Modern Tests of Spacetime Symmetry

outline

- 1. Introduction
- 2. Test of Lorentz violation
- 3. Lorentz violation in astrophysics



24/12/05

Teppei Katori 🥑 @teppeikatori King's College London Ortvay colloquium, Eötvös Loránd University Budapest, Hungary, Dec 5, 2024

ZIMÁNYI SCHOOL 2024





tkatori@cern.ch

Introduction



(I want to say) Einstein is wrong!



I can demonstrate under scientific control and in a repeatable the speed of light. The brain has the energetic power of instantar the problem was to prove it scientifically, today it is possible, I ca









Theory of Special Relativity



Special relativity is a basis of both quantum field theory and general relativity

Special relativity is based on Lorentz symmetry

Lorentz symmetry is isotropy of the spacetime

If the universe has a special direction, space doesn't have Lorentz symmetry and Lorentz transformation is violated → Lorentz violation

All fundamental physics phenomena must be experimentally tested including Lorentz symmetry



Michelson-Morley experiment

The experiment tried to measure the motion of the Earth relative to æther.

The experiment shows the speed of light is constant regardless the motion of the Earth.

This result suggests the isotropy of the space, and Lorentz symmetry.

Lorentz symmetry is valid down to $\Delta c/c \sim 10^{-9}$







Nagel et al, Nature Comm., 6(2015)8174

Michelson-Morley experiment

The experiment has been improved over 100 years.





Nagel et al, Nature Comm., 6(2015)8174

Michelson-Morley experiment

The experiment has been improved over 100 years.

Technology shift (interferometer → optical cavity) around 2000s





Nagel et al, Nature Comm., 6(2015)8174,

Optical cavity experiment

Modern Michelson-Morley experiment

- Saphire crystal resonator
- Whispering gallery mode
- Vacuum insulation, liquid helium cooling to 4K
- Turntable to actively rotate

This experiment is sensitive to the anisotropy of speed of light down to $\Delta c/c \sim 10^{-18}$

Why we keep testing this?

Why do we expect Lorentz violation?







Progress in Particle and Nuclear Physics 125 (2022) 103948

Quantum gravity

Searching Lorentz violation is well motivated

Lorentz violation in Planck scale theories

- string theory
- noncommutative field theory
- quantum loop gravity
- extra dimensions

etc

Lorentz violation is seen as

- spacetime fluctuation
- background field in vacuum etc



Lorentz violating field - background field of the universe (æther)



tkatori@cern.ch

quantum foam

- quantum fluctuation of space-time



Quantum gravity

Searching Lorentz violation is well motivated

Quantum field theory and general relativity are the foundation of modern physics.

Lorentz symmetry is a basis for both quantum field theory and general relativity

How to formulate Lorentz violation in our theories?

Lorentz symmetry could be spontaneously broken, if so, this doesn't violate existing framework of modern physics





Spontaneous symmetry breaking

Magnetization

Nature has many examples of spontaneous symmetry breaking

- Condensed matter (magnetization, crystallization, etc)
- Phase transition in vacuum (Higgs mechanism, spontaneous Lorentz symmetry breaking)





Higgs mechanism





Spontaneous symmetry breaking

Searching Lorentz violation is well motivated

Math is a good approximation of nature There is no perfect symmetry in nature, all somewhat broken

So why space-time symmetry is so perfect?!





Golden ratio and seashell



Fibonacci number and broccoli

Standard-Model Extension (SME)

Search of Lorentz violation is to find anomalous effects due to the couplings of background fields and ordinary fields (electrons, muons, neutrinos, etc)

SME is an effective field theory framework to look for Lorentz violation

 $\mathcal{L} = i\bar{\psi}\gamma_{\mu}\partial^{\mu}\psi - m\bar{\psi}\psi + \bar{\psi}\gamma_{\mu}a^{\mu}\psi + \bar{\psi}\gamma_{\mu}c^{\mu\nu}\psi \cdots$

couplings with background fields

Physics of Lorentz violation

Standard Model

- Spectrum distortion,
- Sidereal time dependence, etc...

e.g.) vacuum Lagrangian for fermion



24h 00min 00sec: Solar day 23h 56min 4.1sec: Sidereal day

Alan Kostelecky, Indiana University

2025 recipient, Norman F. Ramsey Prize

For the development of the Standard Model Extension and for its application to, and inspiration for, a broad set of precision measurement tests across various physical systems, some of which have reached Planck-scale sensitivity.



nobn

noon

Bluhm, Kostelecky, Lane, Russell PRL 2002





Tests of Lorentz violation



EötWash, PRL85(2000)2869, PRL100,(2008)041101

Torsion pendulum

Eötvös experiment \rightarrow EötWash experiment

- Modern torsion balance
- Newton constant
- Equivalent of Principle

etc

Eric G. Adelberger, University of Washington, Seattle

2025 recipient, Einstein Prize

For outstanding contributions to experimental gravity using precision torsionbalance measurements, which have profound implications for fundamental physics.

EötWash





EötWash, PRL97(2006)021603 , PRL122(2019)231301

Torsion pendulum (electron)

Electron spin – background field coupling

- AlNiCo: all magnetic field is from electron spin
- SmCo₅: electron orbital motion creates magnetic field
- Magnetize them to cancel magnetic field, so that the pendulum has net electron spin
- Look for coupling between electron spin and background field







PRL105(2010)151604 and many others

Double gas maser (neutron)

The most sensitive magnetometer

- Optical pump for Rb, K
- Spin transfer to noble gas (Xe, ³He), monitor ³He precession
- Look for coupling between neutron spin and background field







ZEUS, PRD107(2023)092008 and many others

Collider physics (quarks)

HERA p-e⁻ collider

- ZEUS deep-inelastic scattering data
- Monitor sidereal time dependence
- Similar tests are possible for other data







Neutrino physics

Neutrinos in the standard model

The standard model describes 6 quarks and 6 leptons and 3 types of force carriers.



Neutrinos are special because,

1. they only interact with weak nuclear force.



2. interaction eigenstate is not Hamiltonian eigenstate, and propagation of neutrinos changes their species (flavours), called neutrino oscillation.



Neutrino oscillation is an interference experiment (cf. double slit experiment)



For double slit experiment, if path v_1 and path v_2 have different length, they have different phases and it causes interference.



Neutrino oscillation is an interference experiment (cf. double slit experiment)



Neutrino flavour eigenstates are super-position of Hamiltonian eigenstates v_1 and v_2



Neutrino oscillation is an interference experiment (cf. double slit experiment)



Neutrino flavour eigenstates are super-position of Hamiltonian eigenstates v_1 and v_2 Difference in velocities cause quantum interference



Neutrino oscillation is an interference experiment (cf. double slit experiment)



Neutrino flavour eigenstates are super-position of Hamiltonian eigenstates v_1 and v_2 Difference in velocities cause quantum interference The detection may be different flavour (neutrino oscillations)



Neutrino oscillation is an interference experiment (cf. double slit experiment)



Neutrino flavour eigenstates are super-position of Hamiltonian eigenstates v_1 and v_2 Difference in velocities cause quantum interference The detection may be different flavour (neutrino oscillations)

Neutrino propagation may be affected by background fields

 \rightarrow anomalous neutrino oscillation results



LSND, PRD64(2001)112007, MiniBooNE, PRL121(2018)22180

Neutrino oscillation experiments

Neutrino physics \rightarrow Home of anomalies

- Solar and atmospheric neutrino anomalies (Nobel prizes, 2002, 2015)



SNO detector



The Nobel Prize in Physics 2015





noto © Takaaki Kajita Takaaki Kajita Prize share: 1/2

Queen's University /SNOLAB Arthur B. McDonald Prize share: 1/2

The Nobel Prize in Physics 2002





Raymond Davis Jr. Prize share: 1/4

The Nobel Prize in Physics

Masatoshi Koshih Prize share: 1/4









LSND, PRD64(2001)112007, MiniBooNE, PRL121(2018)22180

Neutrino oscillation experiments

Neutrino physics \rightarrow Home of anomalies

- Solar and atmospheric neutrino anomalies (Nobel prizes, 2001, 2015)
- OPERA Neutrino-faster-than-Speed-of-Light (detector problem)
- LSND excess
- MiniBooNE excess





If these anomalous neutrino oscillation data are due to Lorentz MiniBooNE excess violation, data may show sidereal time dependence

LSND, PRD72(2005)076004, MiniBooNE, PLB718(2013)1303, TK, MPLA27(2012)1230024

Neutrino oscillation experiments

Neutrino physics \rightarrow Home of anomalies

- Solar and atmospheric neutrino anomalies (Nobel prizes, 2001, 2015)
- OPERA Neutrino-faster-than-Speed-of-Light (detector problem)
- LSND excess
- MiniBooNE excess



Tests of Lorentz violation – Summary

Limits of SME parameters are summarized in tables <u>https://arxiv.org/abs/0801.0287v17</u>

So far, there is no compelling evidence of Lorentz violation

Table D15. Photon sector, d = 3

Combination		\mathbf{Result}	\mathbf{System}		Ref.	
$ k_{(V)00}^{(3)} $		$(7.32 \pm 2.94) \times 10^{-45} \text{ GeV}$	CMB polarization		$[131], [132]^*$	
$ k_{(V)00}^{(3)} $		$< 1.54 \times 10^{-44}~{\rm GeV}$	"		[133]*	
$ \mathbf{k_{AF}} $		$<7.4\times10^{-45}~{\rm GeV}$	"		[133]*	
$ \mathbf{k_{AF}} $		$< 1.03 \times 10^{-26}~{\rm GeV}$	Satellites		[134]*	
k^Z_{AF} ,,	Table D10. Proton sector, $d = 4$ (part 2 of 2)					
$ k_{(V)10}^{(3)} $	Combination	Result	\mathbf{System}	Ref.]*	
$ k_{(V)11}^{(3)} $	$ \overline{c}^p_{0k} $	$< 1 \times 10^{-8}$	Binary pulsars	[75]*]*	
$k_{(V)10}^{(3)}$	$ \overline{c}_{jk}^p $	$< 1 \times 10^{-11}$	"	[75]*], [138]*	
(1)10	$ ilde{c}_Q $	$< 2 \times 10^{-11} \ {\rm GeV}$	Relativistic Li ions	[72]		
	\overline{c}_{TT}	$(0.24\pm0.30) imes10^{-6}$	Nuclear binding energy	[76]		
VING'S	"	$(-3.3\pm3.5) imes10^{-6}$	Cs interferometer	[77]		
College	$ ilde{c}_Q$	$(-0.3 \pm 2.2) \times 10^{-22} \text{ GeV}$	Cs fountain	[105]	r	
LONDON	$ ilde{c}$	$(-1.8 \pm 2.8) \times 10^{-25} \text{ GeV}$	"	[105]	X	

Table D32. Neutrino sector, d = 4 (part 1 of 13)

Combination	\mathbf{Result}	System	Ref.
$(c_{ m of}^{(4)})_{00}$	$> -4 \times 10^{-19}$	IceCube	[275]*
$ (c_{ m of}^{(4)})_{00} $	$<7.1\times10^{-9}$	SN1987A time of flight	[18]*
"	$< 1.4 \times 10^{-4}$	Fermilab time of flight	[18]*
$\left(c_{ m of}^{(4)} ight)_{00}$	$-8.4\pm1.1^{+1.2}_{-0.9}\times10^{-5}$	OPERA time of flight	[18]*
"	$-1.8 \pm 1.0 \times 10^{-4}$	MINOS time of flight	[18]*
$\left(c_{ ext{of}}^{(4)} ight)_{10}$	$(-1 \text{ to } 4) \times 10^{-17}$	IceCube	[275]*
$ ig(c_{ ext{of}}^{(4)}ig)_{10} $	$<4.4\times10^{-9}$	SN1987A time of flight	[18]*

ible D12. Neutron sector, d = 3, 4 (part 2 of 2)

Bosult	System	Bof
Result	System	nei.
$<(3\pm27\pm27)\times10^{-14}$	Macroscopic matter	$[123]^*$
$< 1 \times 10^{-8}$	Binary pulsars	$[75]^*$
$<1\times10^{-11}$	"	[75]*
$(-4\pm6) imes10^{-6}$	Gravimetry	$[124]^*$
$(-1\pm1) imes10^{-5}$	"	$[124]^*$
$(-1\pm1)\times10^{-5}$	"	$[124]^*$
$(-1.8 \pm 2.2) \times 10^{-14} \text{ GeV}$	Quartz oscillators	[125]
$(1.1 \pm 1.4) imes 10^{-6}$	Nuclear binding energy	[76]
$(7.6\pm 6.7) imes 10^{-6}$	Cs interferometer	[77]
$(4.8 \pm 4.4) imes 10^{-29}$	Ne/Rb/K magnetometer	[107]
$(-2.8 \pm 3.4) \times 10^{-29}$	"	[107]

When do we find Lorentz violation???

Lorentz violation is motivated by Planck scale theories, so it is suppressed with the power of Planck mass ($\sim 10^{19} \text{ GeV}$)

$$\sim \frac{M}{M_{Pl}}, \left(\frac{M}{M_{Pl}}\right)^2$$
, etc

In effective field theory, non-renormalizable operators are the signature of new physics, dimension analysis guides target sensitivity to look for Lorentz violation.

dimension-5 LV operator $< 10^{-19} GeV^{-1}$ dimension-6 LV operator $< 10^{-38} GeV^{-2}$ etc

These numbers can be used as a guidance to design new experiments



Steven Weinberg (CERN Courier Nov. 2017)

"We don't know anything about non-renormalizable interaction terms, but I'll swear they are there!"



Tests of Lorentz violation – Astrophysics

Terrestrial experiments

- controlled, high-precision
- various systems (optics, pendulum, gas, particle physics, etc)

So far, no compelling evidence of Lorentz violation

Astrophysical and cosmological experiments

- not controlled, low-precision
- extreme systems (highest energy, longest distance, etc)
- more sensitive to nonrenormalizable operators



Tests of Lorentz violation in Astrophysics



Progress in Particle and Nuclear Physics 125 (2022) 103948

Lorentz violation in Astrophysics

ELSEVIER

Progress in Particle and Nuclear Physics

Contents lists available at ScienceDire



Review

Quantum gravity phenomenology at the dawn of the multi-messenger era—A review



Highest energy particles – Ultra-high-energy cosmic rays Longest propagating particles – Gravitational waves, cosmic microwave background High-energy and long propagation – Gamma-ray, High-energy astrophysical neutrinos



Amelino-Camelia et al, Nature 393(1998)763, Stecker, et al, PRD91(2015)045009, and many others

Cut-off in high-energy cosmic ray spectrum

Lorentz violation = media in vacuum - Attenuate high-energy cosmic rays?





High-energy astrophysical neutrino spectrum

Amelino-Camelia et al, Nature 393(1998)763, Auger, JCAP01(2022)023, and many others

Cut-off in high-energy cosmic ray spectrum

Lorentz violation = media in vacuum

- Attenuate high-energy cosmic rays?





24/12/05

Huang and Ma, Communications Physics1(2018)62, Amelino-Camelia et al, Nature Astronomy 7(2023)996 and many others

Time delay of high-energy cosmic rays

Lorentz violation = media in vacuum

- Time difference between photons and neutrinos?

Gamma Ray Bursts

- Not identified as point neutrino sources
- Time delay or advance proportion to neutrino energy and sign may fit with data



Could quantum gravity slow down neutrinos?

<u>Giovanni Amelino-Camelia</u> [™], <u>Maria Grazia Di Luca, Giulia Gubitosi</u>, <u>Giacomo Rosati</u> & <u>Giacomo D'Amico</u>

Nature Astronomy 7, 996–1001 (2023) Cite this article





Anomalous neutrino mixings in vacuum

Lorentz violation = media in vacuum

- Neutrino oscillations are affected by media
- If the universe is saturated with background field, they would affect flavours of astrophysical neutrinos

Sensitive to the target signal region of Lorentz violation $(< 10^{-38} GeV^{-2}$ for dimension-6 operators), no anomalous neutrino oscillation is discovered yet





Neutrino astronomy – Summary

Neutrino astronomy has a high potential to look for Lorentz violation. But there are many unknowns;

- Energy spectrum
- Sources (5 known sources, Sun, SN1987, TXS0506+056, NGC1068, Galactic plane)
- Flavour structure

So far, astrophysical neutrino data are low statistics and further data are needed to search Lorentz violation



Hyper-Kamiokande, ArXiv:1805.04163

Hyper-Kamiokande and IceCube-Gen2

New international neutrino astronomy projects around the world







tkatori@cern.ch



24/12/05

Super-Kamiokande



39

IceCube-Gen2, J.Phys.G48(2021)060501

Hyper-Kamiokande and IceCube-Gen2

New international neutrino astronomy projects around the world





IceCube (~1Gton)

Gen2-Optical (~8 Gton)

DeepCore (>7 GeV) IceCube-Upgrade (>3GeV)



The first stage of Gen2 (IceCube upgrade) is ongoing

tkatori@cern.ch

Bluhm, Kostelecky, Lane, Russell PRL 2002

Lorentz violation is motivated from Planck-scale theories

There is a worldwide effort to look for Lorentz violation, using various state-of-the-art techniques, but so far no compelling evidence of Lorentz violation

Neutrino oscillation and neutrino astronomy are powerful tools to look for Lorentz violation

Merry Christmas!



tkatori@cern.ch

Backup



Gravity test

Matter wave interferometer







g



High-energy astrophysical neutrinos

Above ~100 TeV, neutrinos are only particles pointing to their high-energy sources





High-energy astrophysical neutrinos

60TeV- 2PeV astrophysical neutrinos are observed by IceCube Neutrino Observatory



IceCube-Gen2, J.Phys.G48(2021)060501

IceCube event morphology

Track v_{μ} CC $v_{\mu} + N \rightarrow \mu + X$ Cascade v_e CC, v_τ CC, NC $v_e + N \rightarrow e + X$ $v_\tau + N \rightarrow \tau + X$ $v_\chi + N \rightarrow v_\chi + X$

Double cascade v_{τ} CC (L~50m•E/PeV) $v_{\tau} + N \rightarrow \tau + X$ $\tau \rightarrow X'$





Spontaneous symmetry breaking (SSB)

vacuum Lagrangian for fermion
$$L = i \overline{\psi} \gamma_{\mu} \partial^{\mu} \psi$$

In the Standard Model, a phase transition of a scalar field gives nonzero field value in vacuum







$$L = \frac{1}{2} \left(\partial_\mu \phi\right)^2 - \frac{1}{2} \mu^2 \phi^2 - \frac{1}{4} \lambda \phi^4$$

Spontaneous symmetry breaking (SSB)

vacuum Lagrangian for fermion
$$L = i \overline{\psi} \gamma_{\mu} \partial^{\mu} \psi - m \overline{\psi} \psi$$

In the Standard Model, a phase transition of a scalar field gives nonzero field value in vacuum





Particle acquires mass term!

 $L = \frac{1}{2} \left(\partial_{\mu} \phi \right)^{2} - \frac{1}{2} \mu^{2} \phi^{2} - \frac{1}{4} \lambda \phi^{4}$

Spontaneous Lorentz symmetry breaking (SLSB)

vacuum Lagrangian for fermion $L = i \overline{\psi} \gamma_{\mu} \partial^{\mu} \psi - m \psi \psi$

In the Standard Model, a phase transition of a scalar field gives nonzero field value in vacuum

In String Theory, a vector field can be frozen in vacuum by spontaneous symmetry broken





Spontaneous Lorentz symmetry breaking (SLSB)

vacuum Lagrangian for fermion $L = i\bar{\psi}\gamma_{\mu}\partial^{\mu}\psi - m\bar{\psi}\psi + \psi\gamma_{\mu}a^{\mu}\psi$

In the Standard Model, a phase transition of a scalar field gives nonzero field value in vacuum

In String Theory, a vector field can be frozen in vacuum by spontaneous symmetry broken



Lorentz symmetry is spontaneously broken!





 $\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})$







 $\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})$





Under the particle Lorentz transformation:

 $U \overline{\Psi}(x) \gamma_{\mu} a^{\mu} \Psi(x) U^{-1}$







Under the particle Lorentz transformation:

$$\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \rightarrow \mathbf{U}[\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})]\mathbf{U}^{-1}$$

 \neq Ψ(ΛX)γ_μa[∞]Ψ(ΛX)

Lorentz violation is observable when a particle is moving in the fixed coordinate space





Under the particle Lorentz transformation:

$$\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \rightarrow \mathsf{U}[\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})]\mathsf{U}^{-1}$$
$$\neq \overline{\Psi}(\Lambda \mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\Lambda \mathbf{x})$$

Lorentz violation is observable when a particle is moving in the fixed coordinate space







Under the observer Lorentz transformation:

 $\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})$

Under the particle Lorentz transformation:

$$\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \rightarrow \mathsf{U}[\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})]\mathsf{U}^{-1}$$
$$\neq \overline{\Psi}(\Lambda \mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\Lambda \mathbf{x})$$

Lorentz violation is observable when a particle is moving in the fixed coordinate space





Under the observer Lorentz transformation:

 $\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})$ $\mathbf{x} \rightarrow \Lambda^{-1}\mathbf{x}$





tkatori@cern.ch

Under the particle Lorentz transformation:

$$\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \rightarrow \mathsf{U}[\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})]\mathsf{U}^{-1}$$
$$\neq \overline{\Psi}(\Lambda \mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\Lambda \mathbf{x})$$

Lorentz violation is observable when a particle is moving in the fixed coordinate space

Under the observer Lorentz transformation:

$$\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \xrightarrow{\Lambda^{-1}} \overline{\Psi}(\Lambda^{-1}\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\Lambda^{-1}\mathbf{x})$$

Lorentz violation cannot be generated by observers motion (coordinate transformation is unbroken)

all observers agree for all observations



