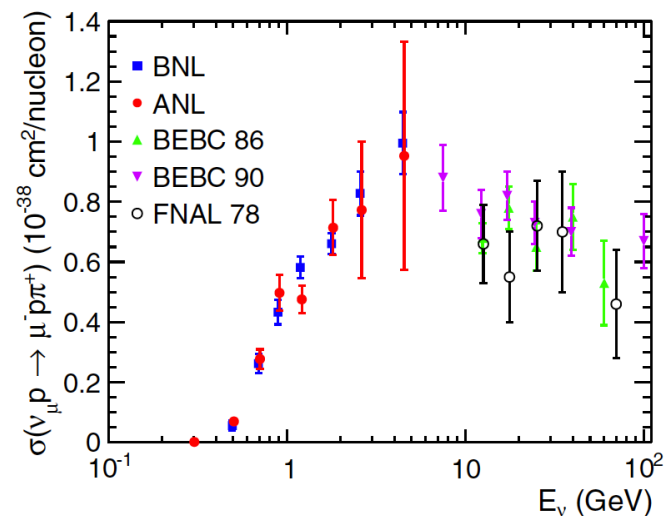


ANL vs. BNL data

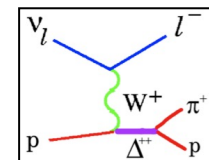
Reanalysis supports ANL result



(a) Published



(b) This analysis



3. MiniBooNE low-energy excess

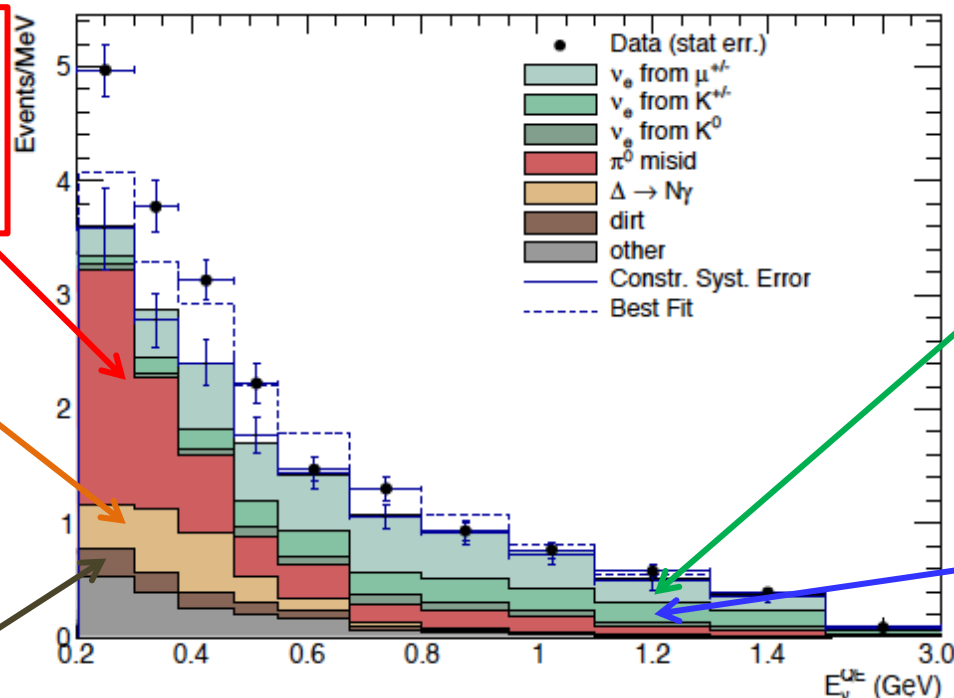
Short-baseline anomalies

- Collection of data suggesting existence of sterile neutrino in 1eV scale
- MiniBooNE has the single highest significance in all anomalies
- Resonance related backgrounds (NC π^0 , NC γ) are dominant at low-E ν_e candidate
- Important for future ν_e appearance oscillation experiments, HyperK and DUNE)

Asymmetric π^0 decay is constrained from measured NC π^0 rate ($\pi^0 \rightarrow \gamma$)

Δ resonance rate is constrained from measured NC π^0 rate

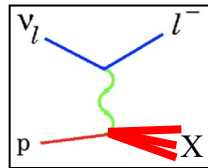
dirt rate is measured from dirt data sample



ν_e from μ decay is constrained from ν_μ CCQE measurement

ν_e from K decay is constrained from SciBooNE high energy ν_μ event measurement

1. Neutrino interaction physics - introduction
2. Charged-Current Quasi-Elastic (CCQE) interaction
3. Neutrino baryonic resonance interaction
- 4. Neutrino shallow- and deep-inelastic scatterings**
5. Conclusions



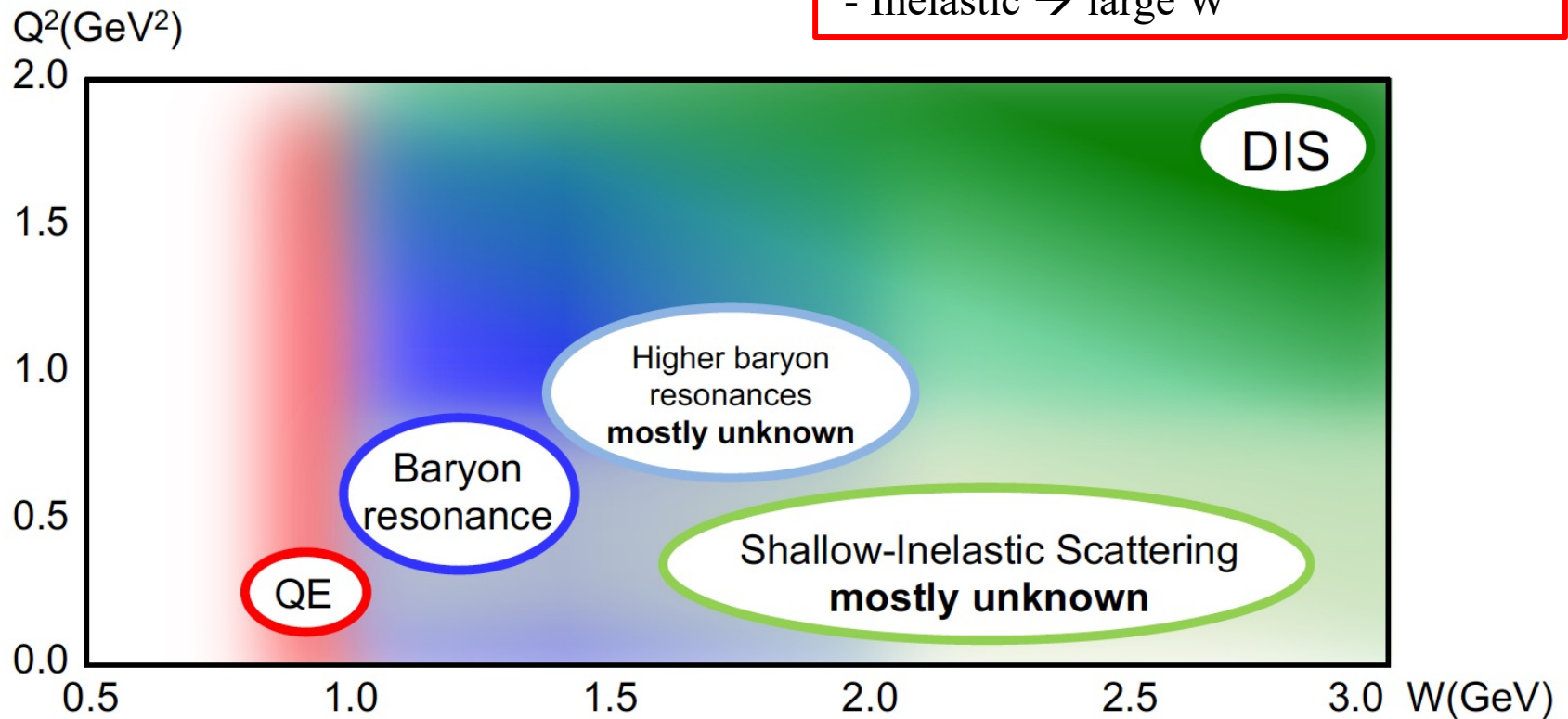
4. Shallow Inelastic Scattering (SIS)

Cross section

- Higher resonances and hadron dynamics
- Quark-Hadron duality (low Q^2 , low W DIS)
- Nuclear dependent DIS

Neutrino experiments around 1-10 GeV are not quite DIS yet

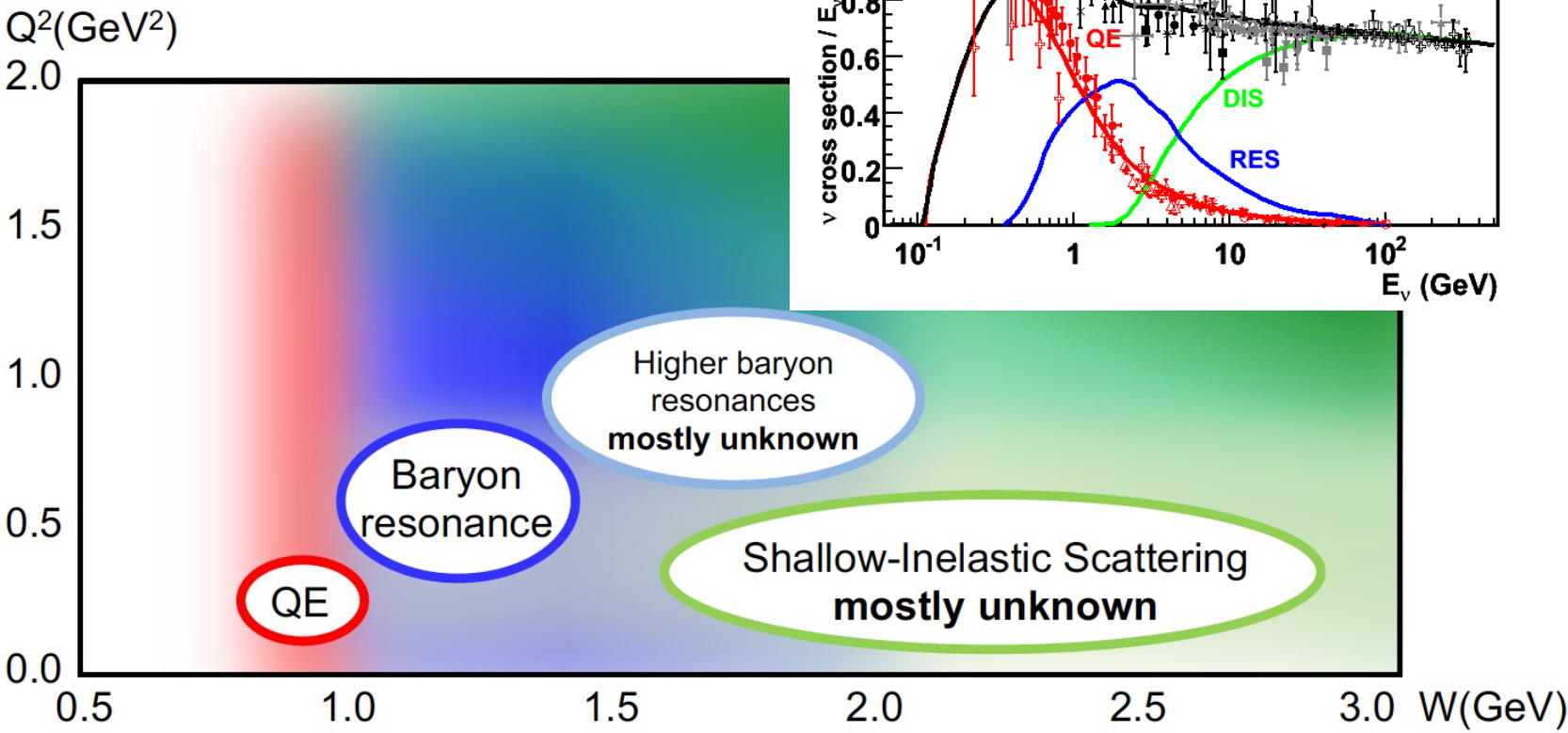
- Shallow $\rightarrow$ low Q^2
- Inelastic $\rightarrow$ large W



4. Shallow Inelastic Scattering (SIS)

Cross section

- Higher resonances and hadron dynamics
- Quark-Hadron duality (low Q^2 , low W DIS)
- Nuclear dependent DIS



4. Higher baryonic resonances

Cross section

- Higher resonances and hadron dynamics
- Quark-Hadron duality (low Q^2 , low W DIS)
- Nuclear dependent DIS

DCC model

- Channels are coupled (πN , $\pi\pi N$, etc), total amplitude is conserved
- Most of axial form factors are unknown

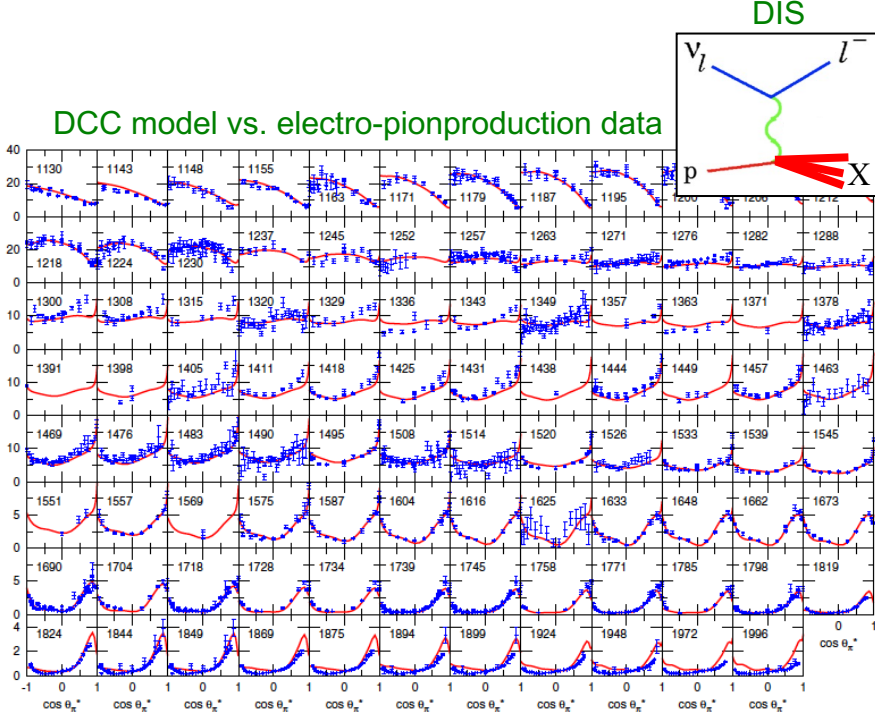
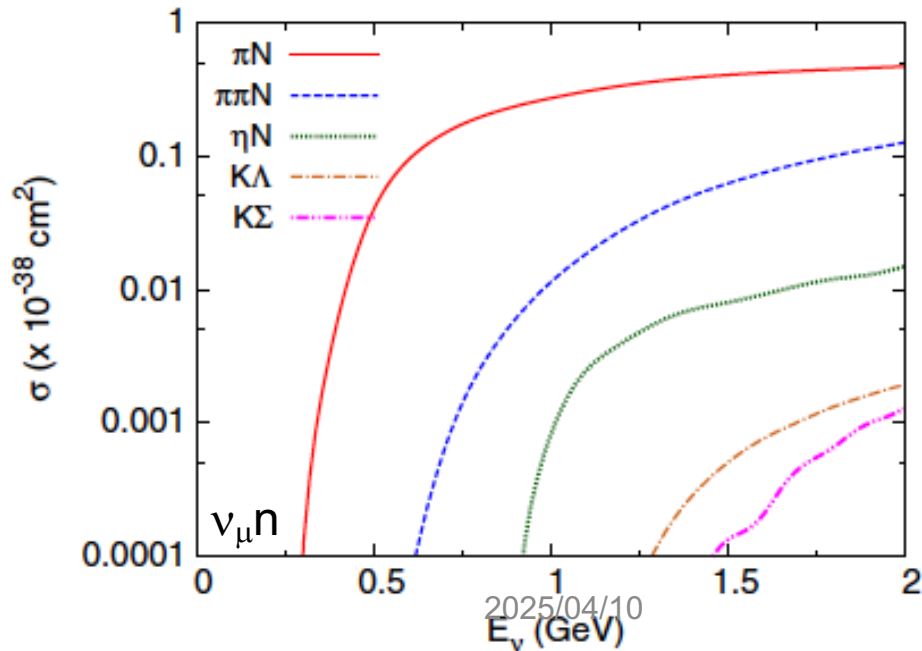
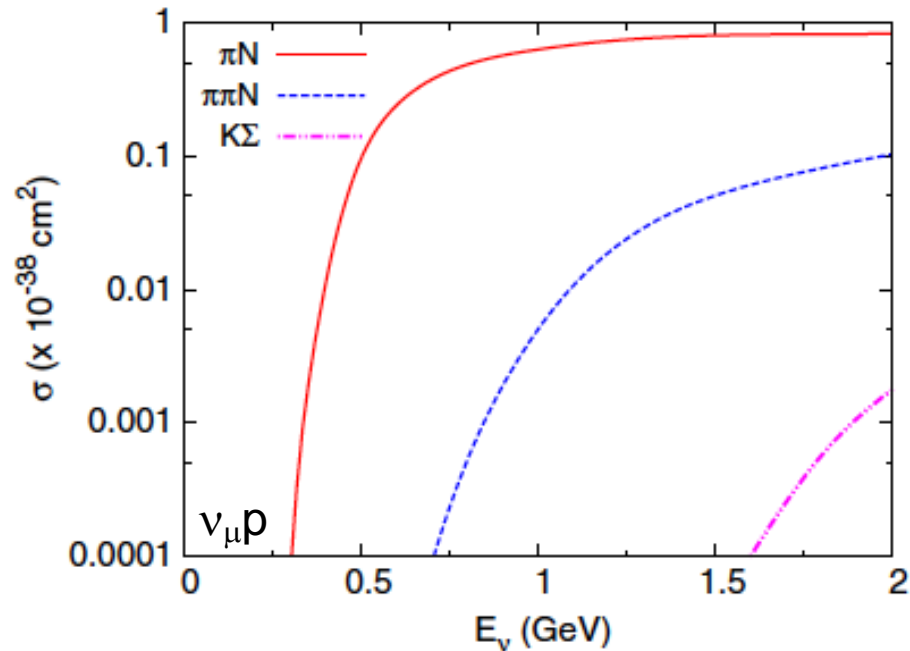


FIG. 8 (color online). Unpolarized differential cross sections, $d\sigma/d\Omega_\pi^*$ ($\mu\text{b/sr}$), for $\gamma n \rightarrow \pi^+ p$. The data are from Refs. [55–78].

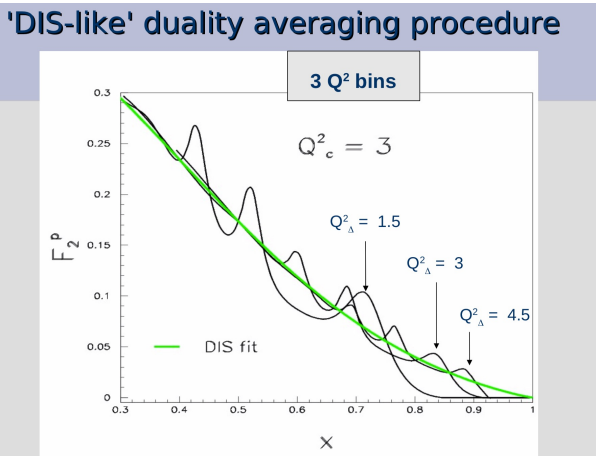
4. Quark-Hadron duality

Cross section

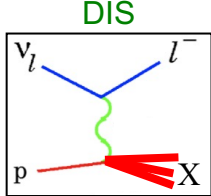
- Higher resonances and hadron dynamics
- Quark-Hadron duality (low Q^2 , low W DIS)
- Nuclear dependent DIS

Bodek-Yang correction is a phenomenological model to reproduce duality-like behavior, accepted by all neutrino simulation

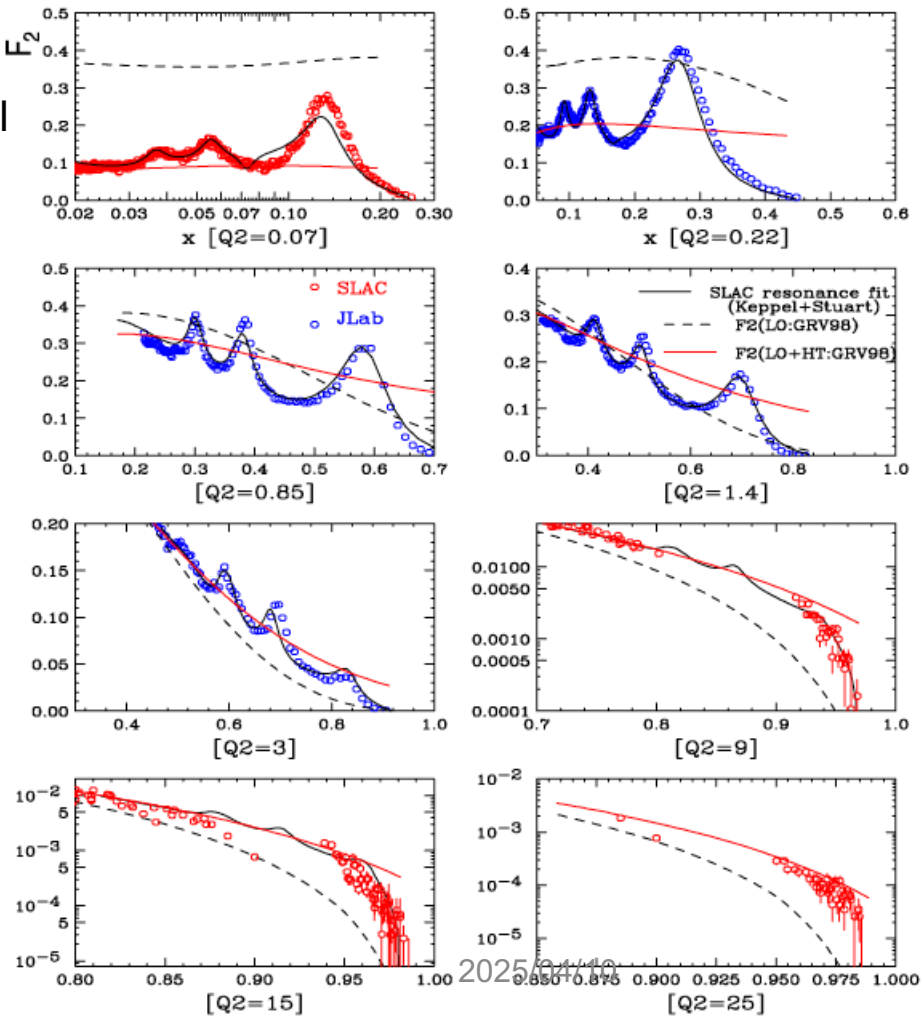
DIS $\neq$ Bjorken limit
DIS = Q^2 average of all resonances



Nachtmann variable $\xi = \frac{2x}{\left(1 + \sqrt{1 + \frac{4x^2 M^2}{Q^2}}\right)}$



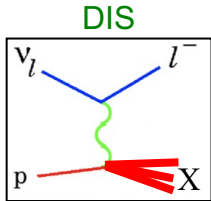
Proton F2 function GRV98-BY correction vs. data



4. Nuclear dependent DIS

Cross section

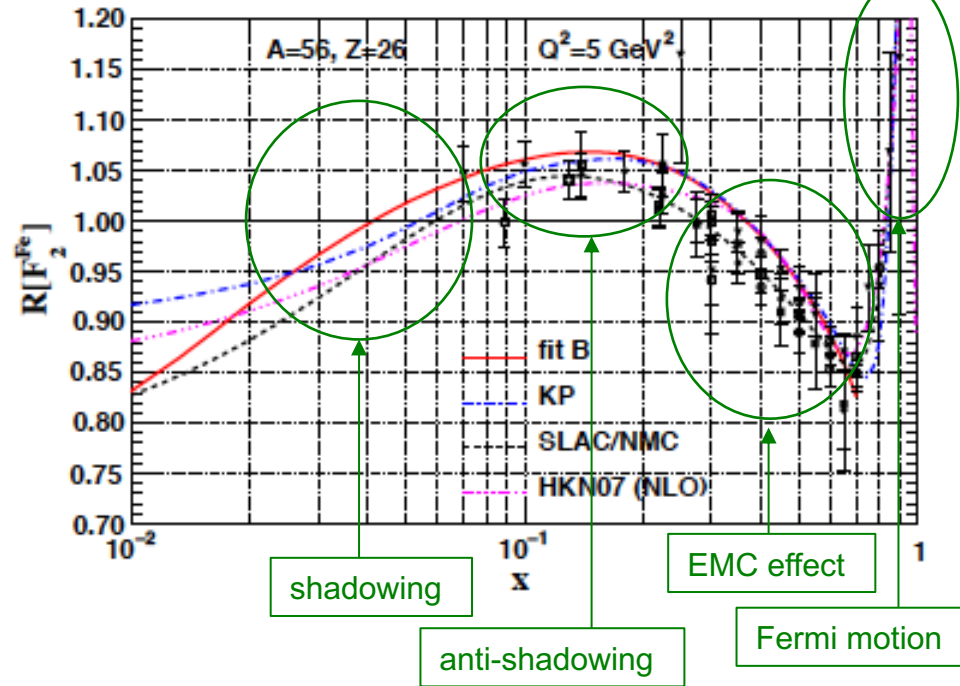
- Higher resonances and hadron dynamics
- Quark-Hadron duality (low Q^2 , low W DIS)
- Nuclear dependent DIS



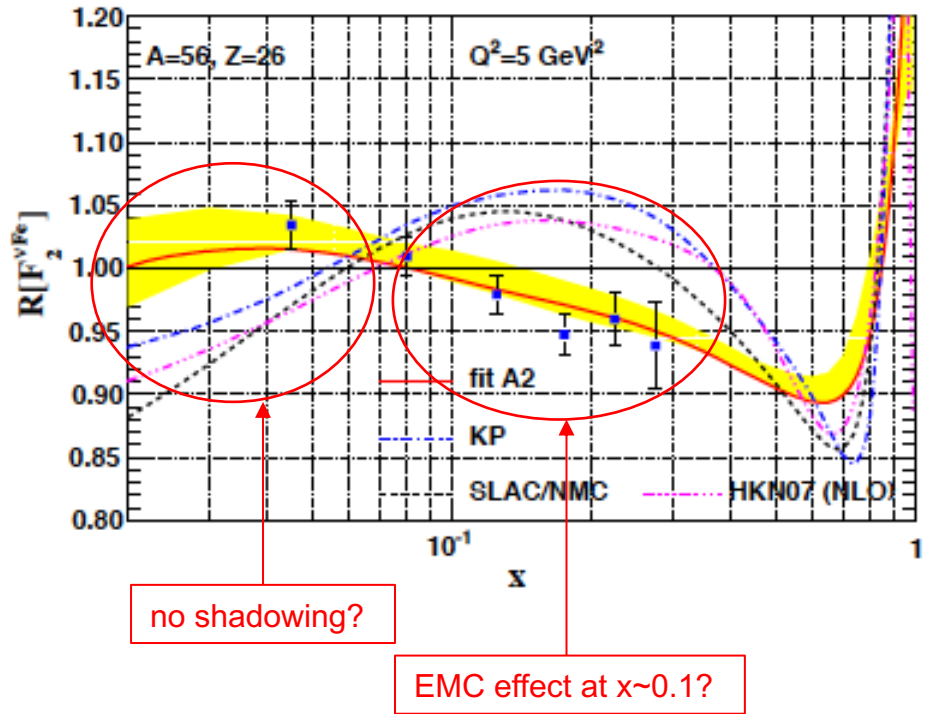
Nuclear PDF

- Shadowing, EMC effect, Fermi motion
- Likely due to nucleon dynamics in nucleus
- Various models describe charged lepton data
- Neutrino data look very different

e^+ -Fe nuclear correction factor



ν -Fe nuclear correction factor



4. Low-W hadronization model

KNO-scaling based model

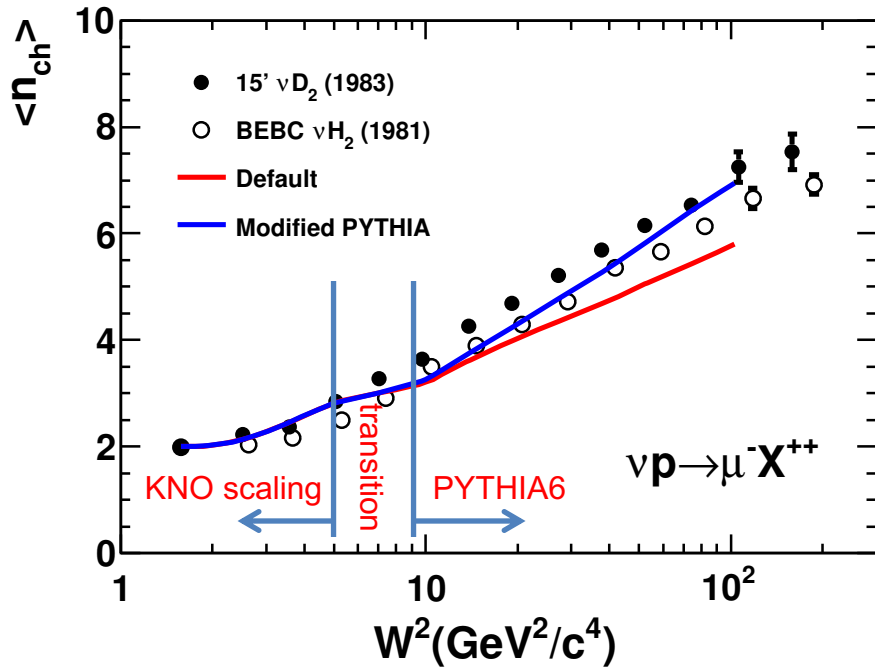
- Data-driven model
- Averaged charged hadron multiplicity $\langle n_{ch} \rangle$
- Averaged neutral hadron multiplicity is chosen from isospin
- Variance of multiplicity is chosen from KNO-scaling law.

$$\langle n_{ch} \rangle = a_{ch} + b_{ch} \cdot \ln(W^2)$$

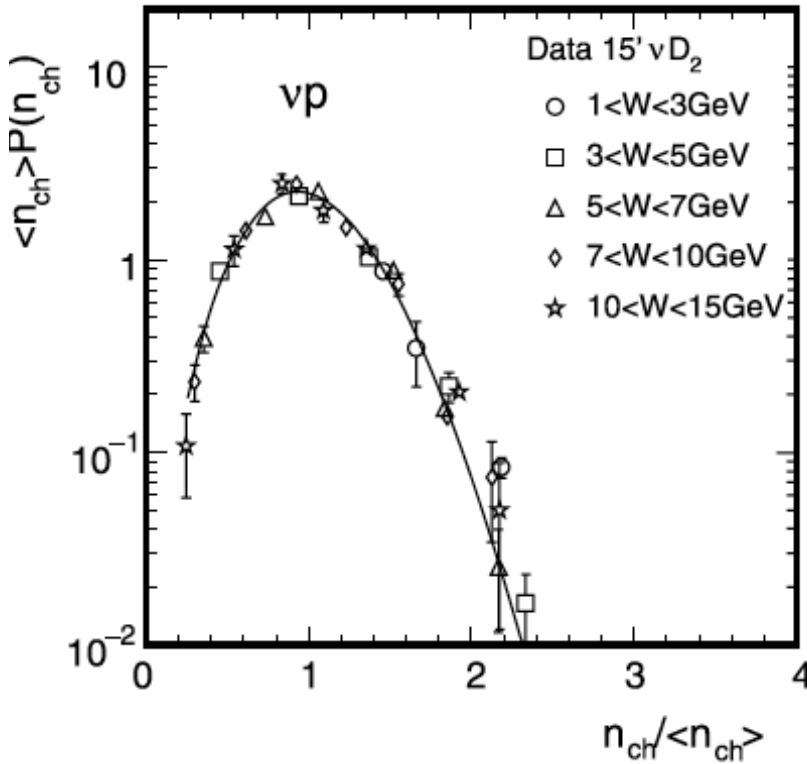
$$\pi^+ + \pi^- : \pi^0 = 2 : 1$$

$$\langle n \rangle \cdot P(n) = \frac{2e^{-c} c^{cn/\langle n \rangle + 1}}{\Gamma(cn/\langle n \rangle + 1)}$$

Neutrino average charged hadron multiplicity



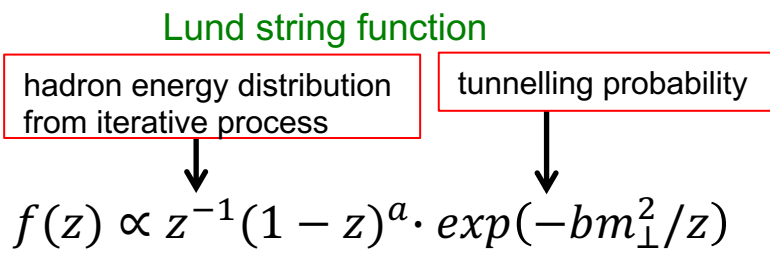
charged hadron multiplicity dispersion



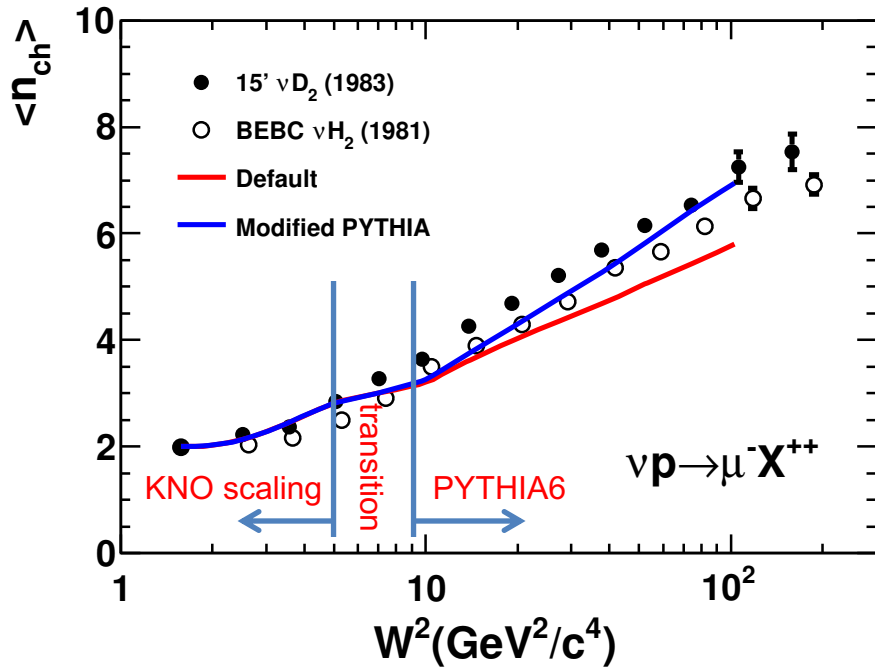
4. High-W hadronization

PYTHIA6

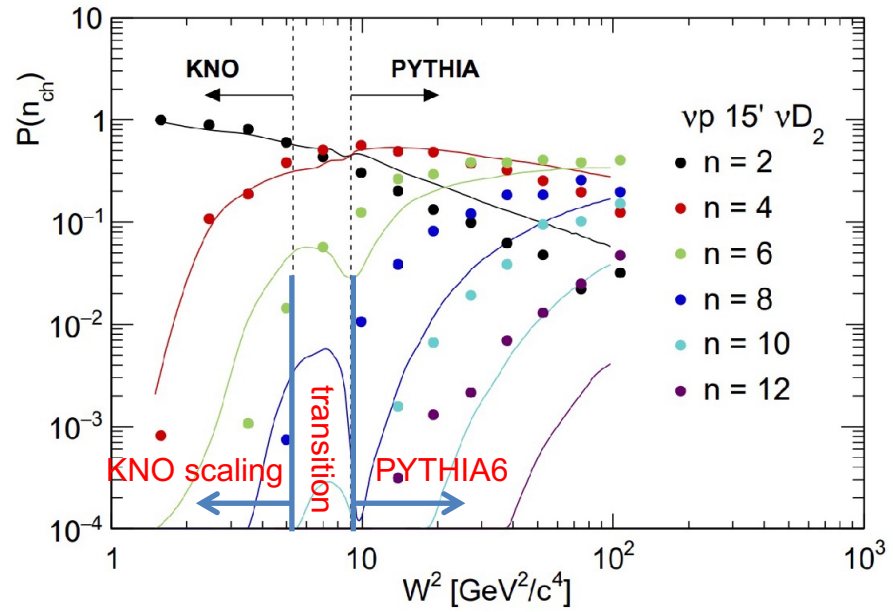
- Lund string function to model $q - \bar{q}$ evolution
- Tuned function can reproduce $\langle n_{ch} \rangle$ data
- So far, PYTHIA tuning cannot fix dispersion discrepancy between Low-W and high-W hadronization models



Neutrino average charged hadron multiplicity



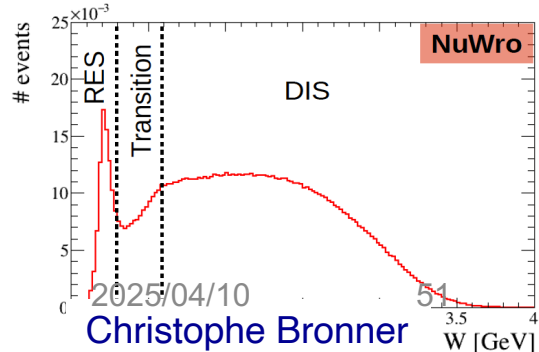
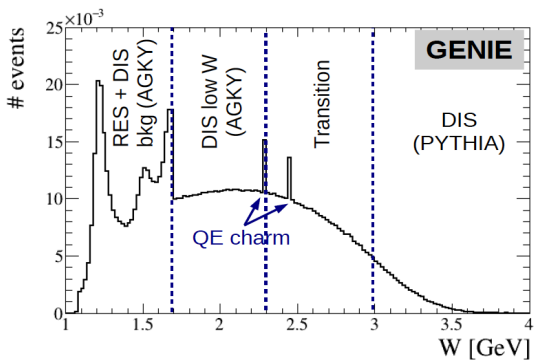
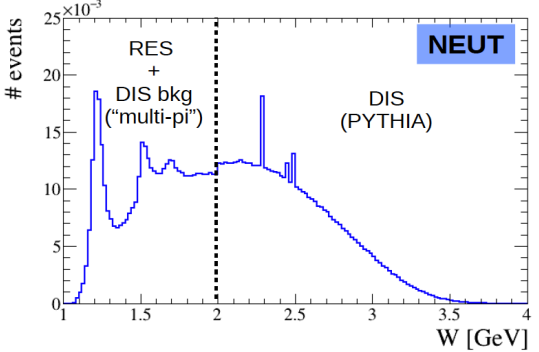
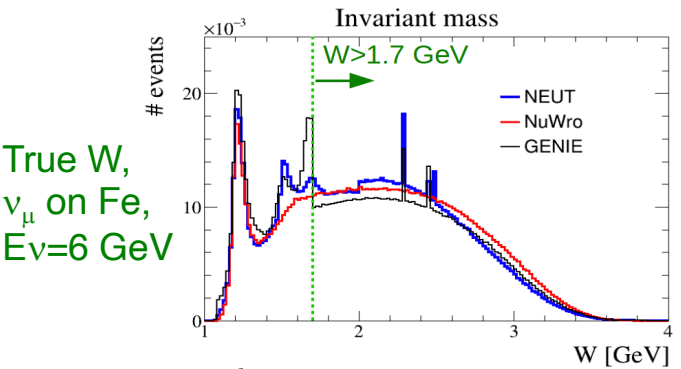
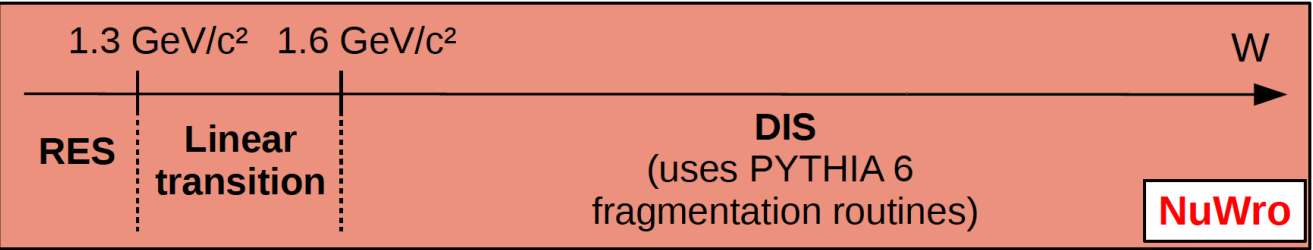
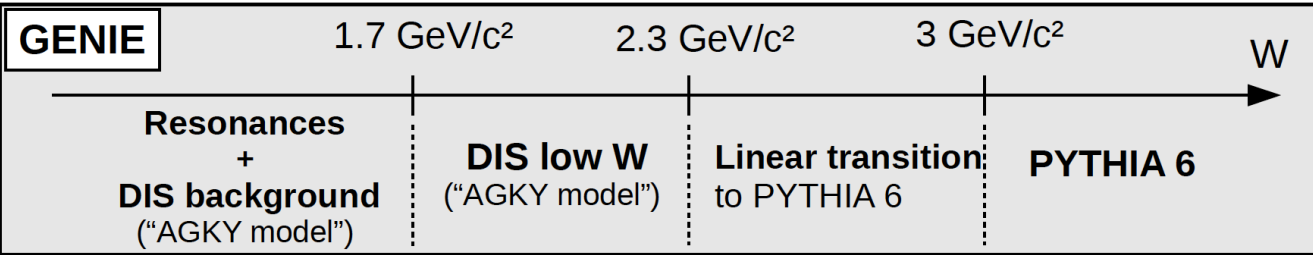
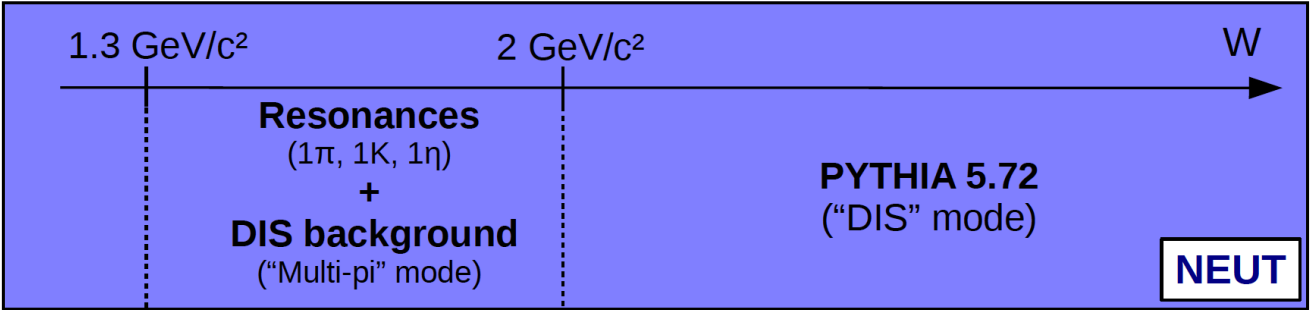
ν -p topological cross section (GENIE)



1. SIS in event generators

Difficult to connect different models

- Cross-section models
- Hadronization models



4. NuSTEC

NuInt conference series

- Oct. 6-10 2025, Mainz (Germany)

<https://nustec.fnal.gov/nuint-conference-series/>

Neutrino Scattering Theory-Experiment Collaboration

<https://nustec.fnal.gov/>

NuSTEC CTGWG seminar and CEWG meetings
<https://indico.fnal.gov/category/990/>

Cross-theory-
generator WG
(CTGWG)

Workshops
and Schools

Cross-
experiment WG
(CEWG)

NuSTEC

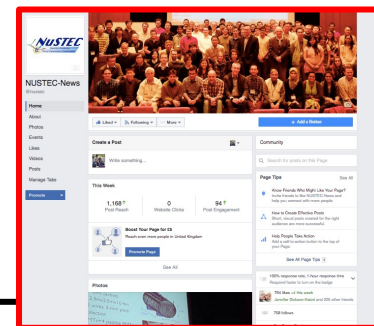
Publications

Long-term
community planning

Outreach



NuSTEC-News Facebook page
<https://www.facebook.com/nuxsec>



NuSTEC, PPNP100(2018)1
<https://www.sciencedirect.com/science/article/pii/S0146641018300061>



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Conclusion

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1 to 10 GeV neutrino interaction measurements are crucial to successful next-generation neutrino oscillation experiments (DUNE, Hyper-K, ORCA, IC-Upgrade)

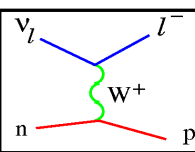
Nucleon correlation physics drastically change neutrino cross sections in QE dominant region.

Currently, the community don't understand tensions in resonant single pion production data.

Systematic errors on SIS and DIS region are important for future experiments.

Thank you for your attention!

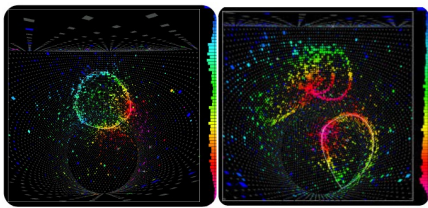
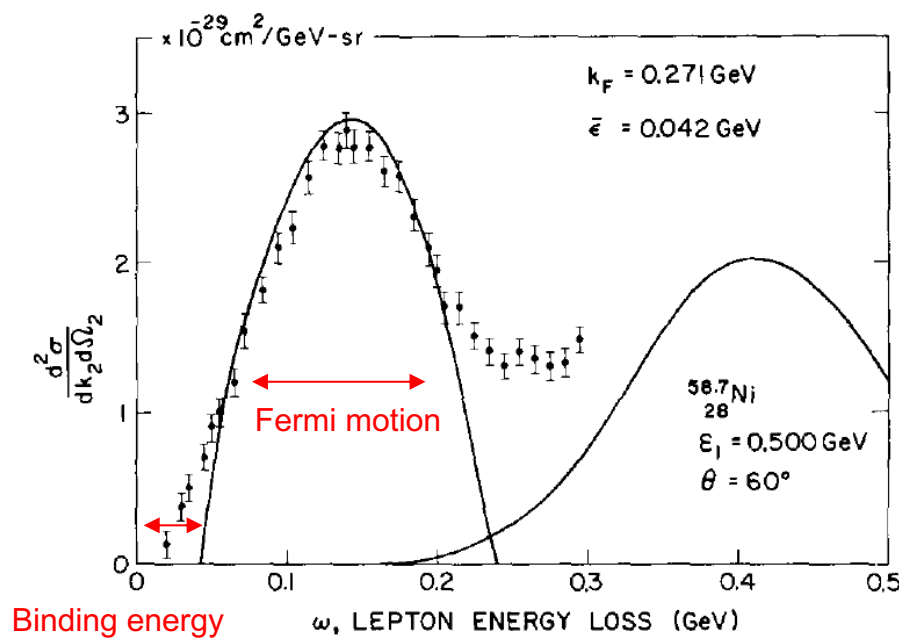
Backup



2. Fermi motion

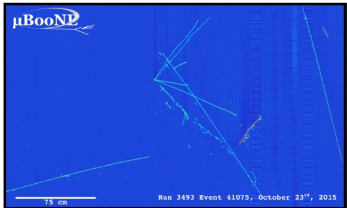
Fermi motion

- Measured energy is smeared from the true energy if you assume nucleon at rest
- High resolution detector can measure all outgoing hadrons
 → initial nucleon momentum can be reconstructed (no Fermi motion smearing)



Cherenkov detectors:
 Assuming QE interaction
 Using lepton only

$$E_{QE} = \frac{2M\epsilon + 2ME_l - m_l^2}{2(M - E_l + |k_l| \cos \theta_l)}$$



Tracking detectors:
 Calorimetric sum
 Using All detected particles

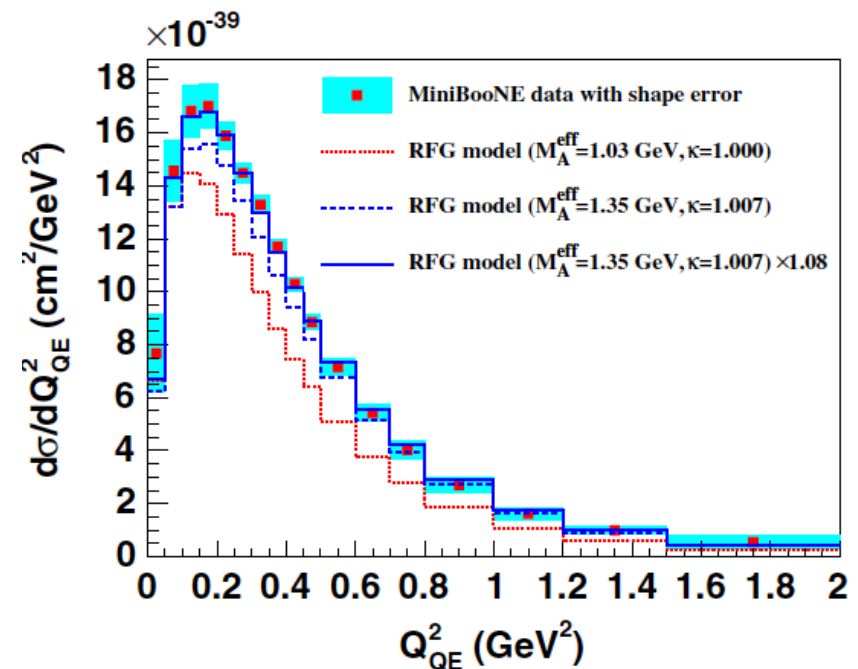
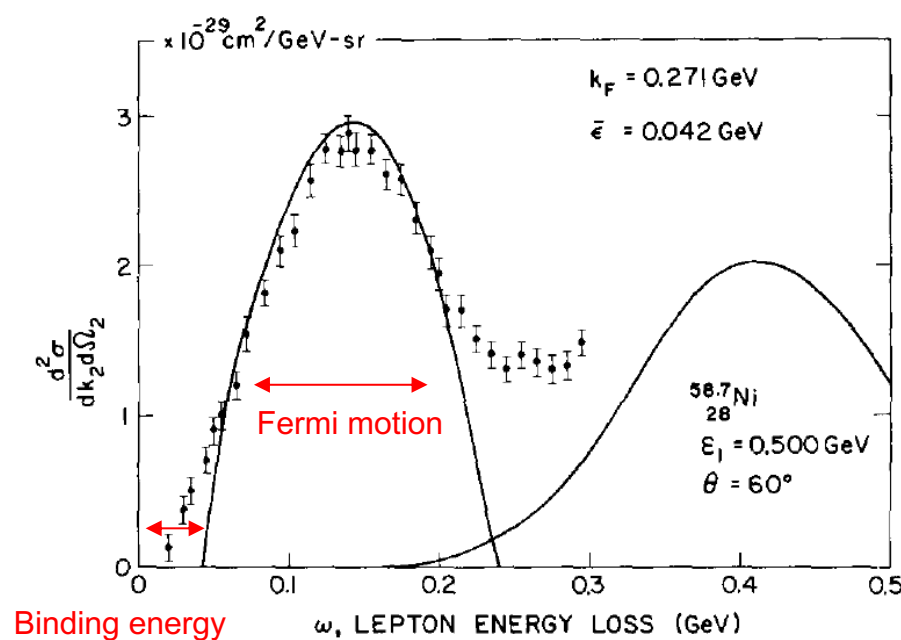
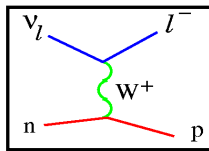
$$E_{cal} = E_l + E_p^{kin} + \epsilon$$

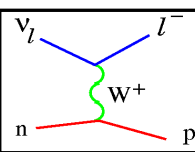
[1p0π]

2. Pauli blocking

Pauli blocking

- Low momentum transfer reaction is forbidden.
- data show more suppression than what Pauli blocking can $\rightarrow$ RPA(?)
- In the global Fermi gas model, Pauli blocking looks unphysical

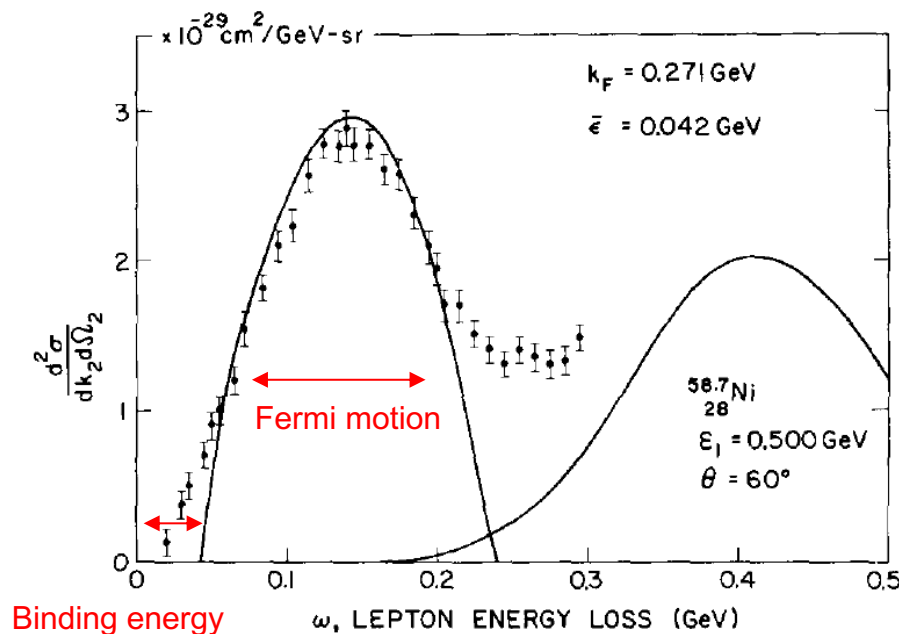




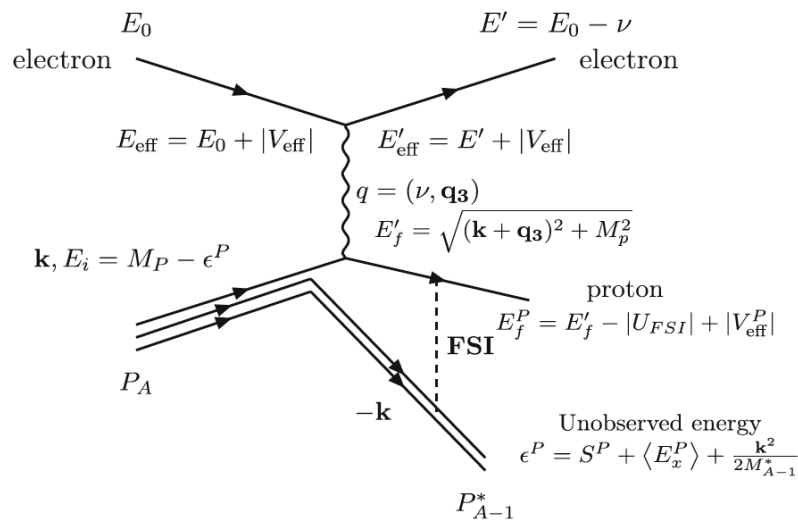
2. Nuclear Shell structure and binding energy

Binding energy ~ unobserved energy

- Energy to cost to release 1 nucleon, not constant
- Separation energy + excitation energy + recoil energy
 - Separation energy: energy to release 1 nucleon from the shell (~15 MeV, depends)
 - Excitation energy: energy used to excite leftover target nucleus (~1 MeV)
 - Recoil energy: kinetic energy of recoil target nucleus (~2-3 MeV)



Electron scattering on proton



2. New paradigm of lepton scattering experiments

Flux-averaged differential cross-section

- Incomplete kinematics, reconstruction of E_ν , Q^2 , q^3 , W , x , y ,... depends on models

Electron scattering

- well defined energy, well known flux
- reconstruct energy-momentum transfer
- measure each process

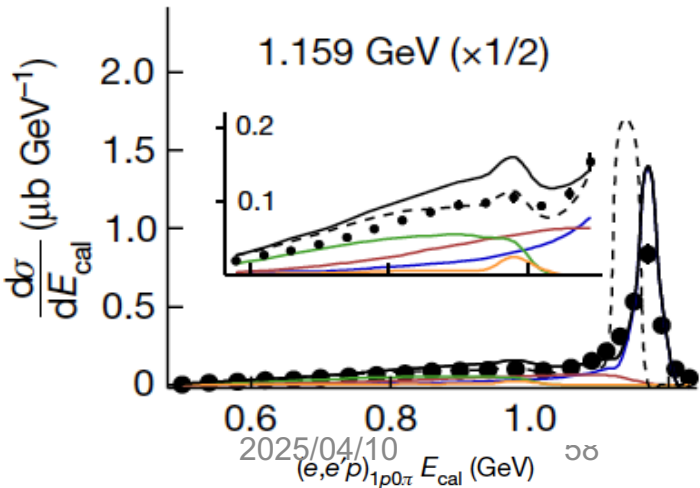
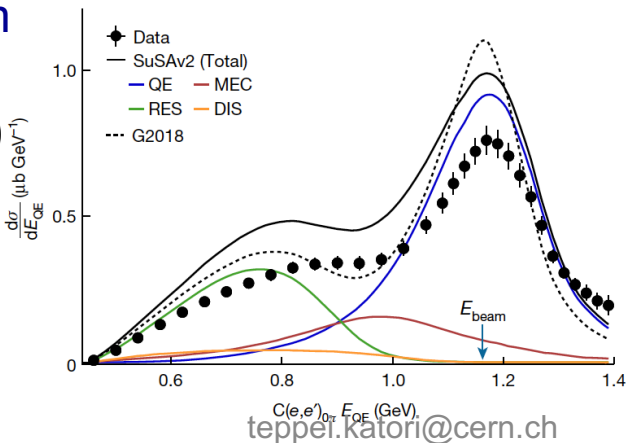
Neutrino experiment don't reconstruct E_ν (and Q^2) with great precision

Reconstructed beam electron energy spectrum by

- QE kinematic (HyperK)
- Calorimetric (DUNE)

$$E_\nu^{QE} = \frac{ME_\nu - 0.5m_\mu^2}{M - E_\mu + p_\mu \cos\theta}$$

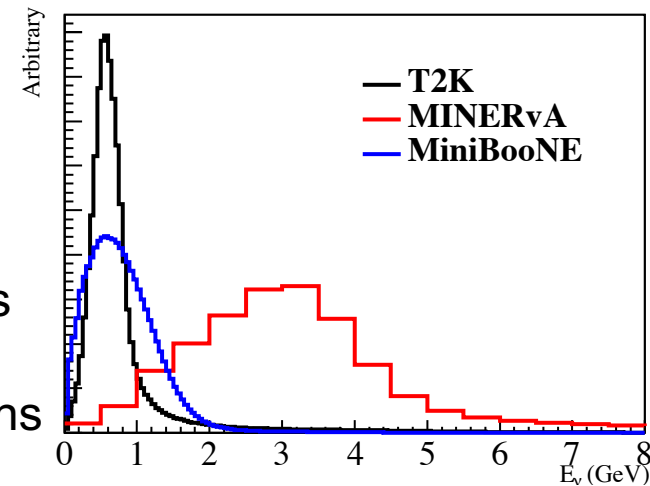
$$E_\nu^{Cal} = E_\mu + \sum^{all} E_{had}^i$$



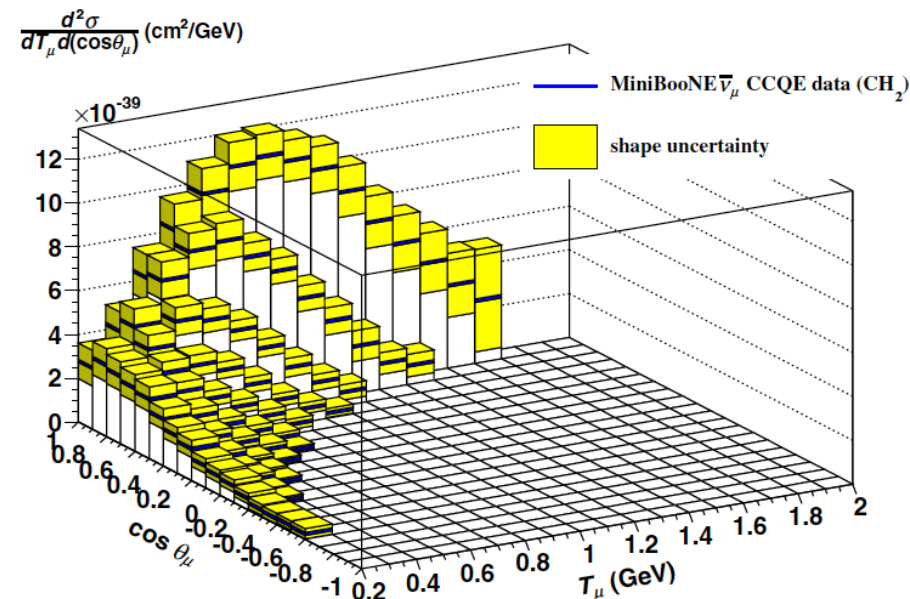
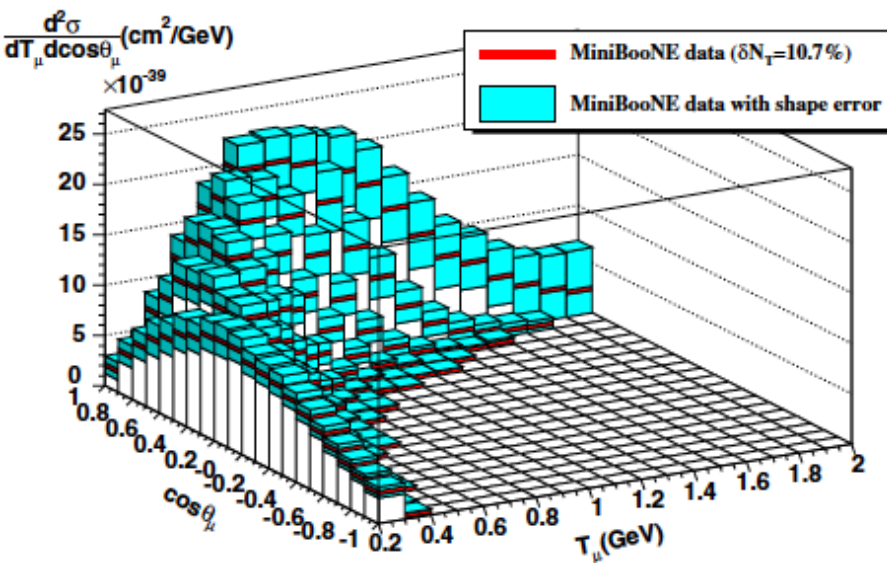
5. MiniBooNE

Mineral oil (CH₂) Cherenkov detector

- 4 π coverage, $\langle E \rangle \sim 800$ MeV beam up to 2 GeV
- Measure Cherenkov radiations from charged particles
- Some calorimetric (scintillation)
- Measured first flux-integrated differential cross sections
- Solved CCQE puzzle



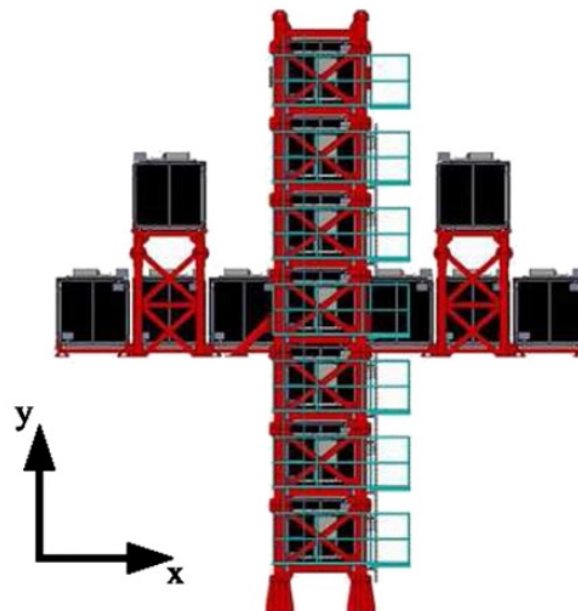
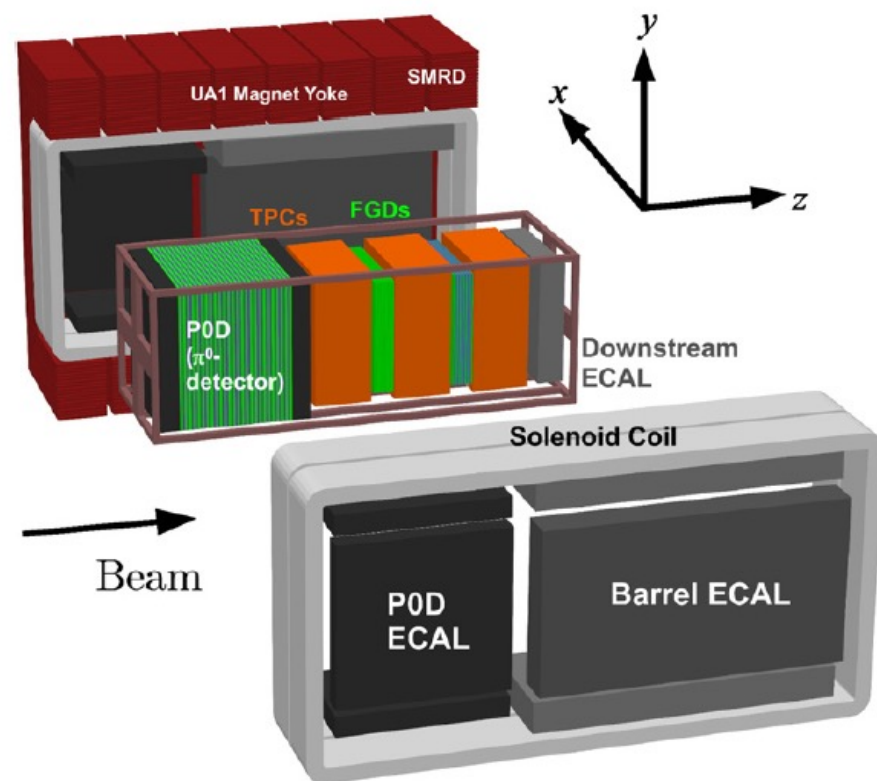
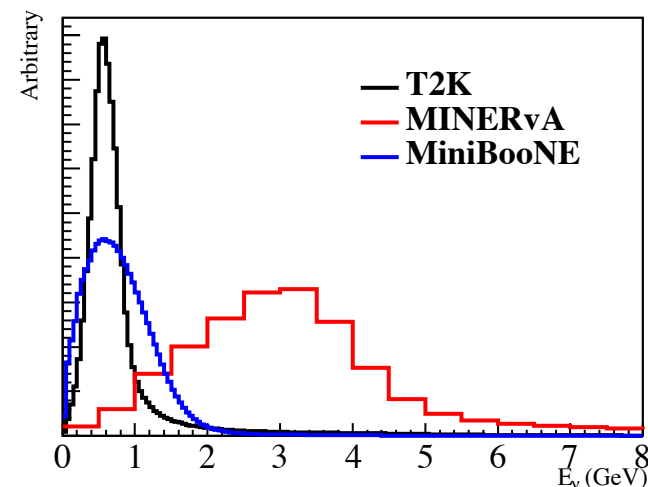
neutrino and anti-neutrino CCQE-like double differential cross sections



5. T2K near detector complex

INGRID, FGD, P0D, ECal, TPC, SMRD

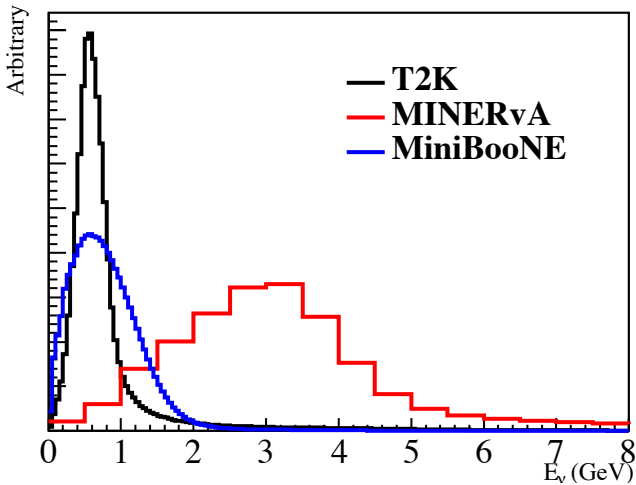
- Plastic scintillation trackers (except gas TPC)
- 0.2T magnet for momentum measurement
- $\langle E \rangle \sim 600$ MeV off-axis beam
- variety of targets (CH, H₂O, Pb, Ar)
- Limited coverage (combination of sub-detectors)



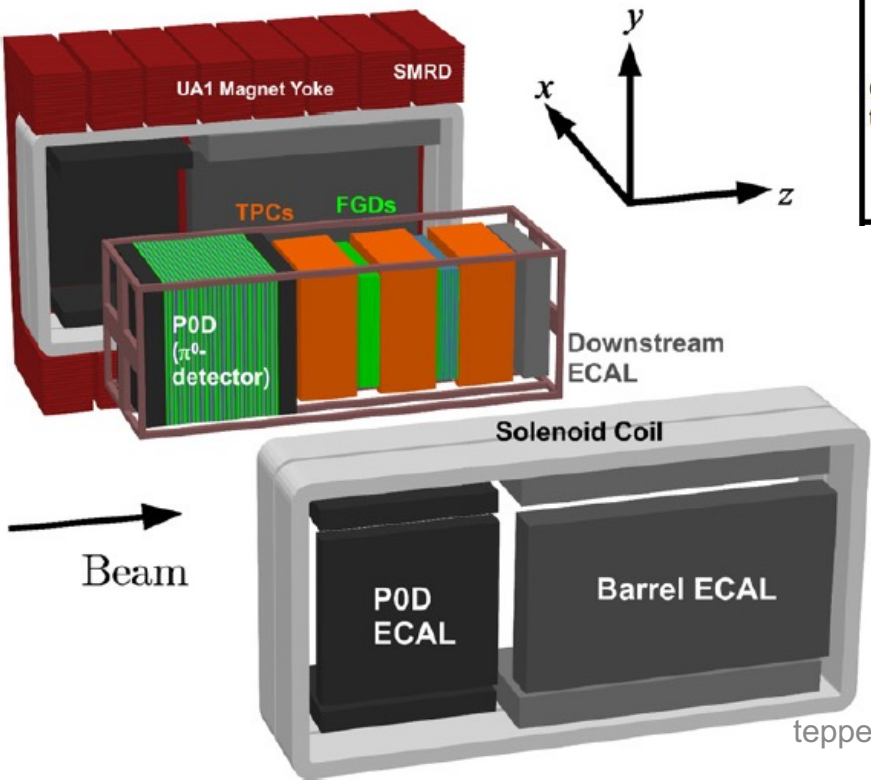
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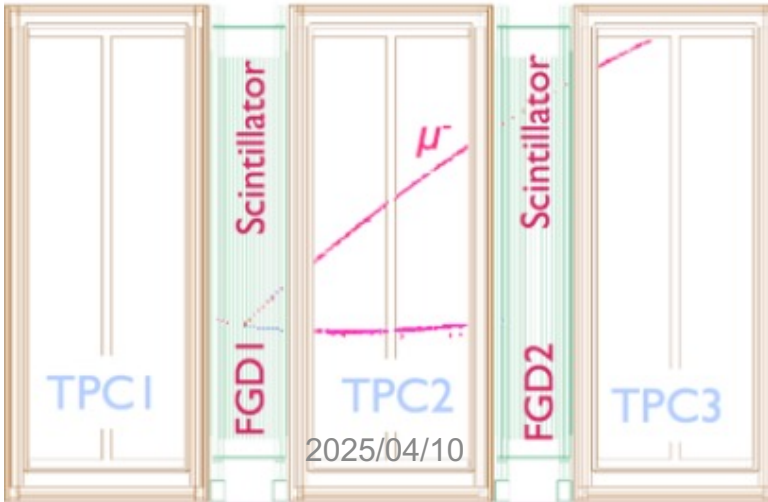


neutrino CC0 π double differential cross sections



	(1)	(2)	(3)	(4a)	(4b)
CCQE topology					

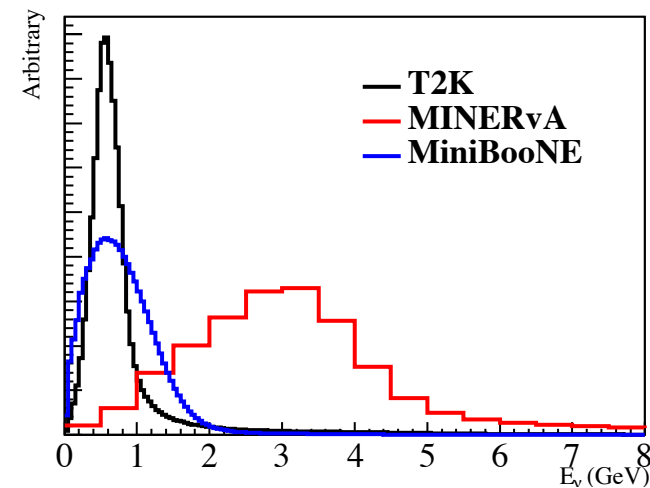
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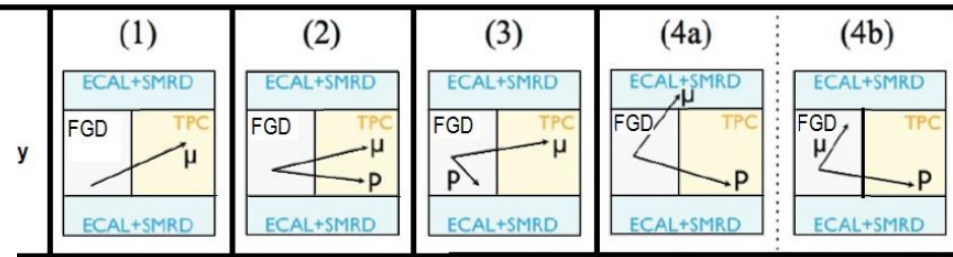
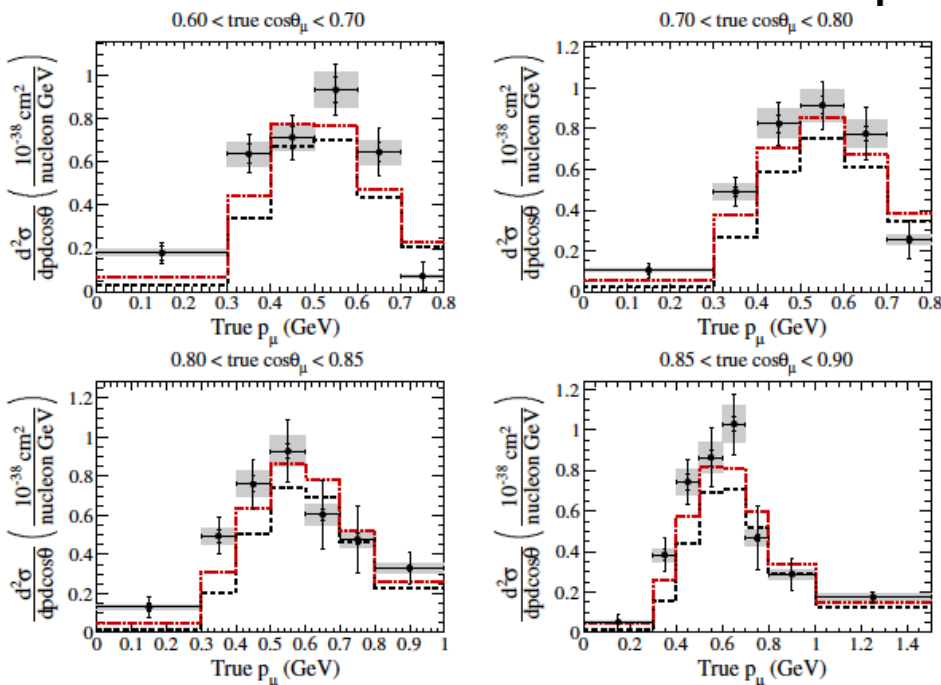
5. T2K near detector complex

INGRID, FGD, P0D, ECal, TPC, SMRD

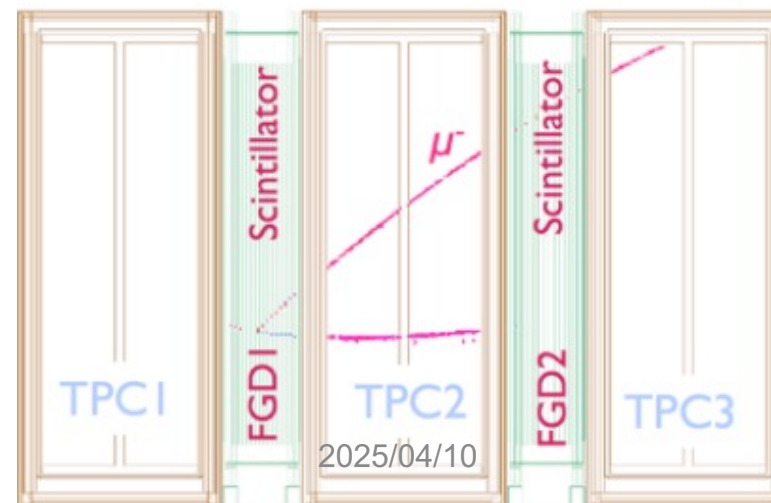
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neutrino CC0 π double differential cross sections



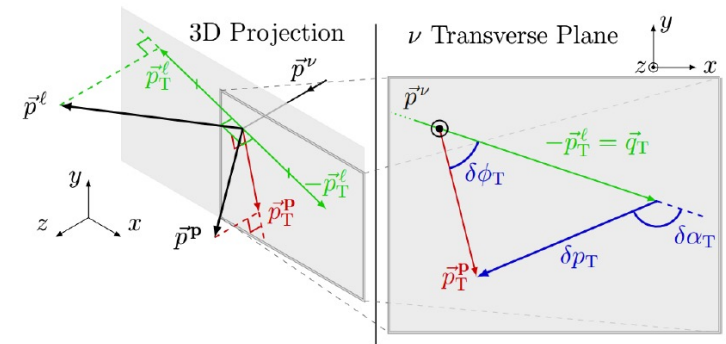
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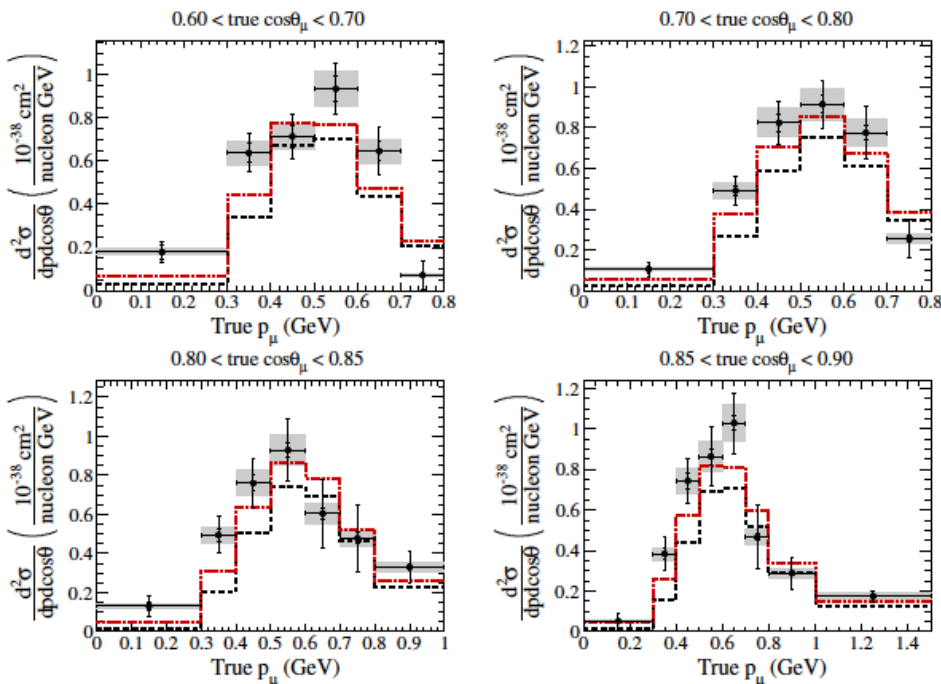
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INGRID, FGD, P0D, ECal, TPC, SMRD

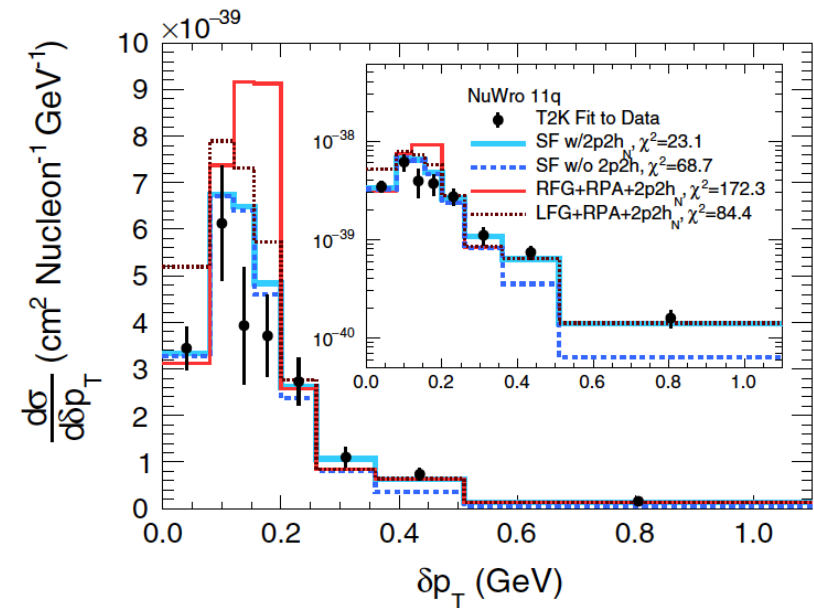
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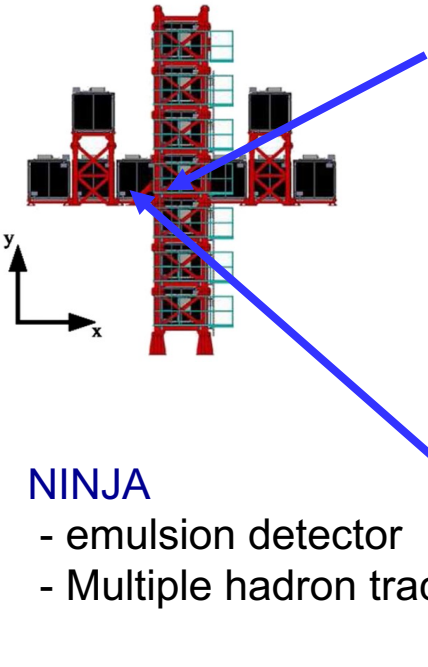
neutrino CC0 π double differential cross sections



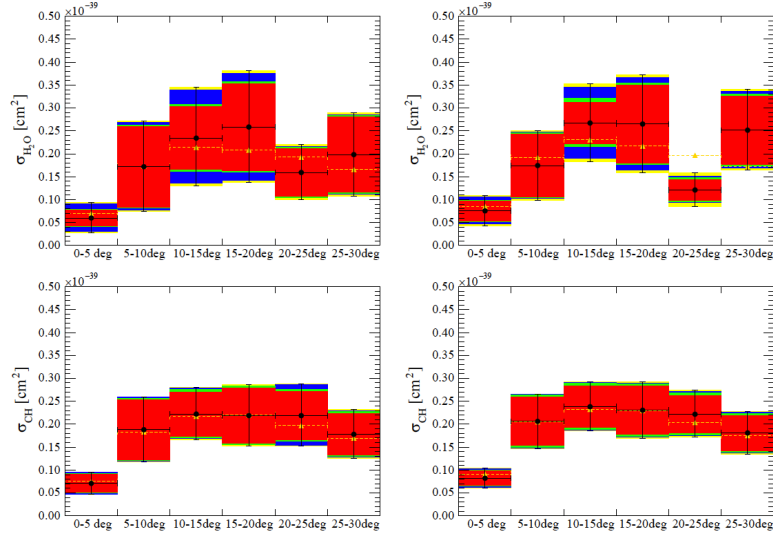
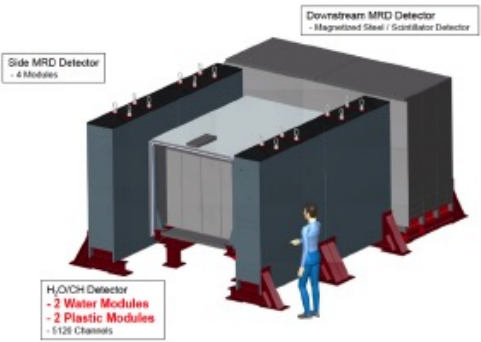
neutrino CC0 π 1p differential cross sections



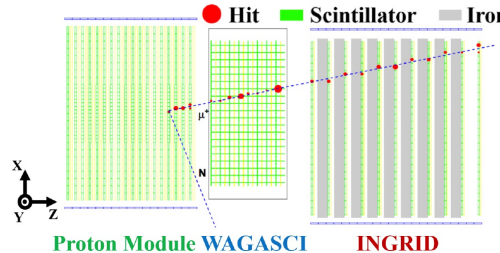
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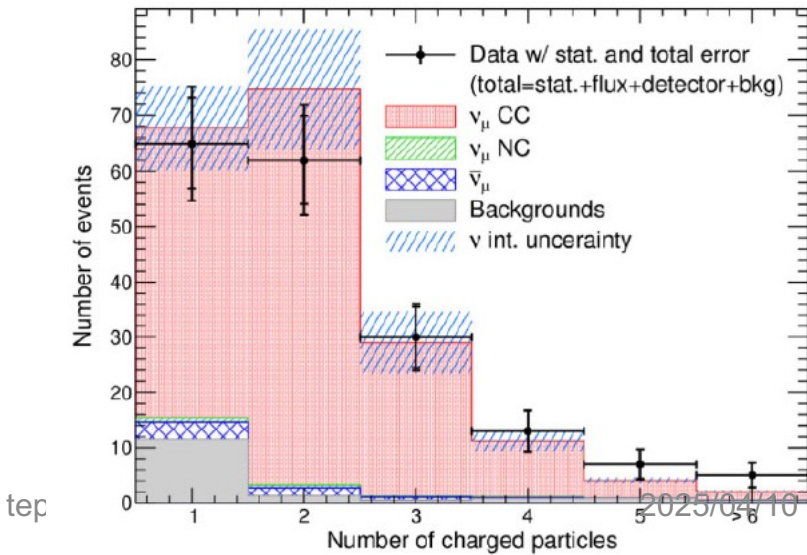
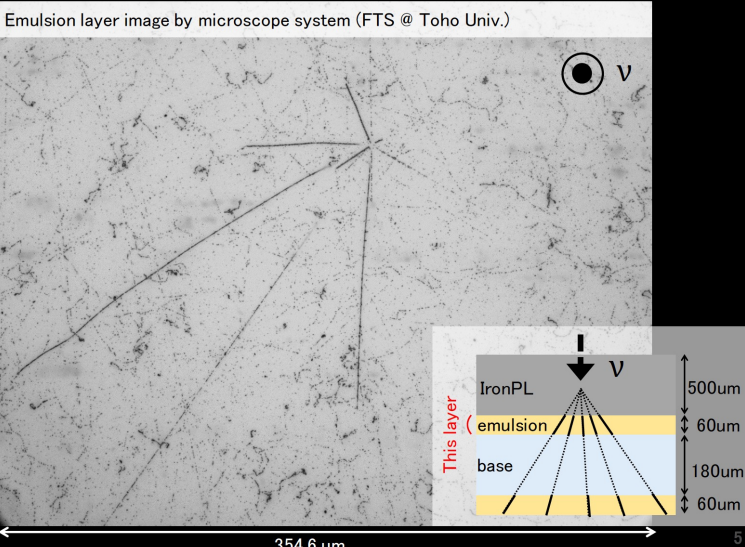
WAGASCI
- water target



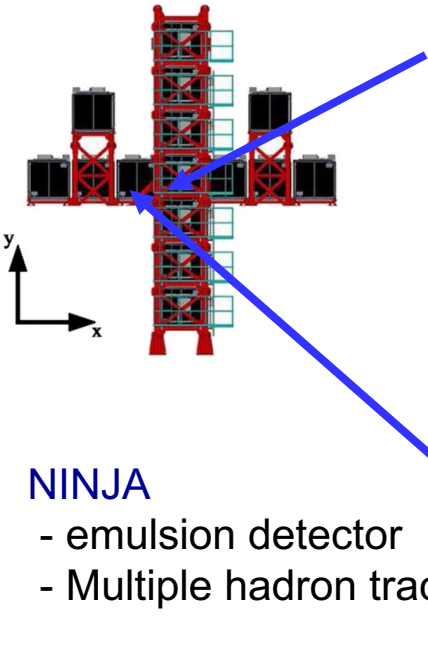
NINJA
- emulsion detector
- Multiple hadron tracks



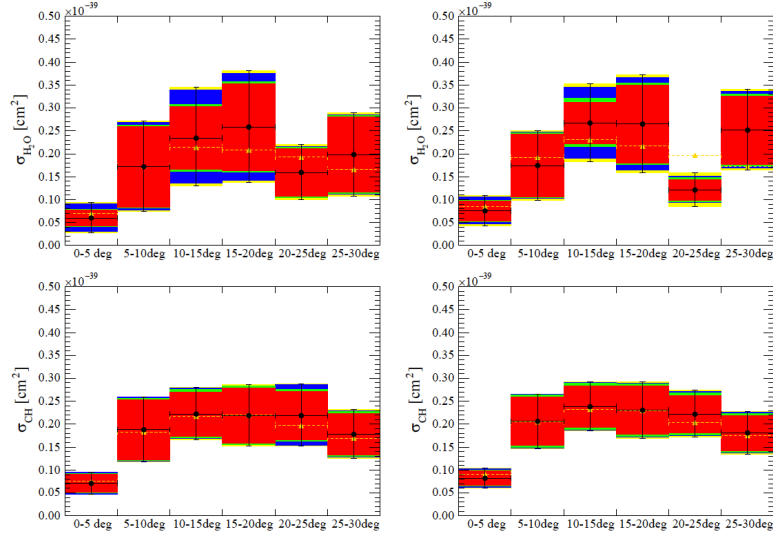
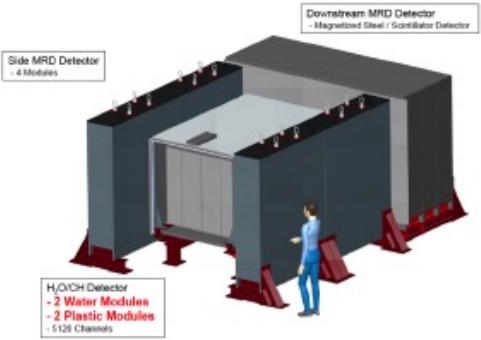
An example of ν – iron interaction (2016 NINJA iron target run)



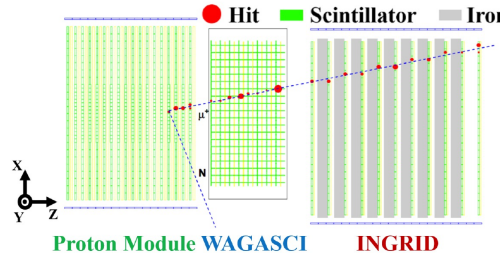
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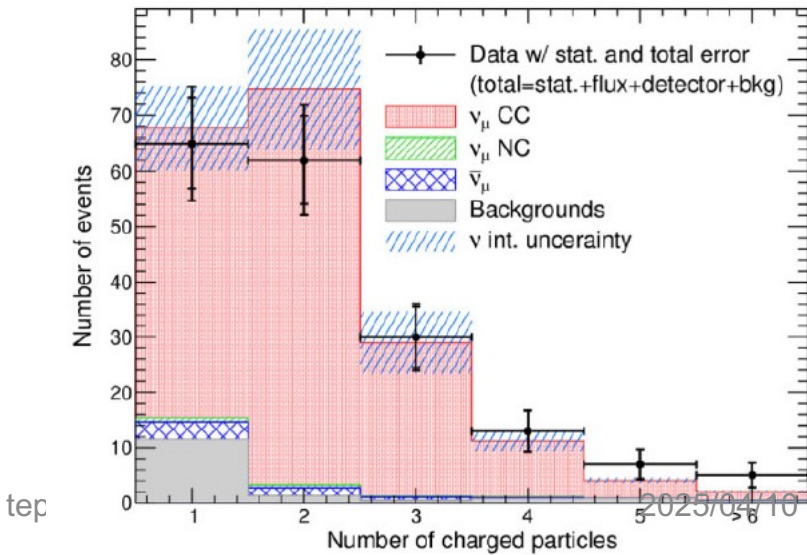
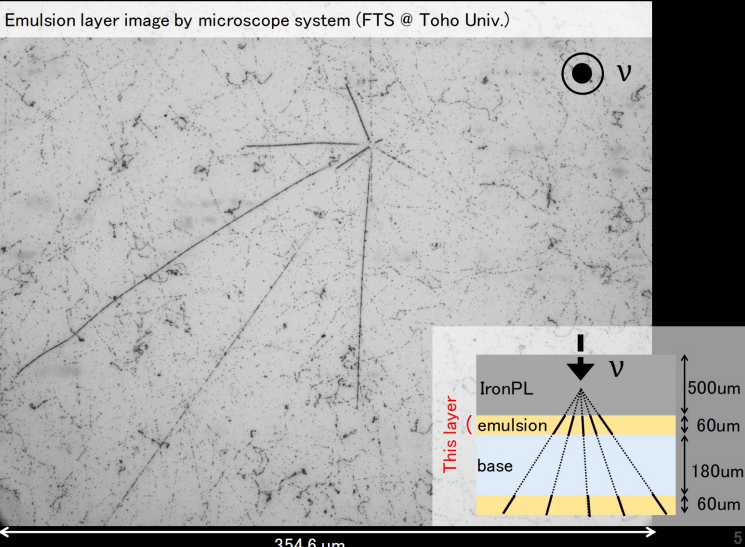
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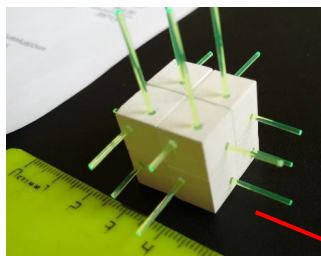
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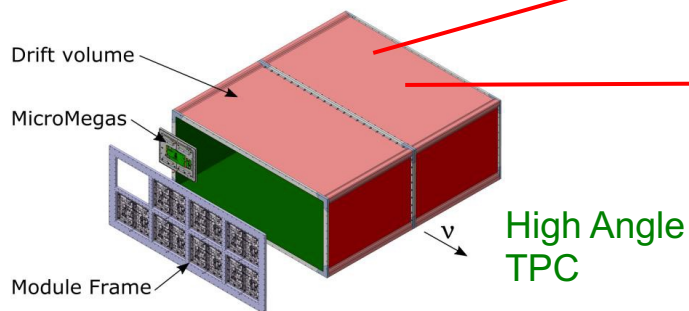
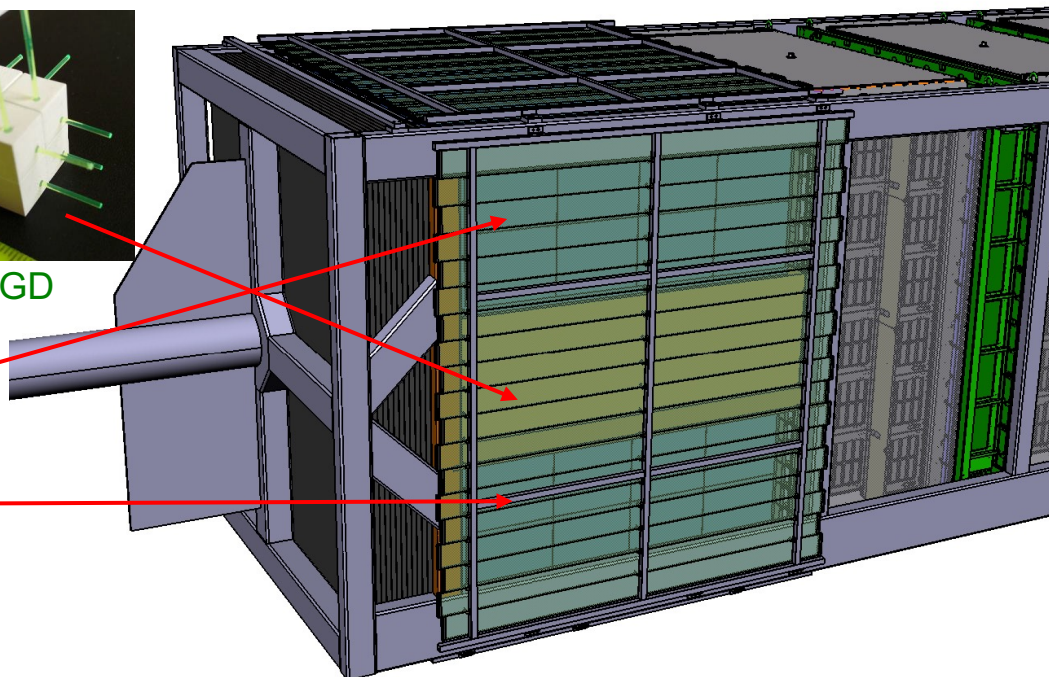
5. ND280 Upgrade

ND280 Upgrade

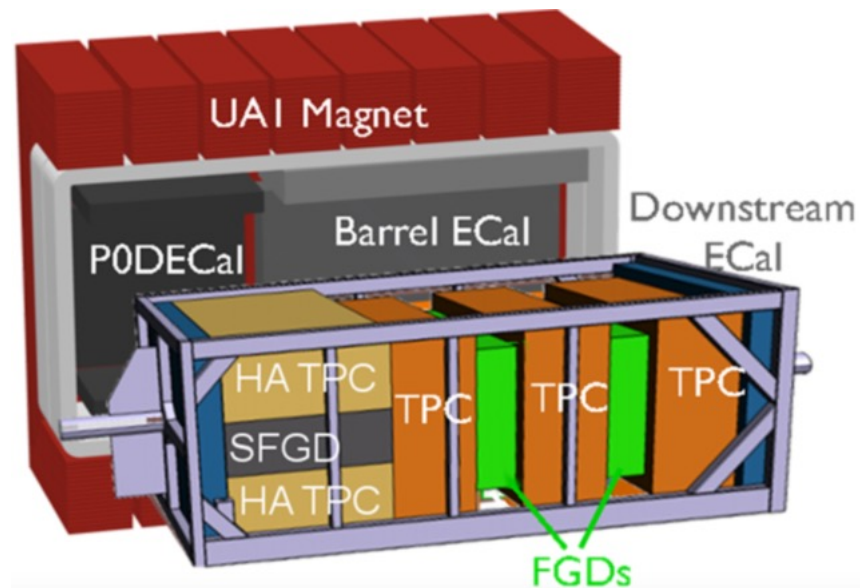
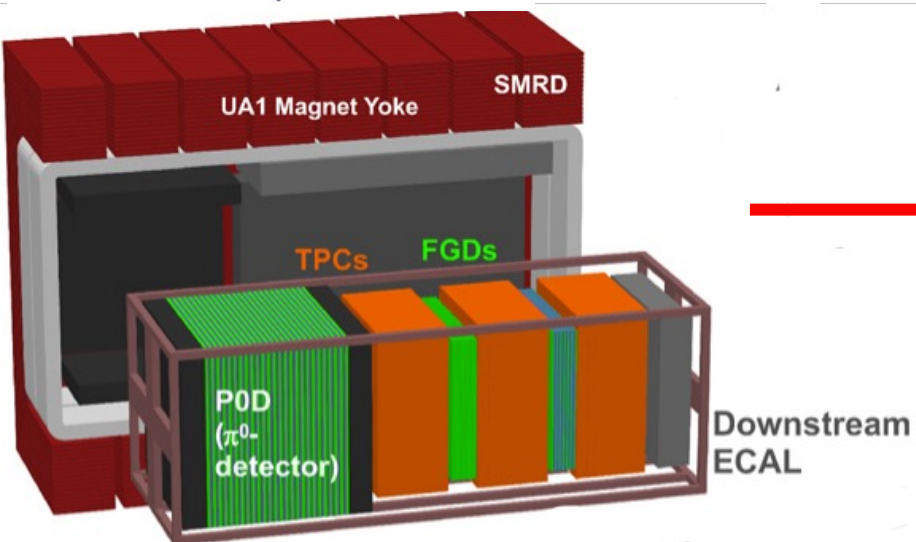
- Out: P0D detector
- In: High Angle TPC (HATPC)
- In: SuperFGD



SuperFGD



High Angle TPC



5. ND280 Upgrade

ND280 Upgrade

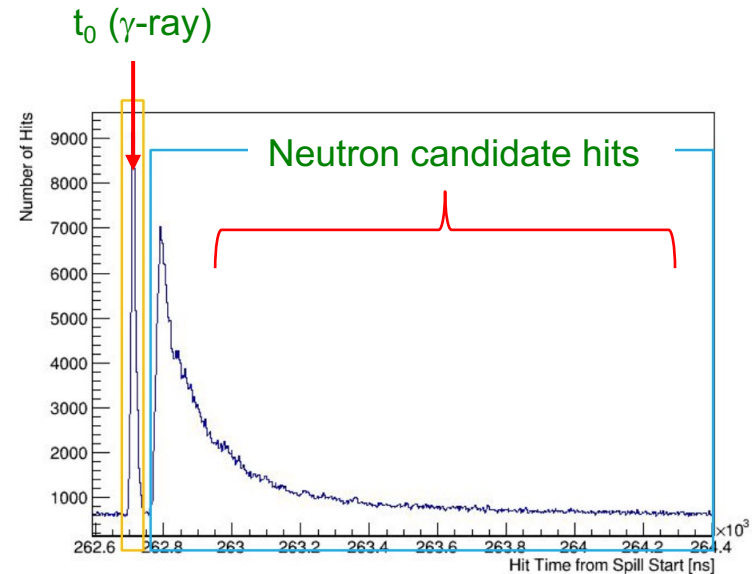
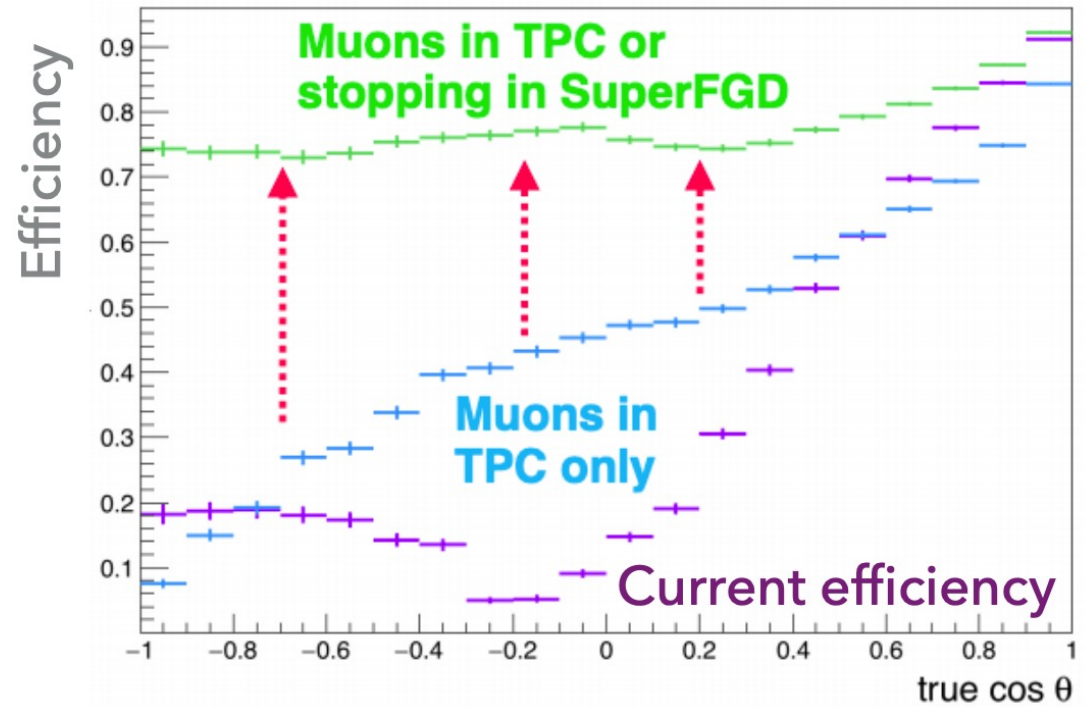
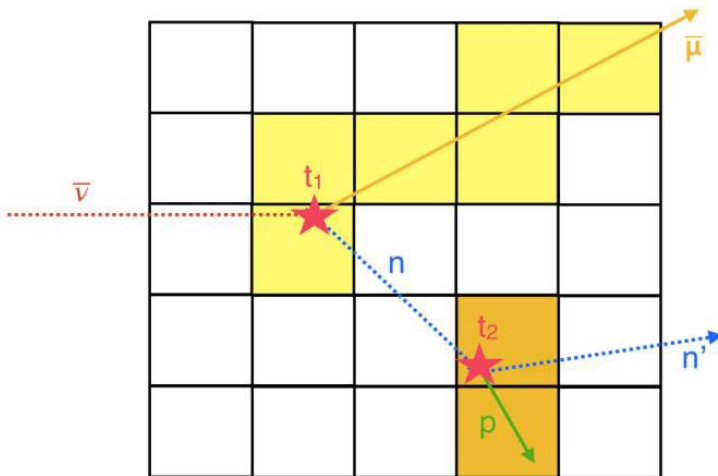
- Out: P0D detector
- In: High Angle TPC (HATPC)
- In: SuperFGD

4π coverage

- It matches with Hyper-K phase space

Neutron tagging

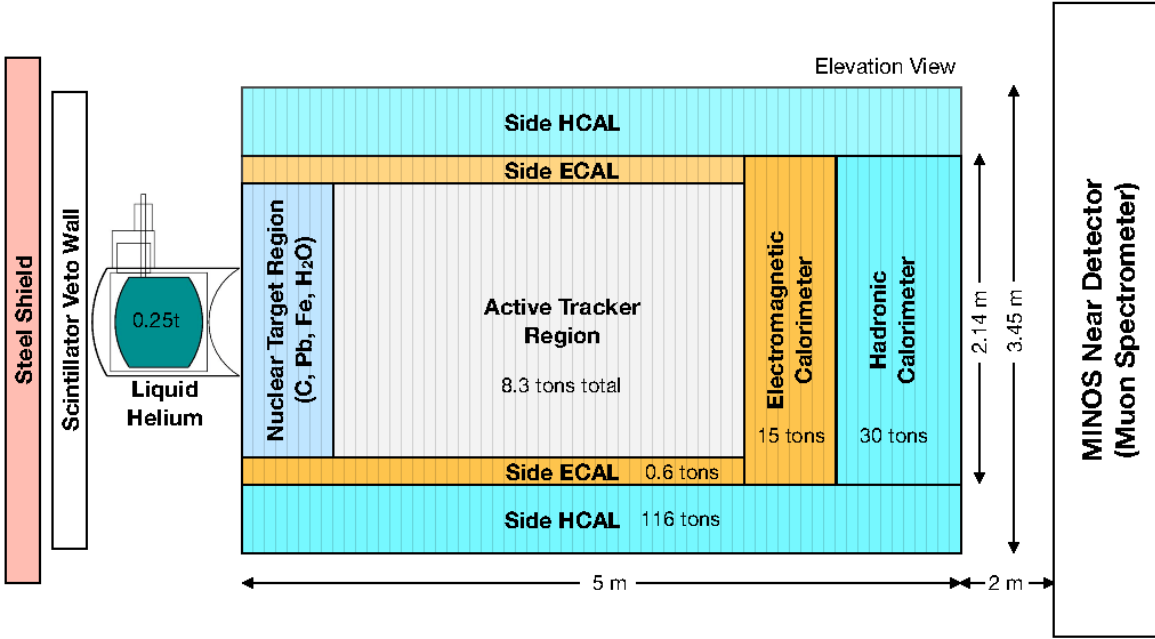
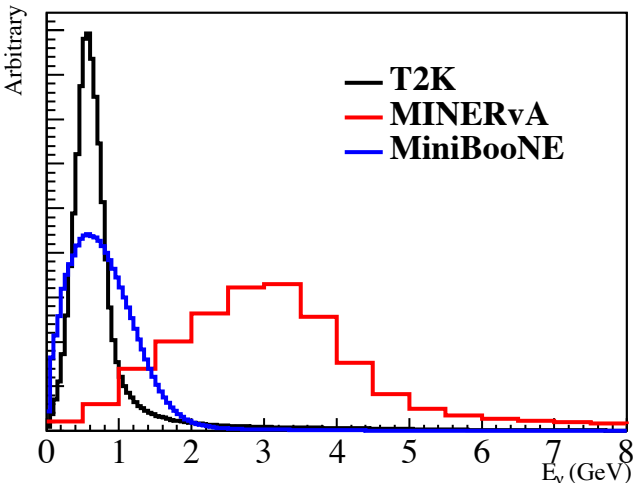
- SuperFGD beam test at LANL
- ToF to measure energy



5. MINERvA

Scintillation tracker

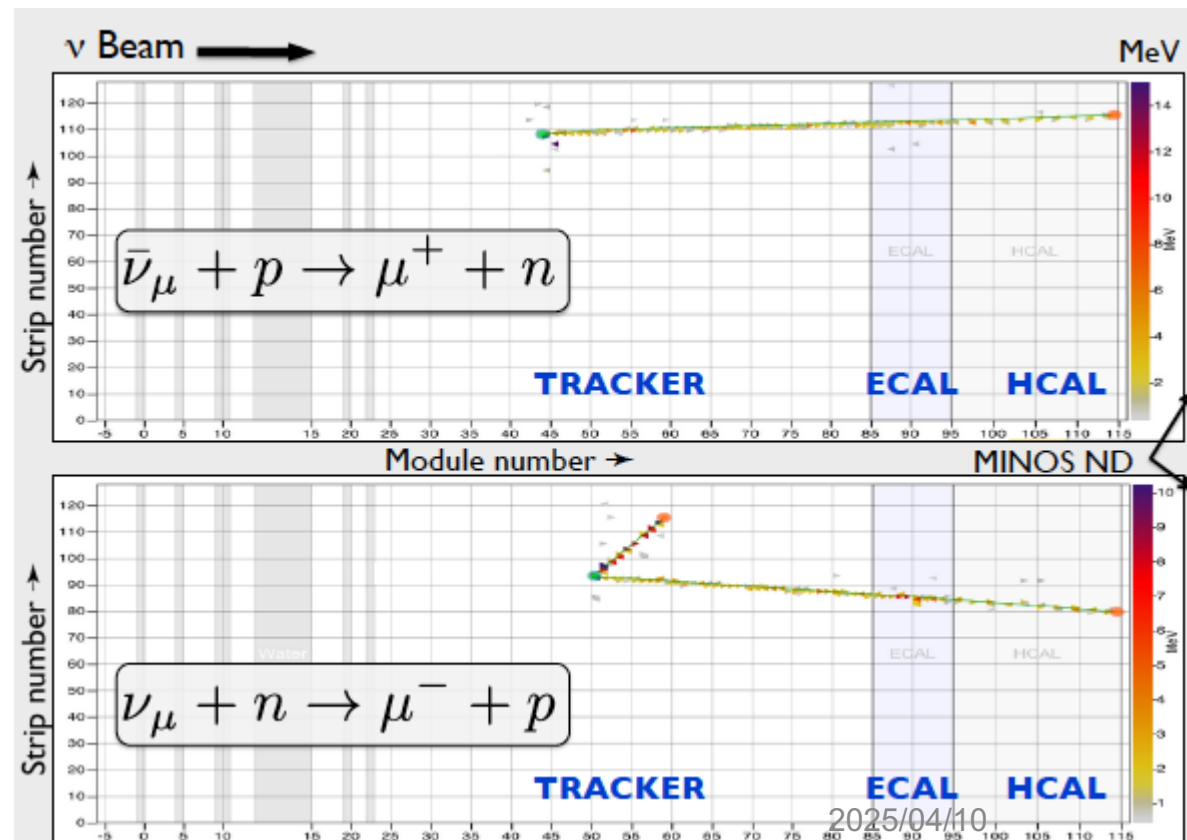
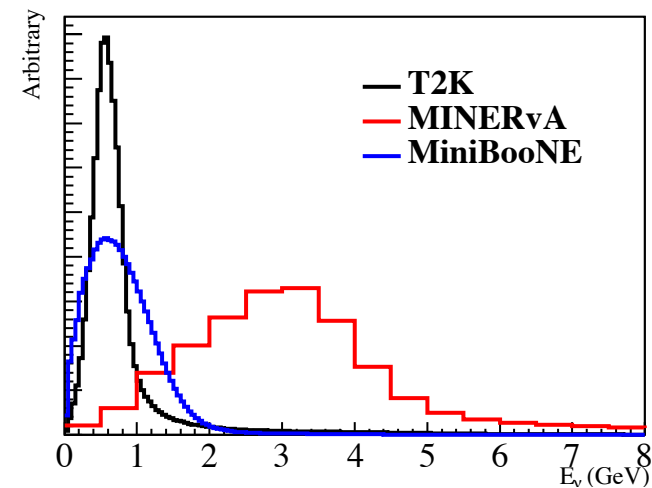
- $\langle E \rangle \sim 3.5$ GeV on-axis beam
- variety of targets (CH, Pb, Fe)
- Small acceptance due to MINOS ND
- charge separation by MINOS ND
- internal flux constraint (DIS, ν -e)



5. MINERvA

Scintillation tracker

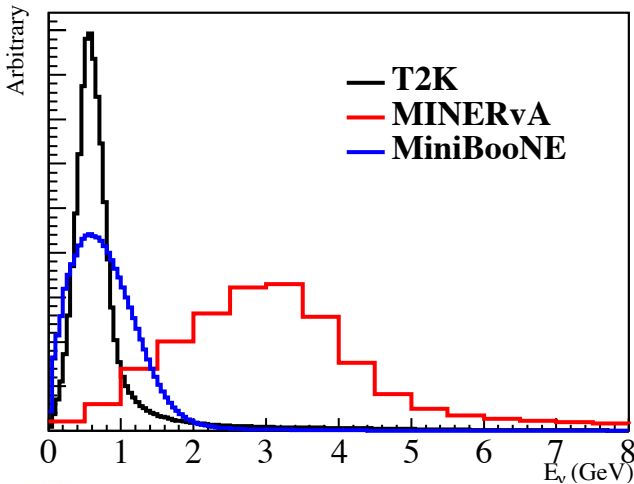
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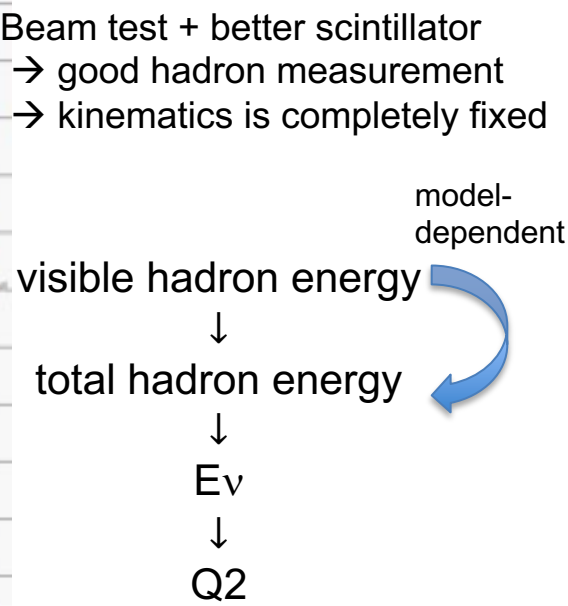
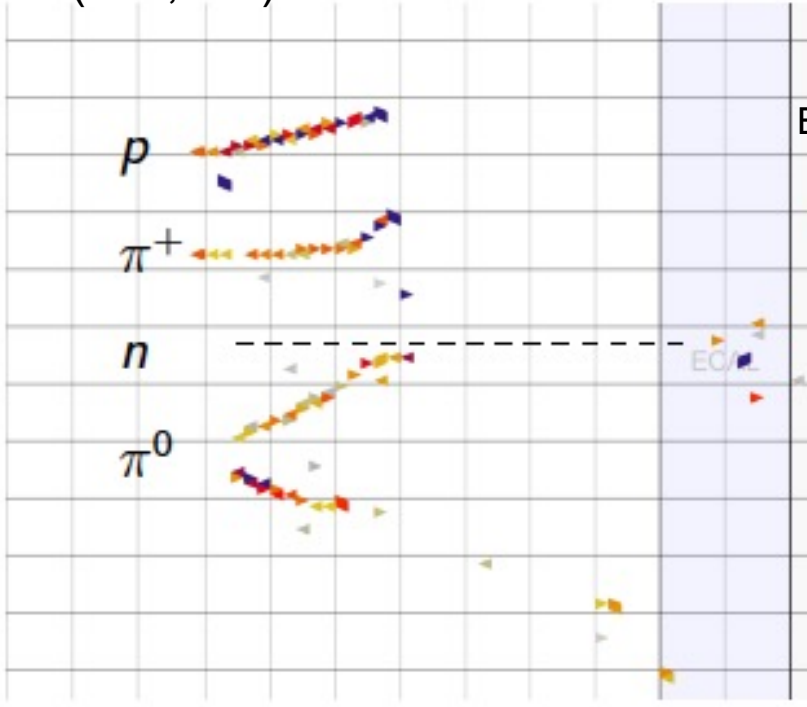


Kinetic energy

Kinetic energy

0

Total energy



On average, we see *available* hadronic energy $E_{avail} \neq q_0$:

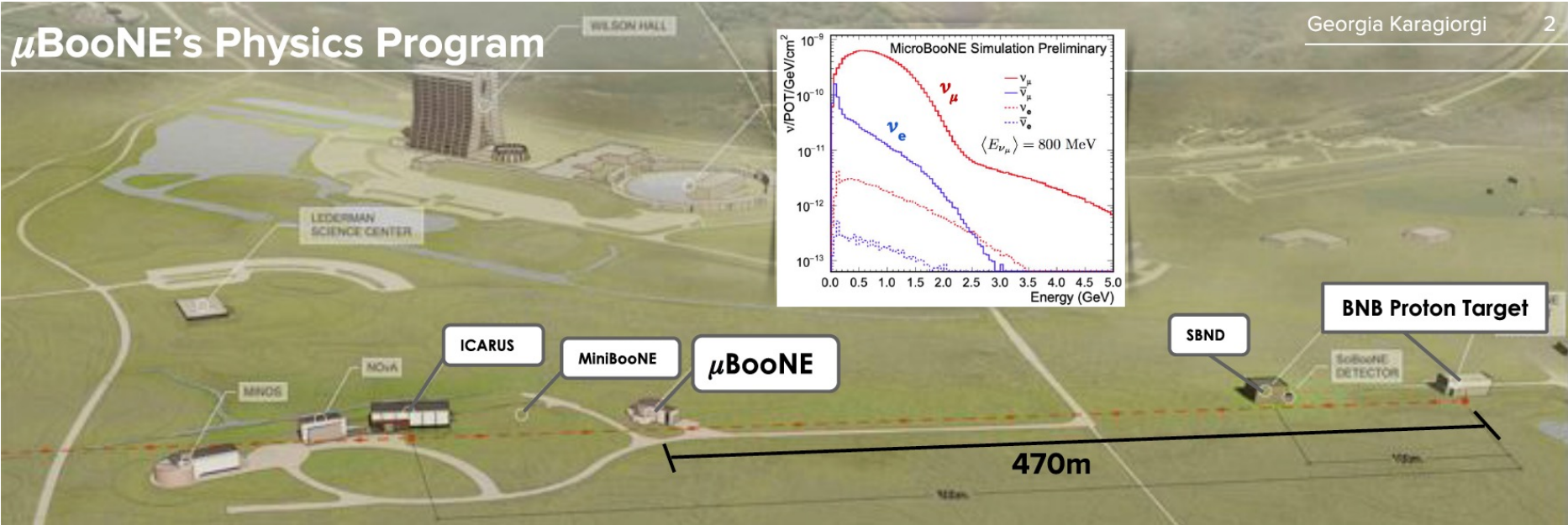
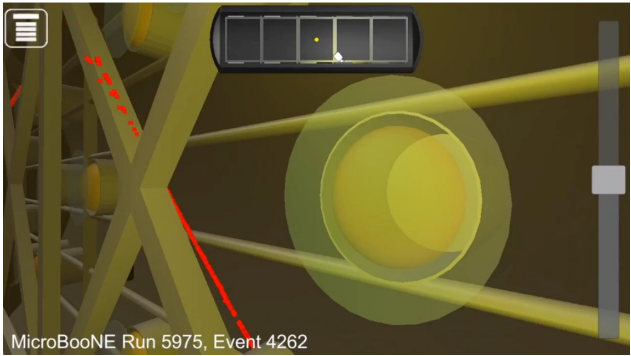
$$E_{avail} = \sum (\text{Proton and } \pi^\pm \text{ KE}) + (\text{Total } E \text{ of other particles except neutrons})$$

5. MicroBooNE

86ton LArTPC

- $\langle E \rangle \sim 800$ MeV BNB on-axis beam
- Single phase LArTPC, 3-wire-plane reading
- 3mm pitch
- photon detection system
- ArgoNeuT, LArLAT, SBND, ICARUS, protoDUNE, DUNE

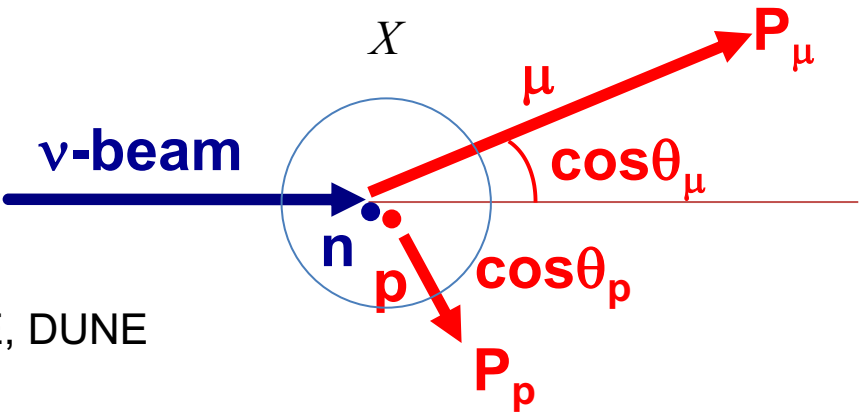
VENu (Virtual Environment of Neutrinos)
<http://venu.physics.ox.ac.uk/>
- smart phone app for MicroBooNE data



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MicroBooNE CC $\mu + p$ differential cross section

Outgoing proton kinematics are measured to reconstruct Fermi motion

Multiple Coulomb scattering to estimate escaping muon energy

Large cosmic ray background, but mostly understood

Low statistics for hadron measurements

